

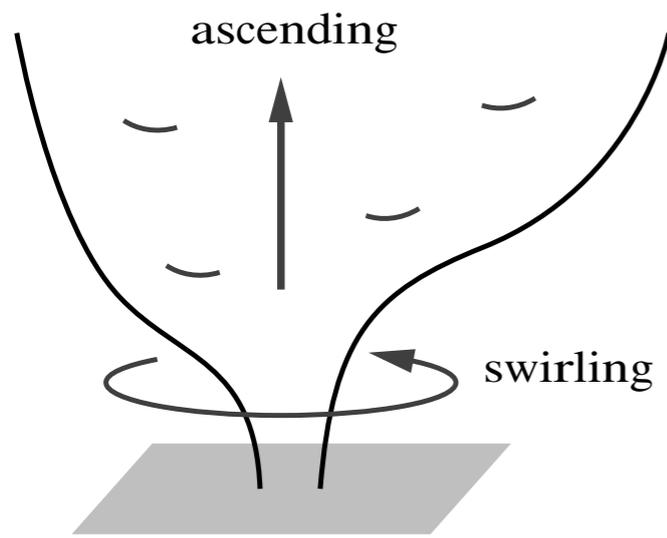
GFD Seminar  
Shikotsu-Lake, 15 March 2025

# **Modelling turbulence based on response-function and multiple-scale formulation**

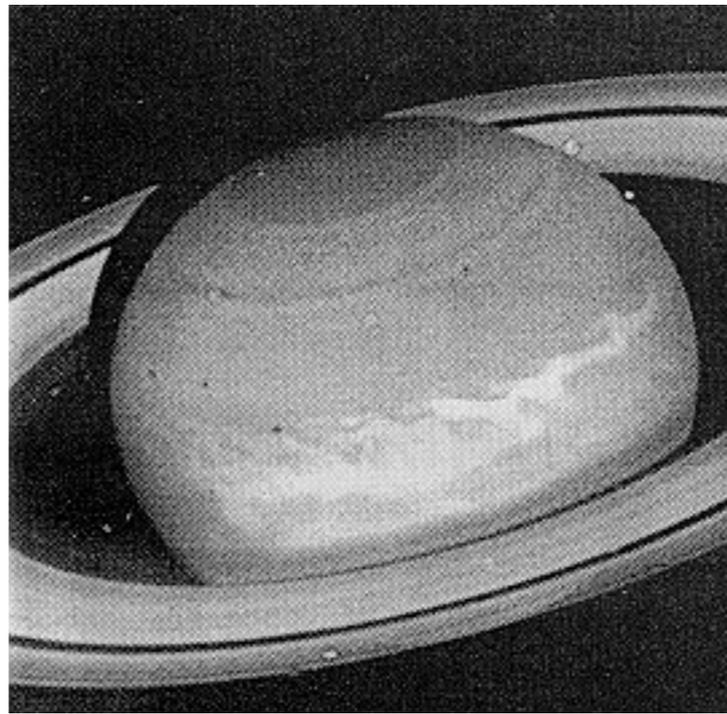
***Nobumitsu YOKOI***

*Institute of Industrial Science (IIS), Univ. of Tokyo*

# **Large-scale structures in Turbulence**



Typhoon, tornado

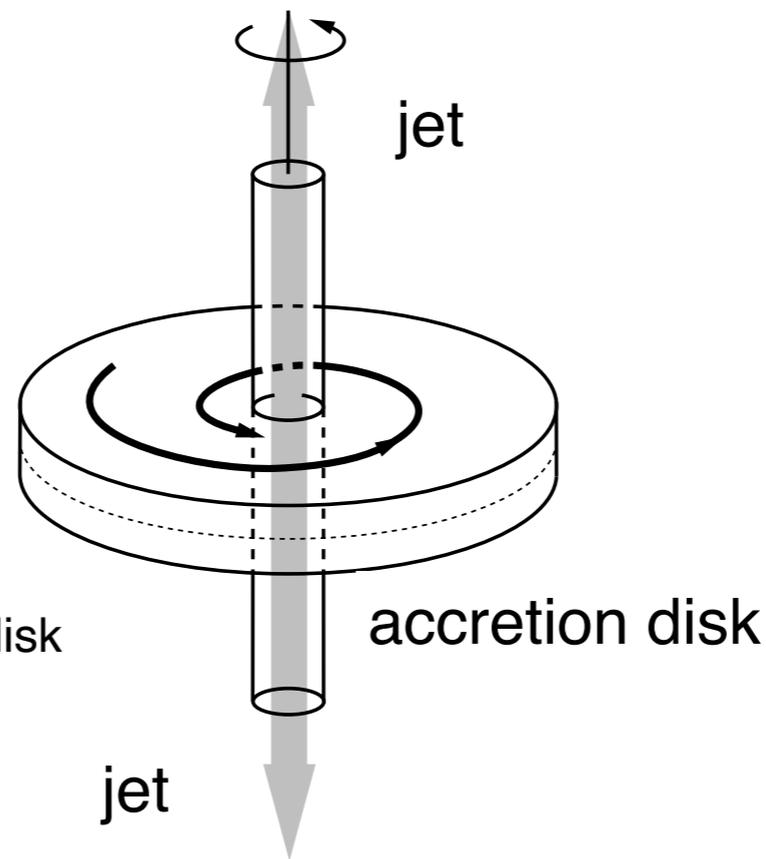
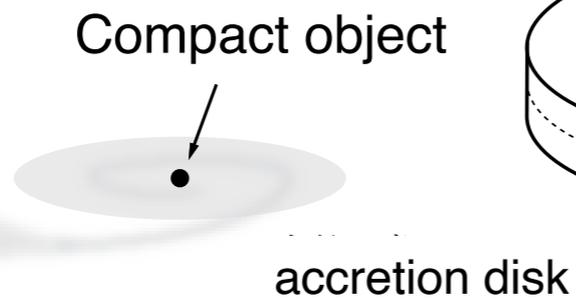


White spot on Saturn

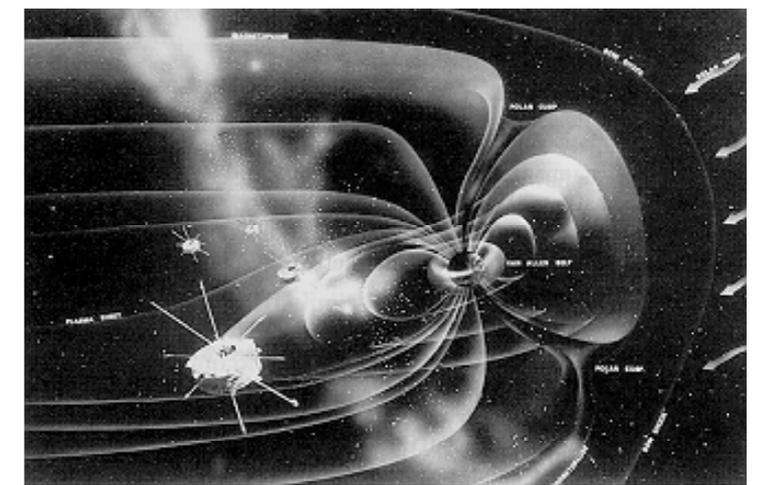


Magnetic field configuration in galaxy (M51)

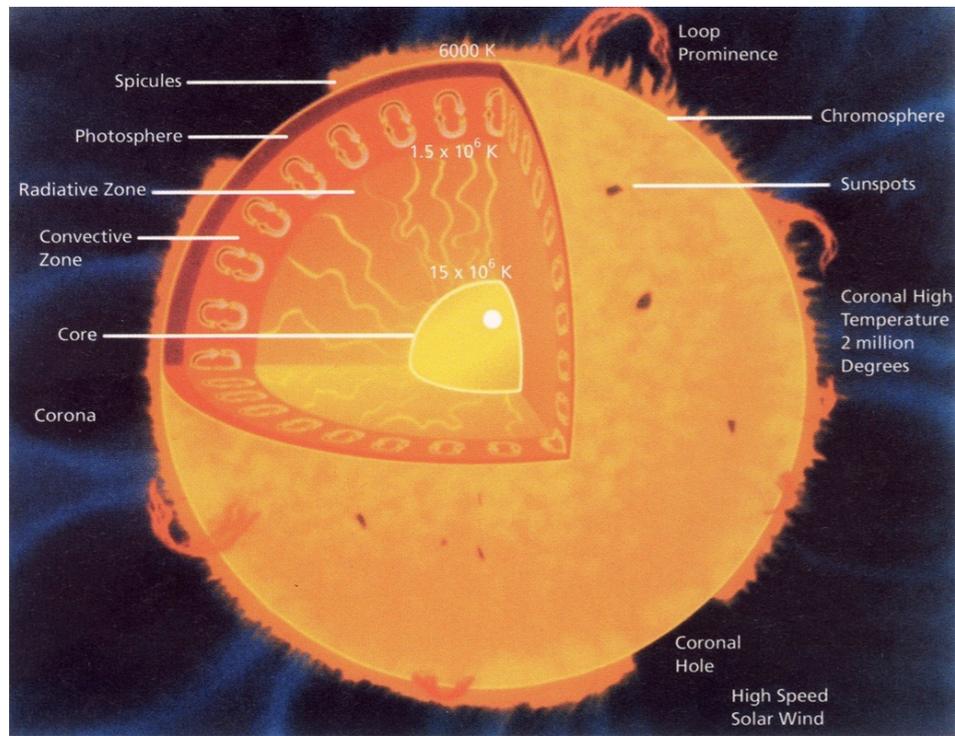
Accompany star



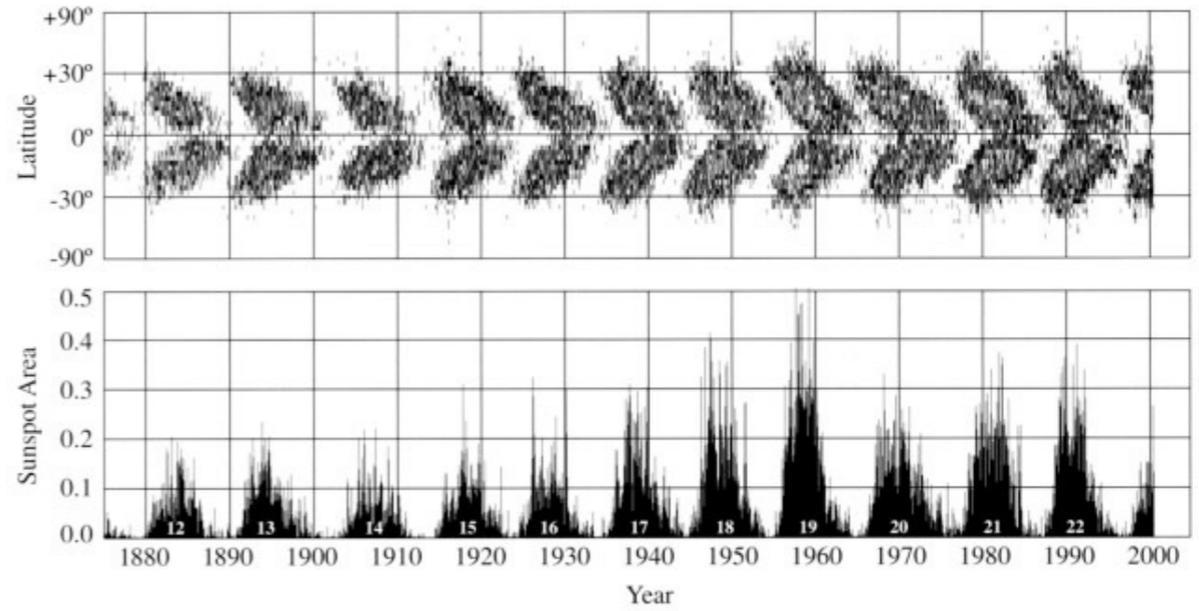
Accretion disk and its jet



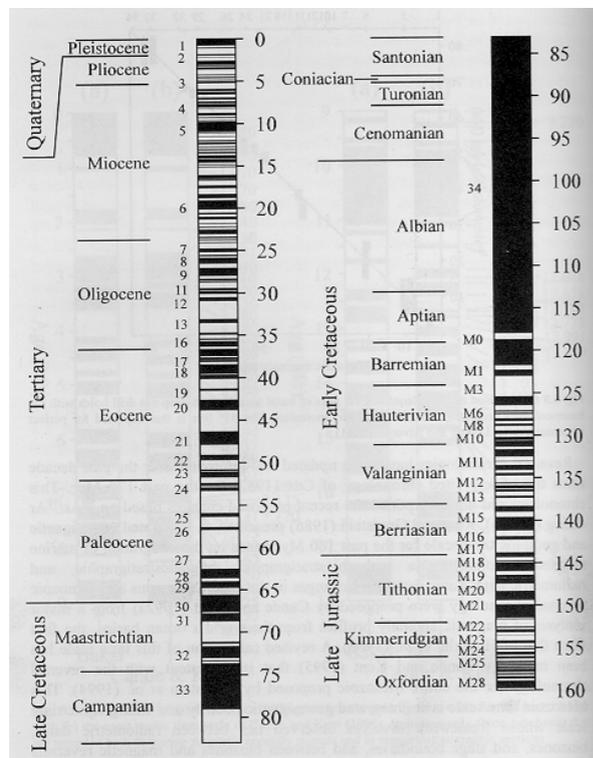
Solar winds



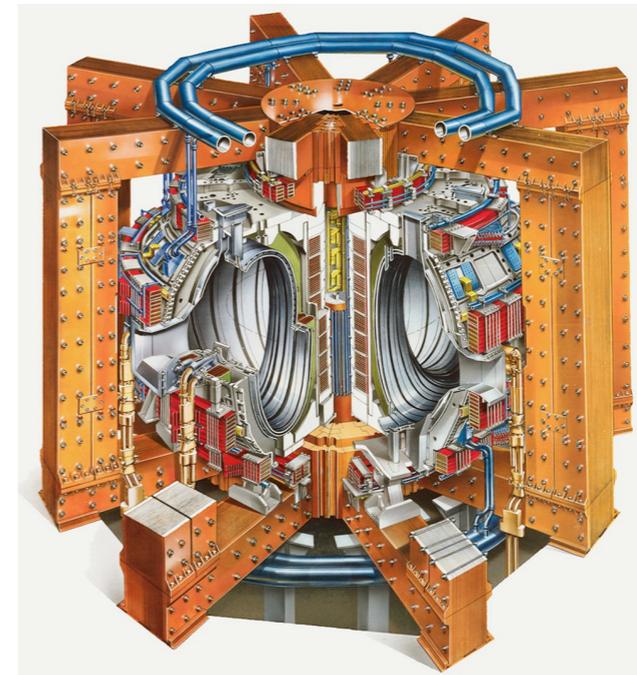
Stellar convection and dynamos



Periodic variation of sunspot

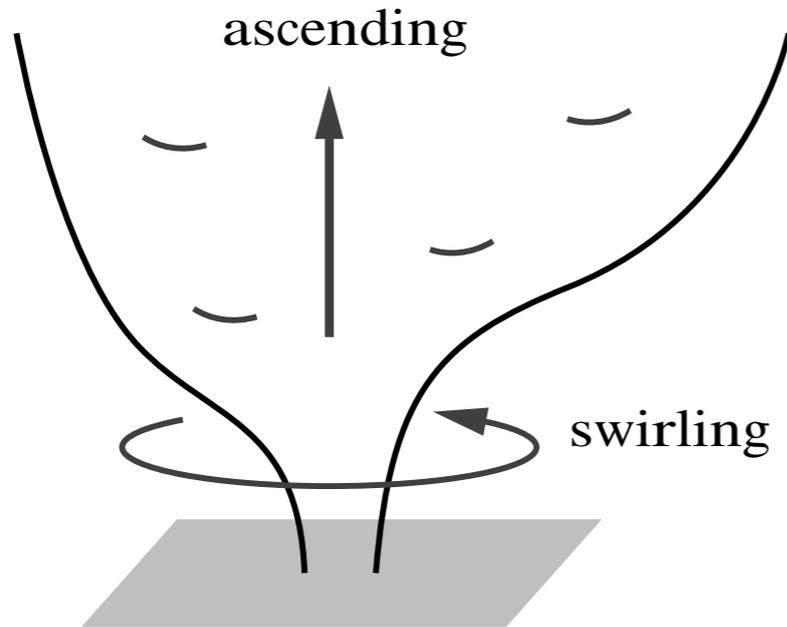


Polarity reversal of the geomagnetism

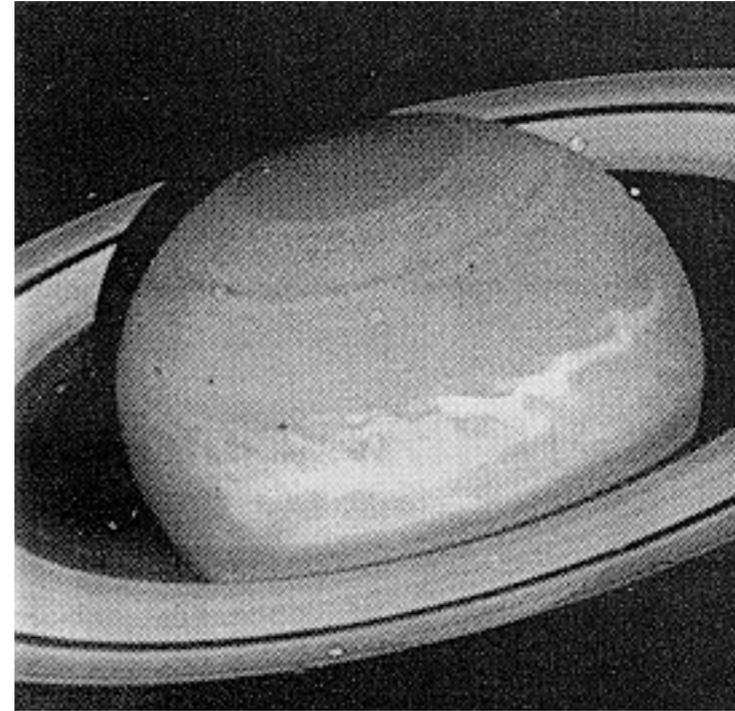


Joint European Torus (JET)

# Vortex structure, large-scale magnetic field



Typhoon, tornado



White spot on Saturn

Structure formation



Breakage of symmetry

What suppresses turbulent viscosity and/or anomalous resistivity?

Suppression of cascade

# Global magnetic fields of galaxies

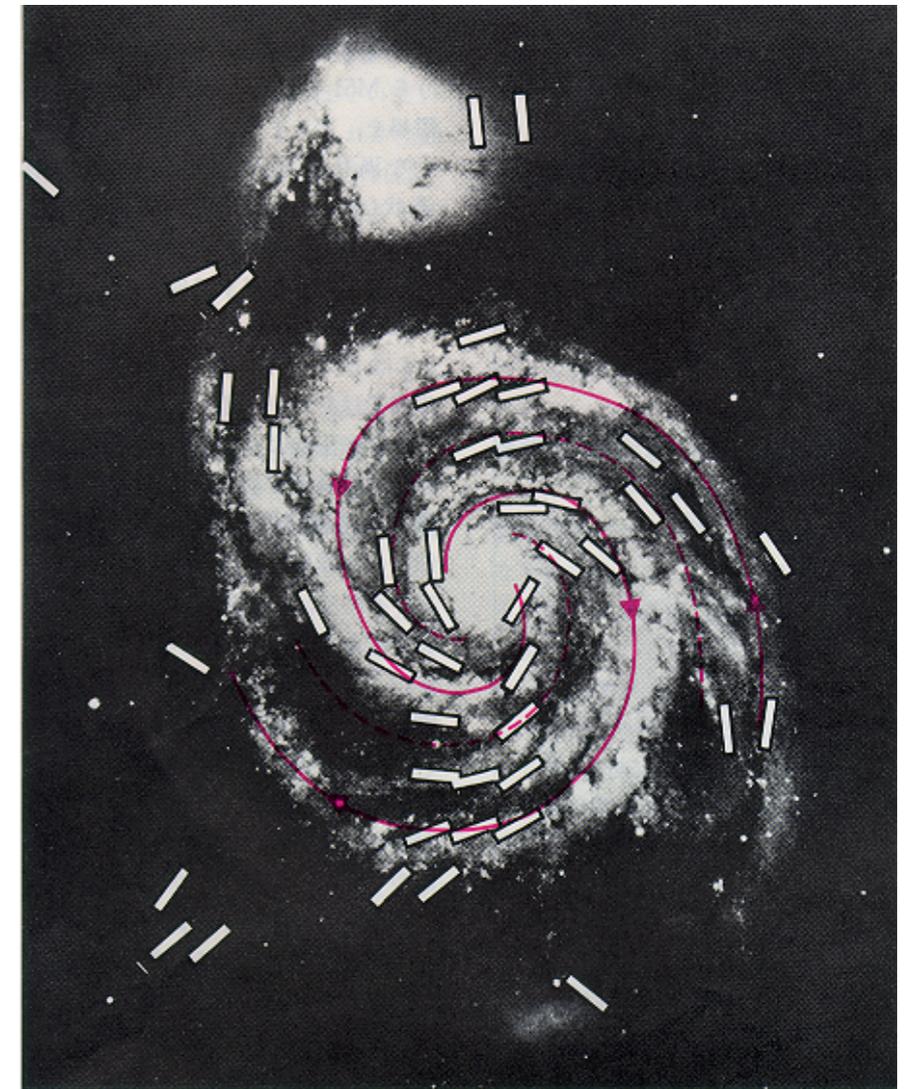
Galactic magnetic field

Toroidal magnetic field  
is observable

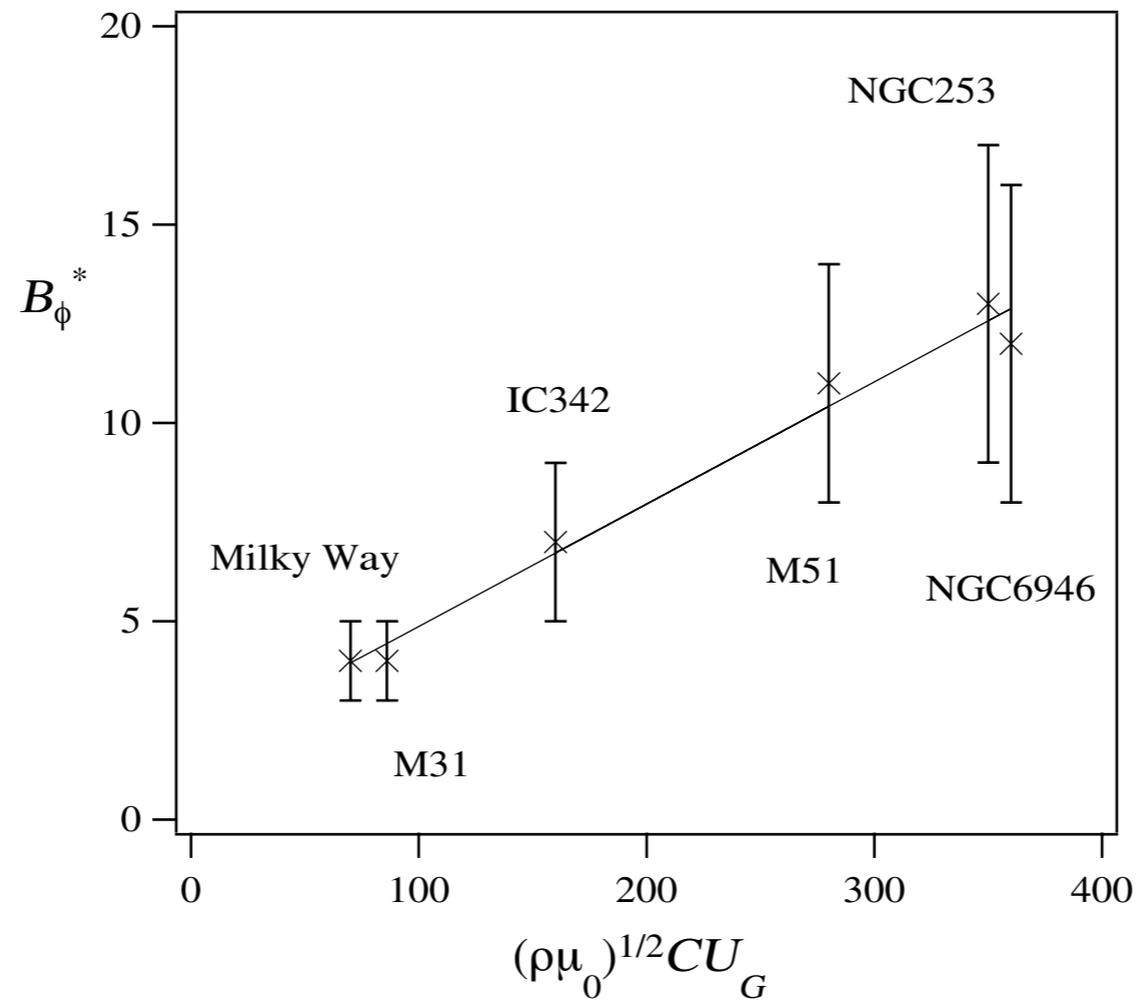
Magnetic-field strength  
a few  $\mu\text{G}$

Origin of galactic magnetic field

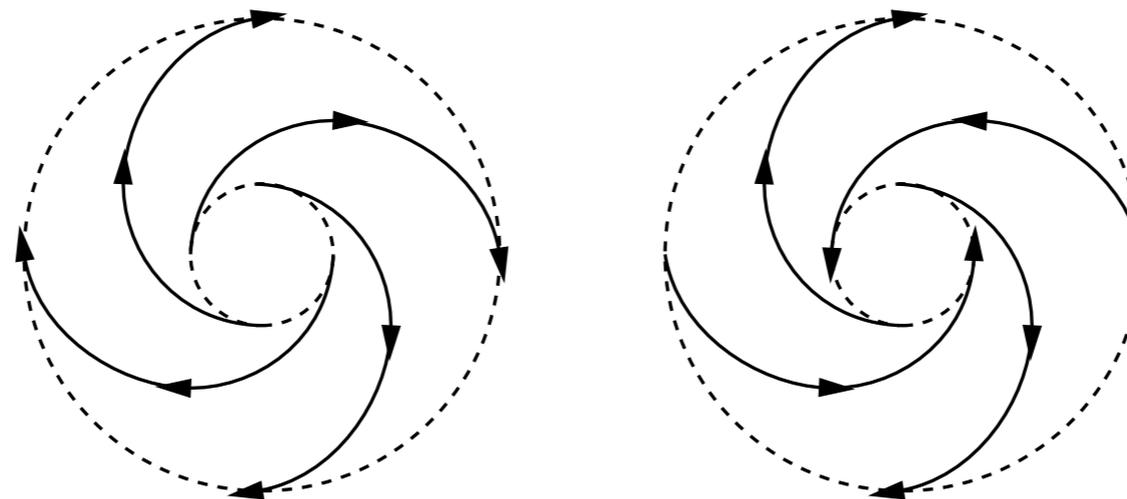
Primordial or Dynamo?



Magnetic field configuration in galaxy (M51)



Rotational speed and magnetic-field strength of galaxies

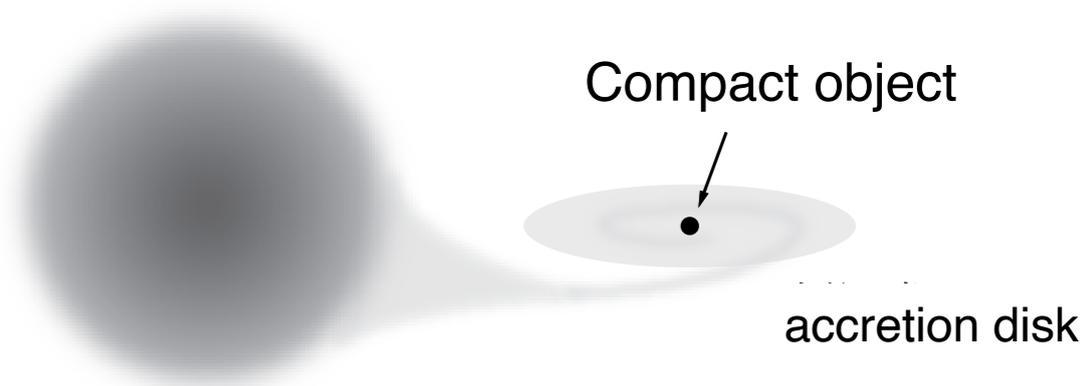


ASS (axisymmetric spiral) field BSS (bisymmetric spiral) field

# Accretion Disks

Plasmas  
accreting to central compact object  
with rotating

Accompany star



Astrophysical body	Compact object
Young stellar object	Protostar
Cataclysmic variables	White dwarf
X-ray binaries	Neutron star Black hole
Active galactic nuclei	Supermassive black hole

Conditions for accretion: Angular momentum loss

Turbulent viscosity

Alfvén wave

Jet

# Astrophysical jets

Bipolar jets

Mechanism for driving jet

Rotation (Vortical motion)



Toroidal magnetic-field generation  
due to the cross-helicity effect

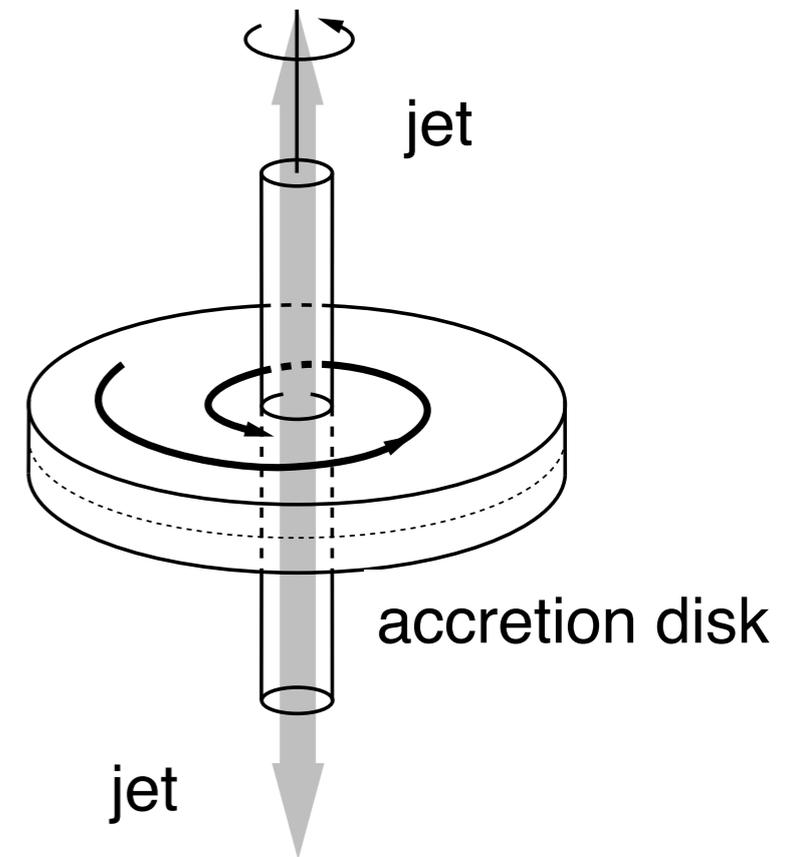


Driving jet by magnetic energy

High collimation

Collimation  $O(1)/O(10^6)$

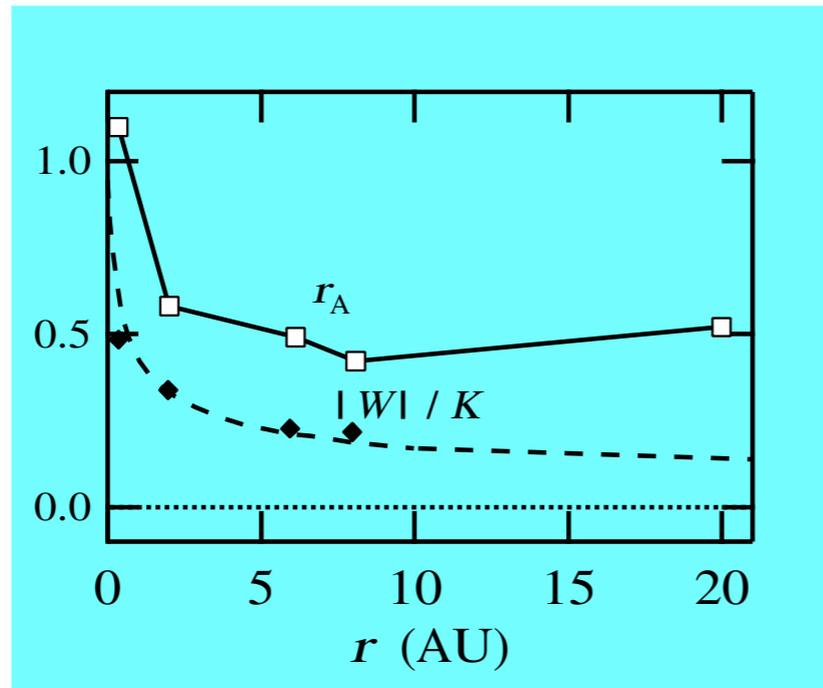
Magnetic confinement?



Jet from AGN 3C348



# Evolution of solar-wind turbulence



cross correlation and Alfvén ratio  
against heliocentric distance  
(Roberts *et al.* 1990)

## (i) Alfvén ratio

$$\text{Alfvén ratio } r_A \equiv \frac{\langle \mathbf{u}'^2 \rangle}{\langle \mathbf{b}'^2 \rangle} \simeq 0.5 \quad \text{for } r \gtrsim 3 \text{ AU}$$

Near the Sun

Equipartition  
(Alfvénic)



Far from the Sun

Magnetic dominance  
(super Alfvénic)

## (ii) Cross-helicity

Normalized cross helicity  $\frac{|W|}{K} = \frac{|\langle \mathbf{u}' \cdot \mathbf{b}' \rangle|}{\langle \mathbf{u}'^2 + \mathbf{b}'^2 \rangle / 2}$

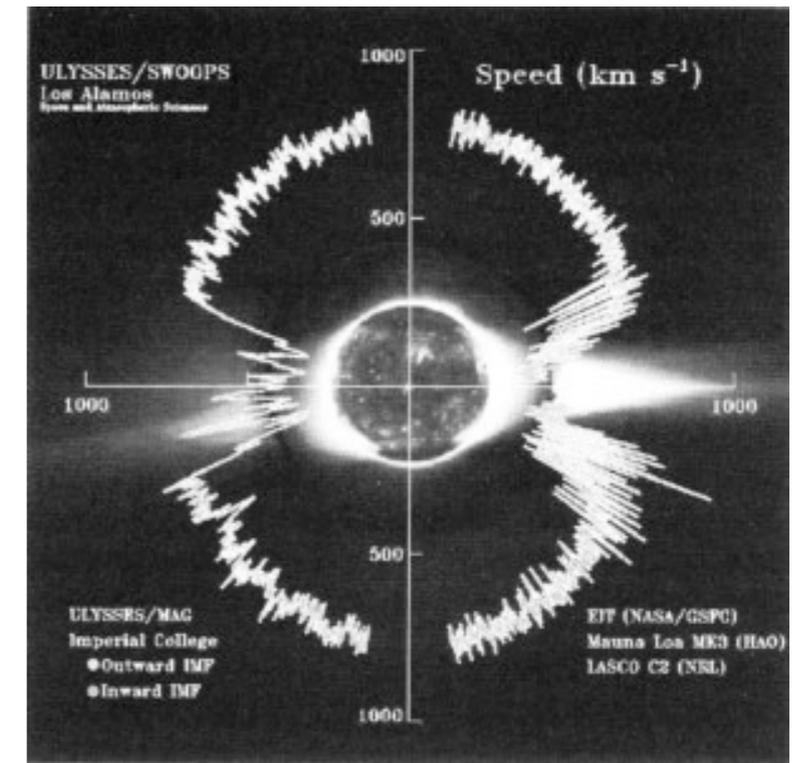
As the heliocentric distance increases

- **High-speed wind**  
from higher-latitude regions  
**weak velocity shear**

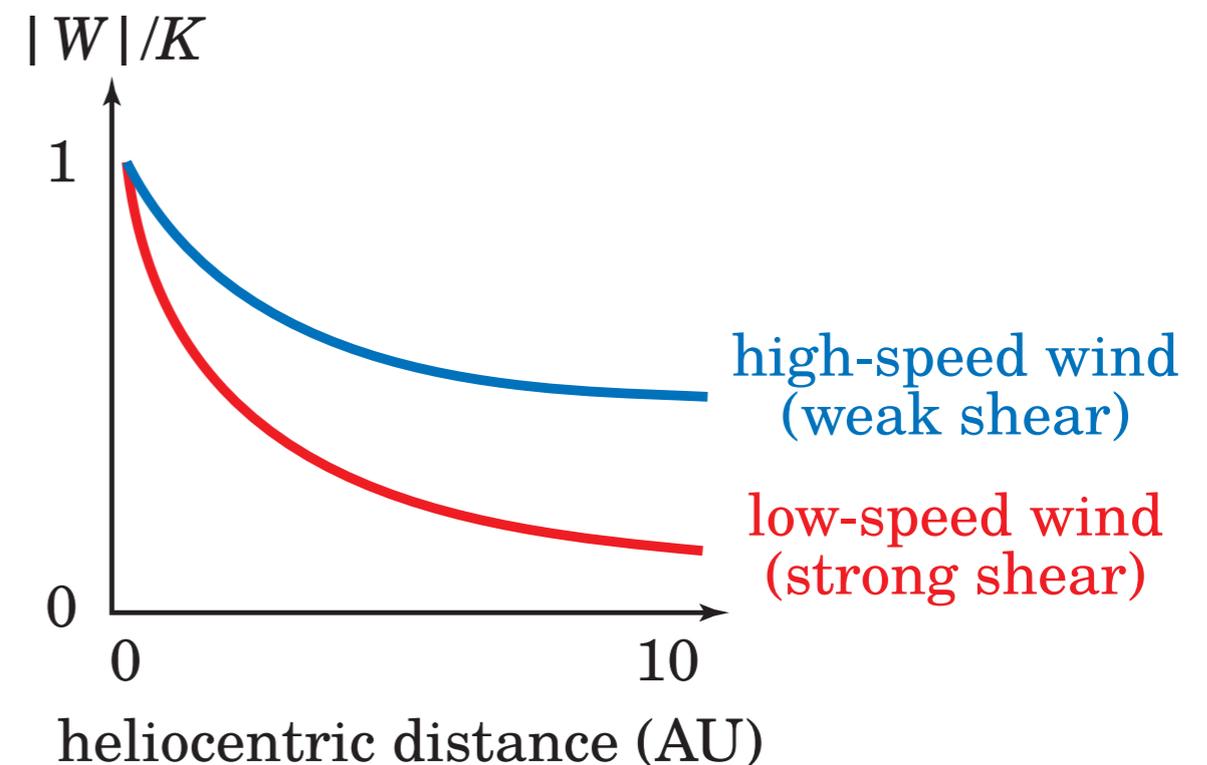
cross helicity remains to be relatively  
large value of **0.4~0.7**

- **Low-speed wind**  
from lower-latitude regions  
**strong velocity shear**

cross helicity decreases to  
small value of **0.1~0.3**



solar wind  
(Ulysses mission, 1991-1996)

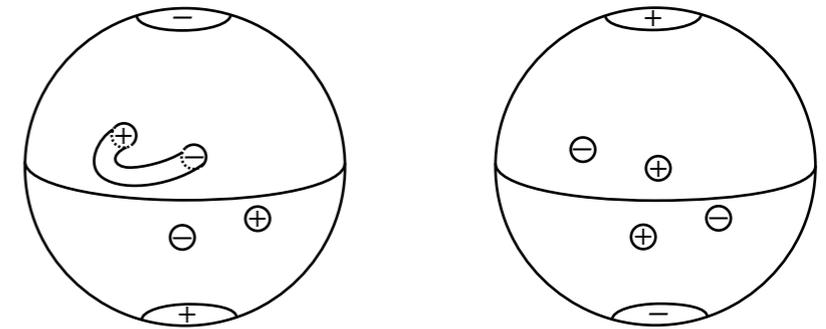


# Magnetic activity of the Sun

Solar sunspot

Emergence of strong magnetic field  
(thousands G)

Toroidal magnetic field

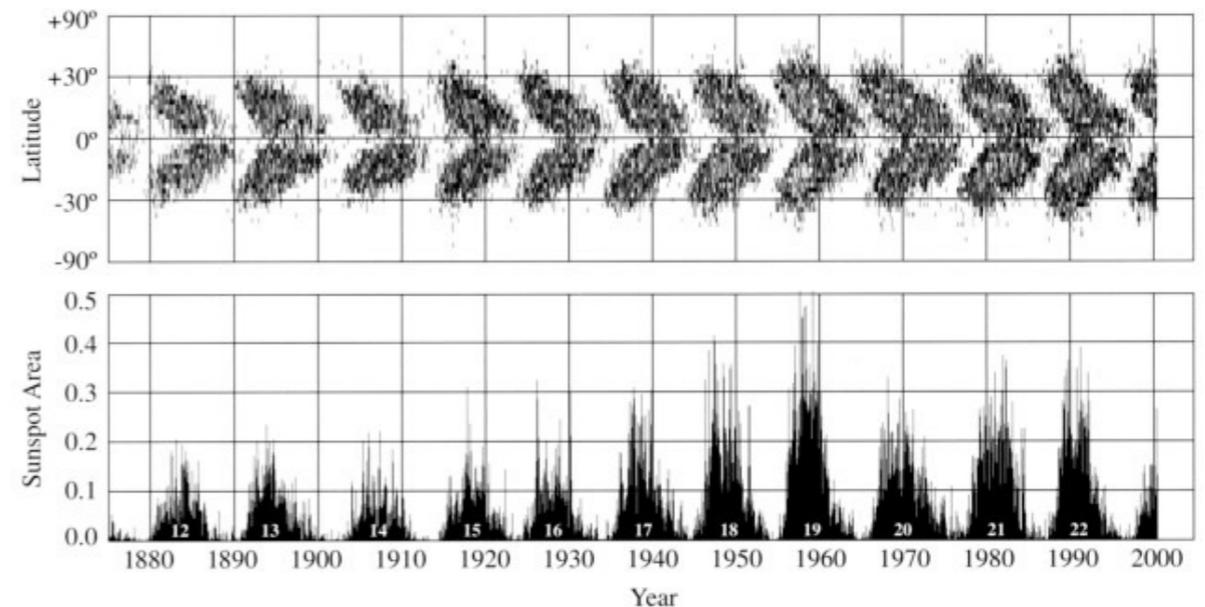


Polarity law of sunspot

Periodic behaviour  
of sunspot number

Periodic variation  
of magnetic field

Polarity reversal



Periodic variation of sunspot

What generates and sustains the solar magnetic field

← Fluid plasma motions in the Sun

# Solar interior

## Core

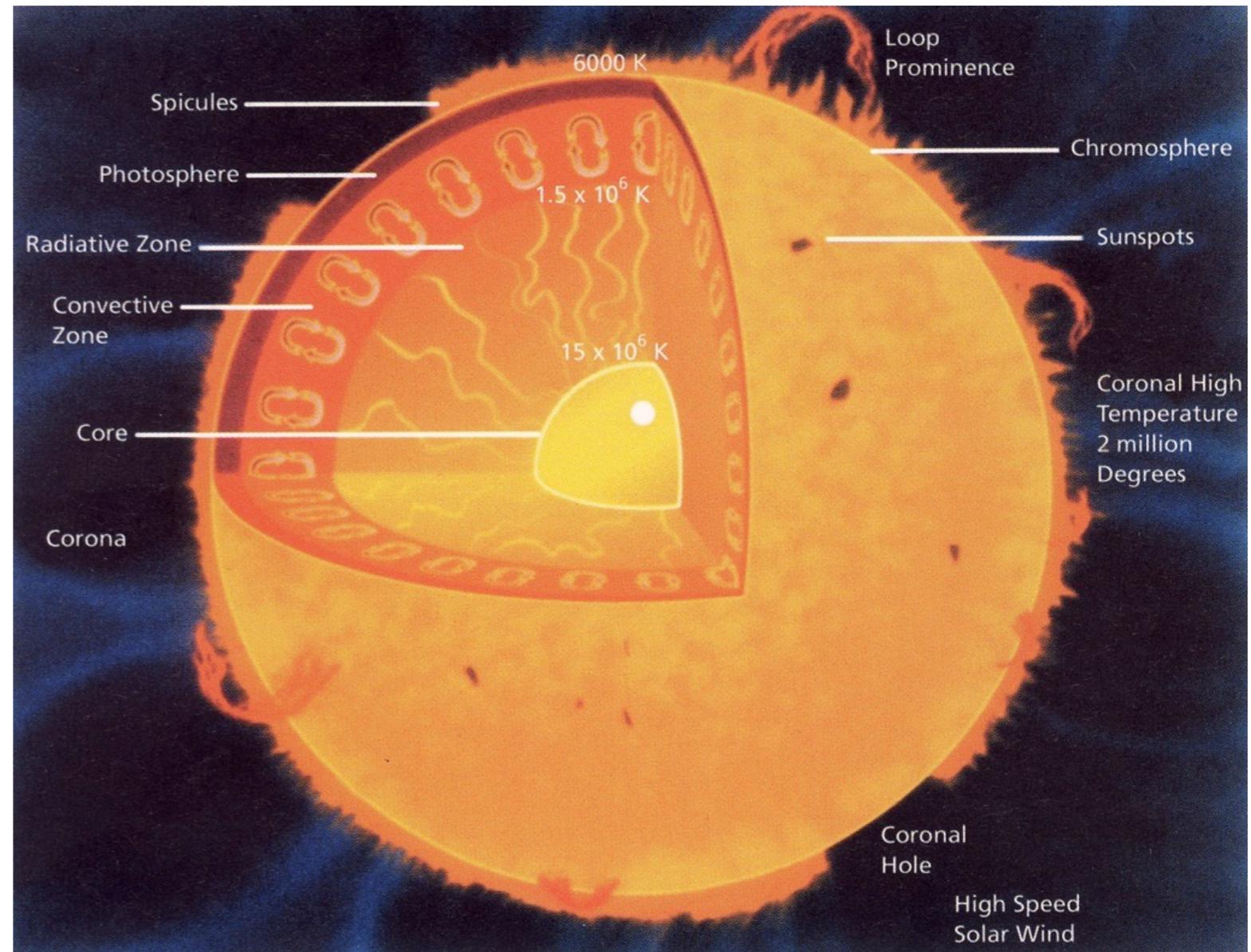
Thermal fusion

## Radiative zone

Approximately rigid rotation

## Convective zone

Differential rotation



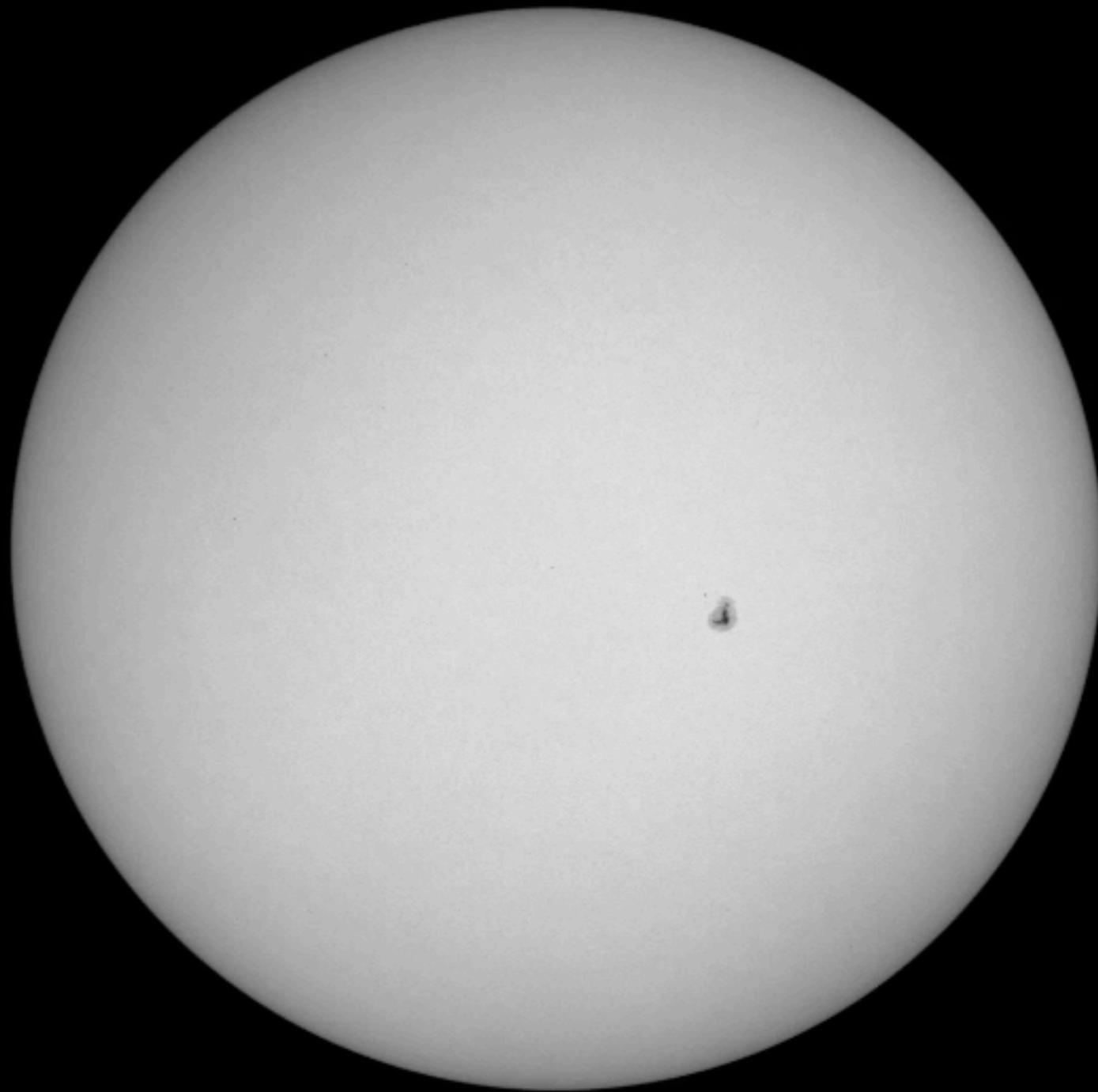
Motion in the convective zone



Turbulent electromotive force

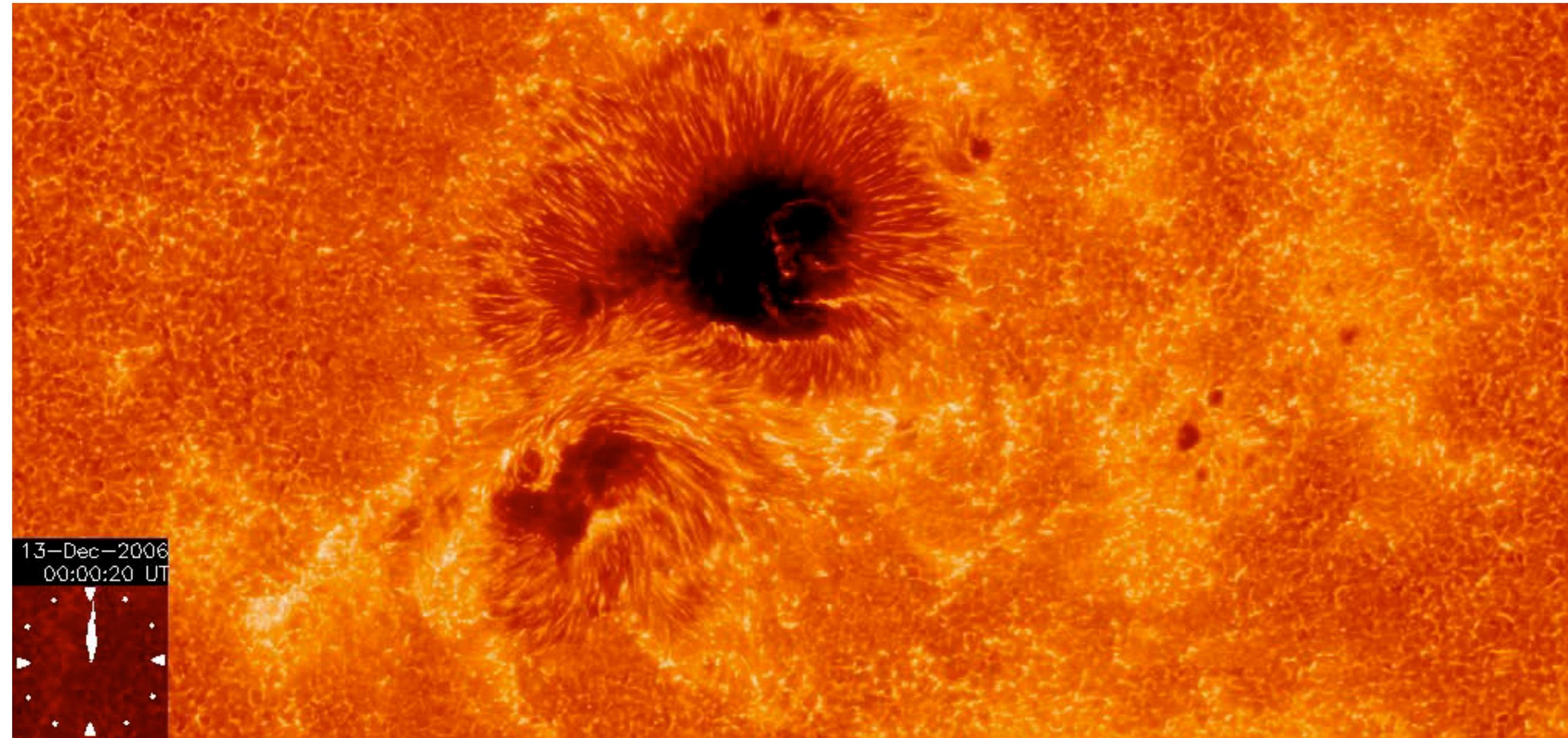


Solar magnetic field

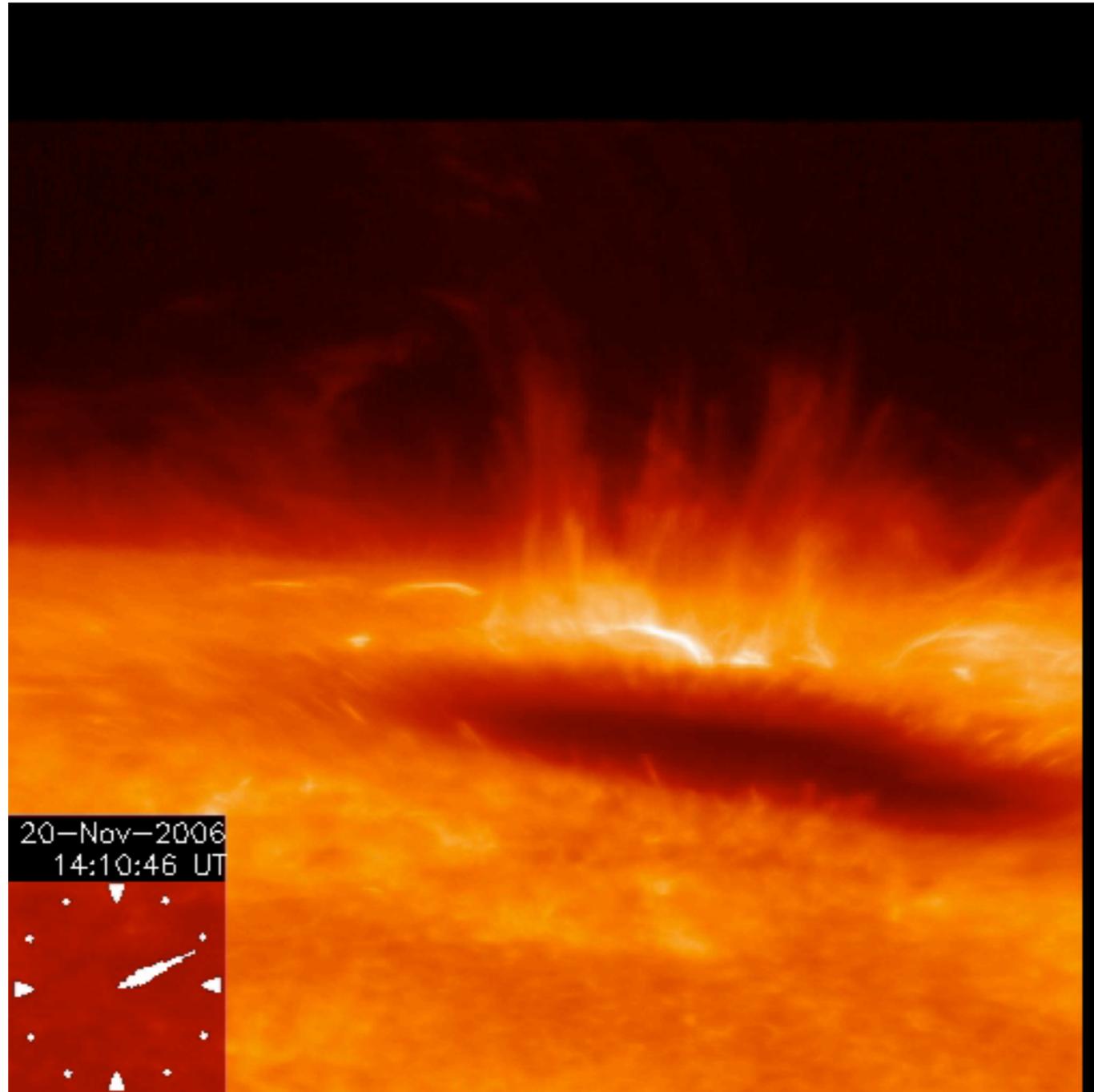


171,000 km

Okamoto,  
NAOJ



Courtesy of the HINODE Group, NAOJ



Courtesy of the HINODE Group, NAOJ

# Geomagnetism

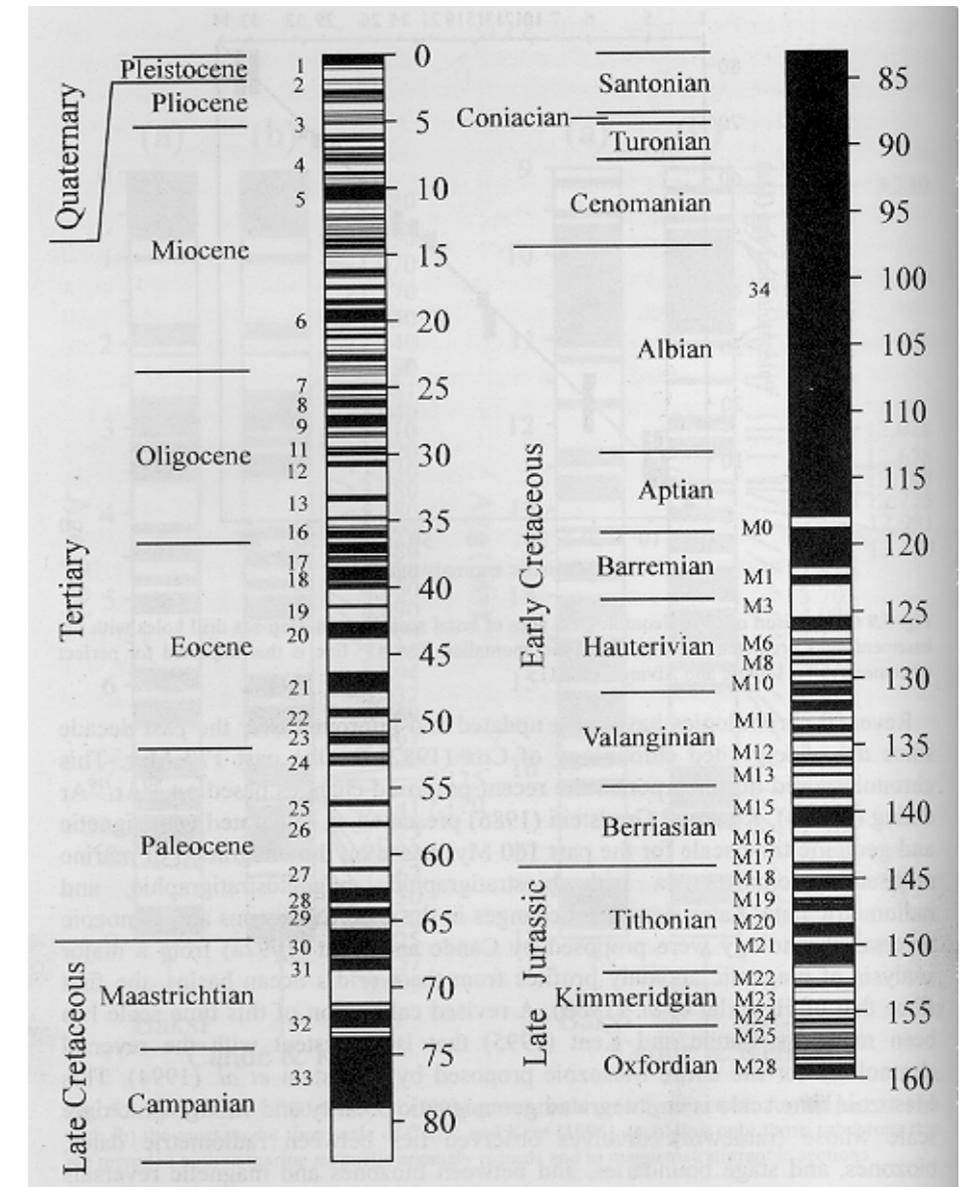
Polarity reversal of geomagnetic field

Irregular

Magnetic energy  $\gg$  Kinetic energy

Motion of melted iron  
in the outer core

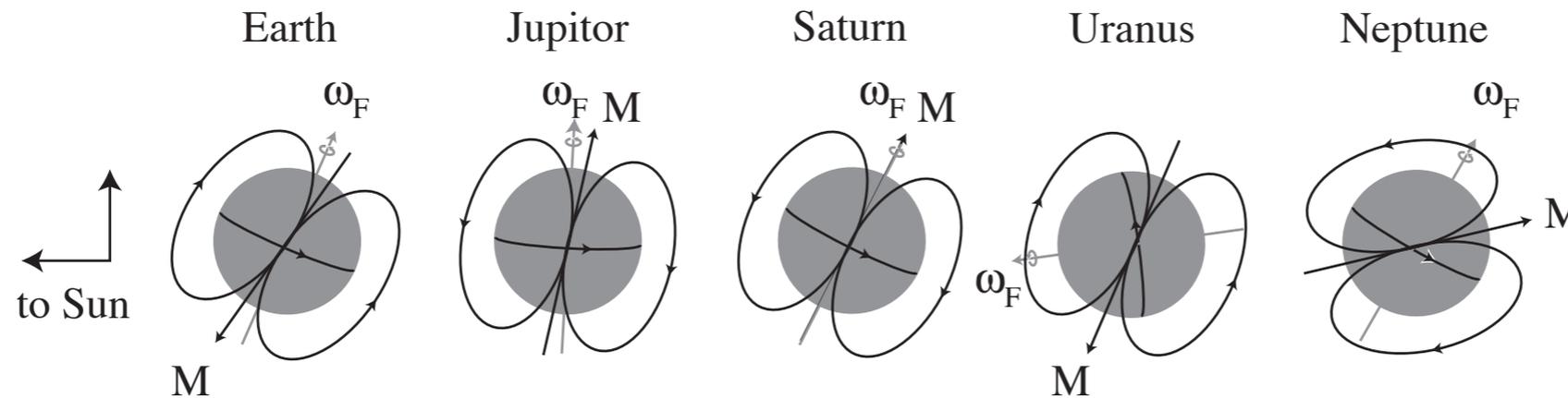
Only poloidal magnetic fields  
are observable



Polarity reversal of the geomagnetism

# Geomagnetism

## Planetary magnetic field

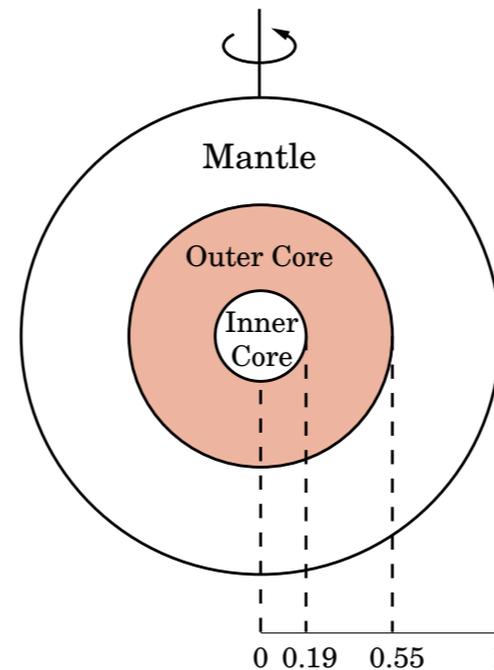


Directions of rotation axis ( $\omega_F$ ) and dipole moment (M)

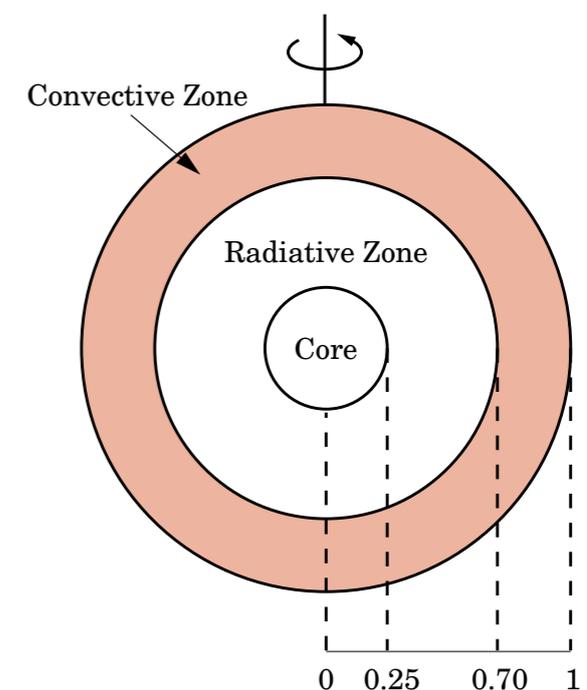
## Differences

Earth      Outer core  
 Melted iron      Thick shell  
 Fast rotation  
 Irregular reversal

Sun      Convection zone  
 Ionized H      Thin shell  
 Relatively slow rotation  
 Periodic reversal



Earth's outer core



Solar convective zone

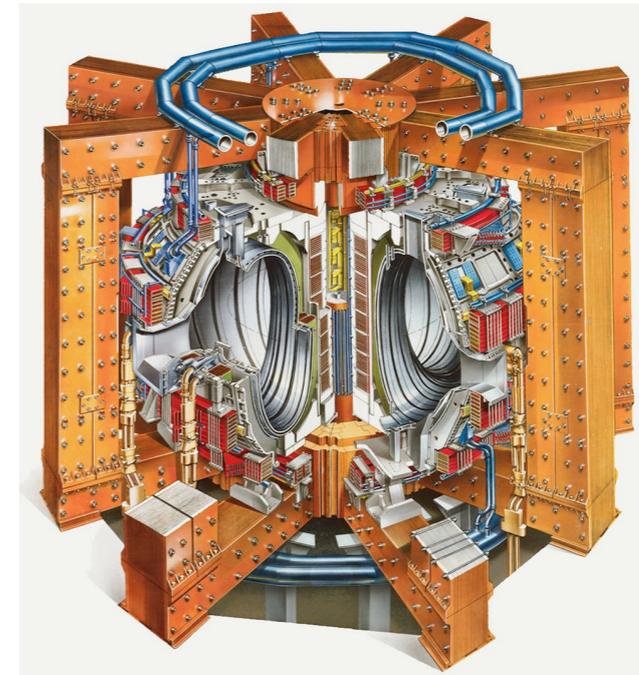
# Fusion plasmas

Reversed field pinch

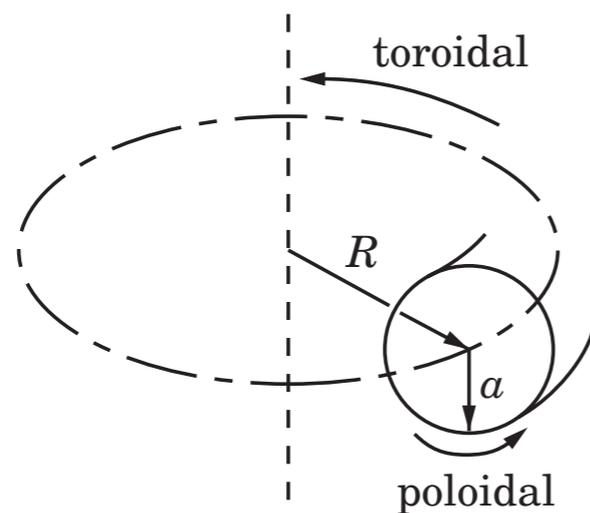
Tokamaks

Helical systems

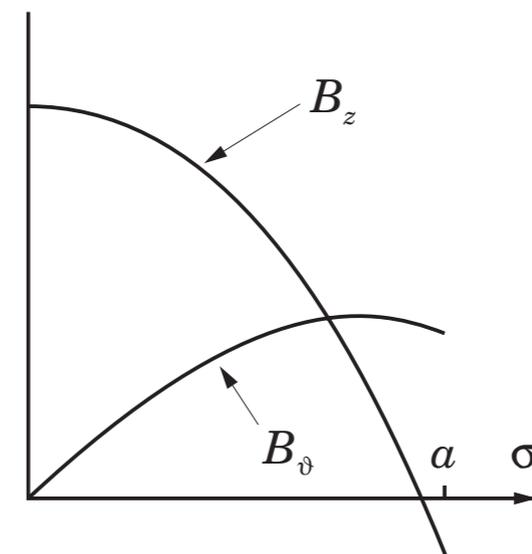
Inertial confinement



Joint European Torus (JET)



Poloidal and toroidal directions



Radial profiles of magnetic-field in RFP

# Mean-field equations in compressible MHD

Yokoi, N., J. Plasma Phys. **84**, 735840501 & 775840603 (2018a,b)

Density	$\frac{\partial \bar{\rho}}{\partial t} + \nabla \cdot (\bar{\rho} \mathbf{U}) = -\nabla \cdot \langle \rho' \mathbf{u}' \rangle$	Turb. mass flux
Momentum	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} U^\alpha + \frac{\partial}{\partial x^a} \bar{\rho} U^a U^\alpha \\ = -(\gamma_0 - 1) \frac{\partial}{\partial x^\alpha} \bar{\rho} Q + \frac{\partial}{\partial x^\alpha} \mu S^{a\alpha} + (\mathbf{J} \times \mathbf{B})^\alpha \\ - \frac{\partial}{\partial x^\alpha} \left( \underbrace{\bar{\rho} \langle u'^a u'^\alpha \rangle}_{\text{Reynolds stress}} - \frac{1}{\mu_0} \underbrace{\langle b'^a b'^\alpha \rangle}_{\text{Turb. Maxwell stress}} + U^a \underbrace{\langle \rho' u'^\alpha \rangle}_{\text{Turb. energy flux}} + U^\alpha \underbrace{\langle \rho' u'^a \rangle}_{\text{Turb. mass-energy correl.}} \right) + R_U^\alpha \end{aligned}$	
Internal energy	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} Q + \nabla \cdot (\bar{\rho} \mathbf{U} Q) = \nabla \cdot \left( \frac{\kappa}{C_V} \nabla Q \right) - \nabla \cdot (\bar{\rho} \langle q' \mathbf{u}' \rangle + Q \langle \rho' \mathbf{u}' \rangle + \mathbf{U} \langle \rho' q' \rangle) \\ - (\gamma_0 - 1) \left( \bar{\rho} Q \nabla \cdot \mathbf{U} + \underbrace{\bar{\rho} \langle q' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. energy dilatation}} + Q \underbrace{\langle \rho' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. mass dilatation}} \right) + R_Q \end{aligned}$	
Magnetic field	$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B} + \langle \mathbf{u}' \times \mathbf{b}' \rangle) + \eta \nabla^2 \mathbf{B}$	Turb. electromotive force

# Contents

- **Turbulence properties**
- **Theoretical formulation**
  - **Non-linearity: Closure with response functions**
  - **Inhomogeneities, anisotropies, non-equilibrium properties: Multiple-scale analysis**
- **Illustrative examples**
  - **Angular-momentum transport by inhomogeneous kinetic helicity**
  - **Convective transport**
    - **Plumes as coherent fluctuations**
    - **Non-equilibrium properties along coherent motions**
  - **Dynamos**
- **Turbulence modelling**

# Turbulence properties



群盲撫象 葛飾北斎 Hokusai

# Turbulence

Nonlinear

Cascade

Dissipative

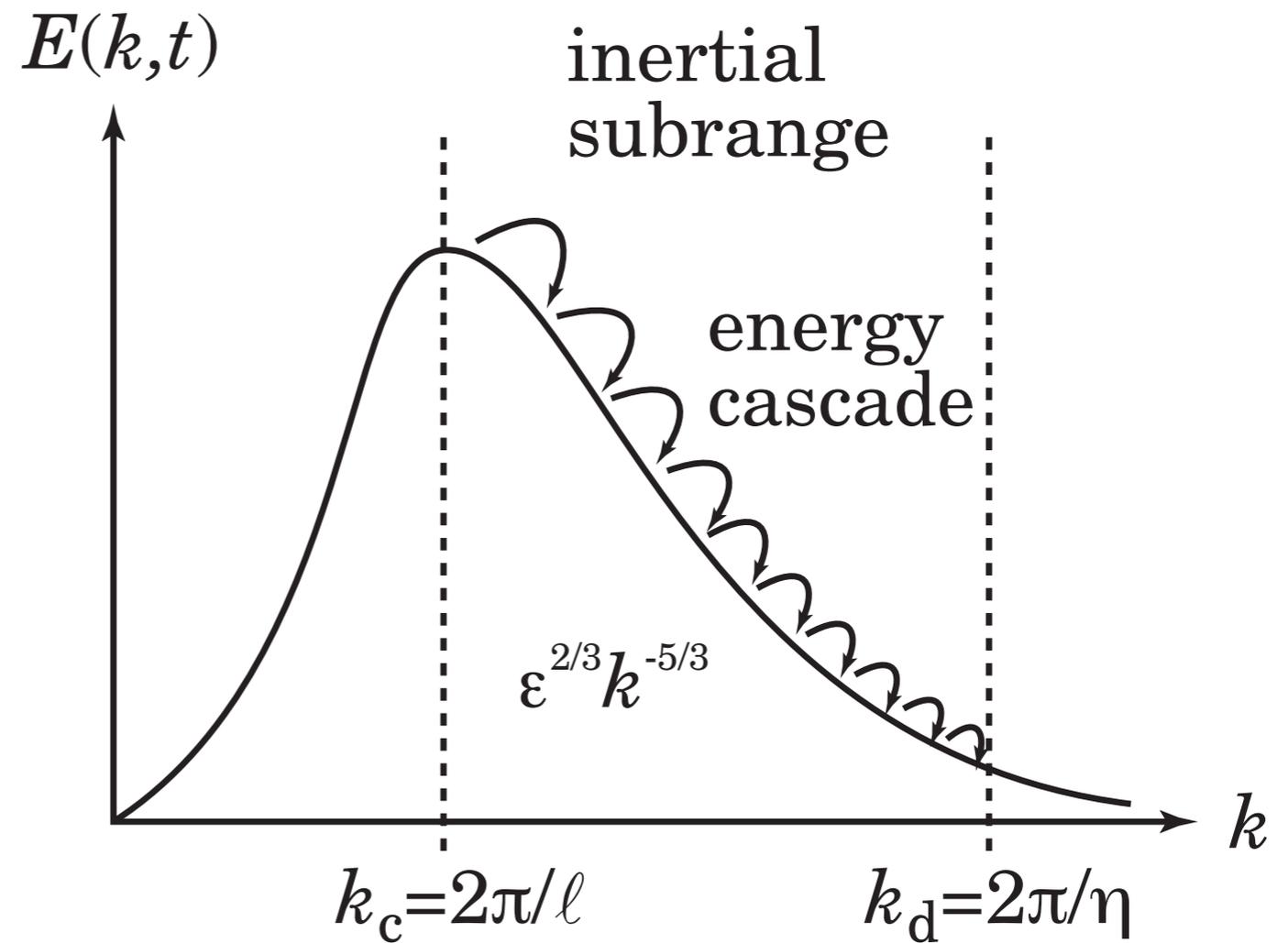
Anomalous transport

Local vs. non-local

Homogeneous  
vs. inhomogeneous

Isotropic vs. anisotropic

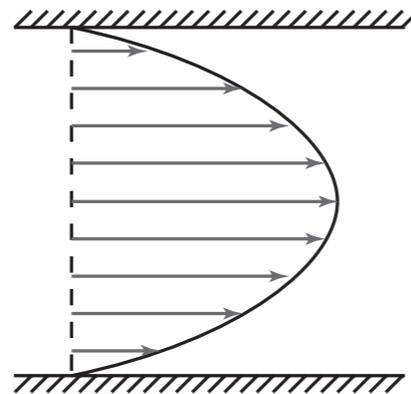
Equilibrium vs. non-equilibrium



Simple picture of cascade

# Effects of turbulence

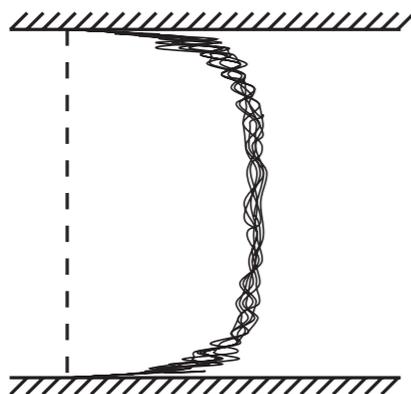
Laminar pipe flow



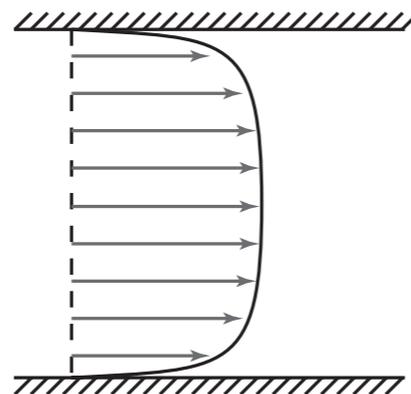
$$\left( \frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) \mathbf{u} = -\frac{1}{\rho} \nabla p + \nu \nabla^2 \mathbf{u}$$

Turbulent pipe flow

Large-scale structure destroyed



Instantaneous



Mean

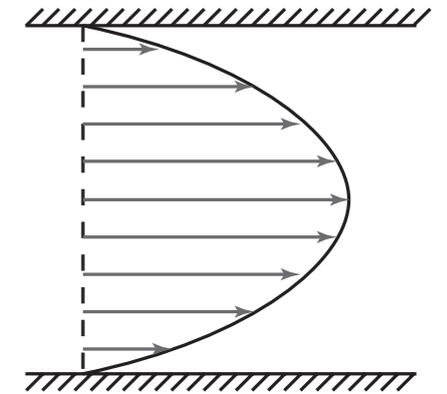
$$\begin{cases} \mathbf{u} = \mathbf{U} + \mathbf{u}', & \mathbf{U} = \langle \mathbf{u} \rangle \\ p = P + p', & P = \langle p \rangle \end{cases}$$

# Enhancement of transport

Equation of mean velocity

$$\frac{DU_\alpha}{Dt} \equiv \left( \frac{\partial}{\partial t} + \frac{\partial}{\partial x_a} \right) U_\alpha = -\frac{\partial P}{\partial x_\alpha} - \frac{\partial}{\partial x_a} \langle u'_a u'_\alpha \rangle + \nu \frac{\partial^2 U_\alpha}{\partial x_a^2}$$

Laminar



Reynolds stress  $\langle u'_\alpha u'_\beta \rangle = \frac{2}{3} K \delta_{\alpha\beta} - \nu_T \left( \frac{\partial U_\alpha}{\partial x_\beta} + \frac{\partial U_\beta}{\partial x_\alpha} \right)$  (Model)

$\nu_T$  : eddy viscosity (turbulent viscosity)

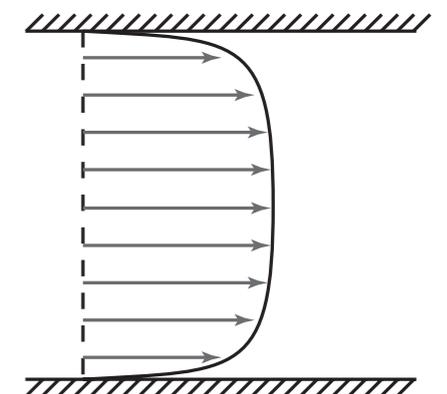
(Boussinesq, 1877)

$$\longrightarrow \frac{\partial U_\alpha}{\partial t} + U_a \frac{\partial U_\alpha}{\partial x_a} = -\frac{\partial P}{\partial x_\alpha} + \frac{\partial}{\partial x_a} \left[ (\nu + \nu_T) \left( \frac{\partial U_\alpha}{\partial x_a} + \frac{\partial U_a}{\partial x_\alpha} \right) \right]$$

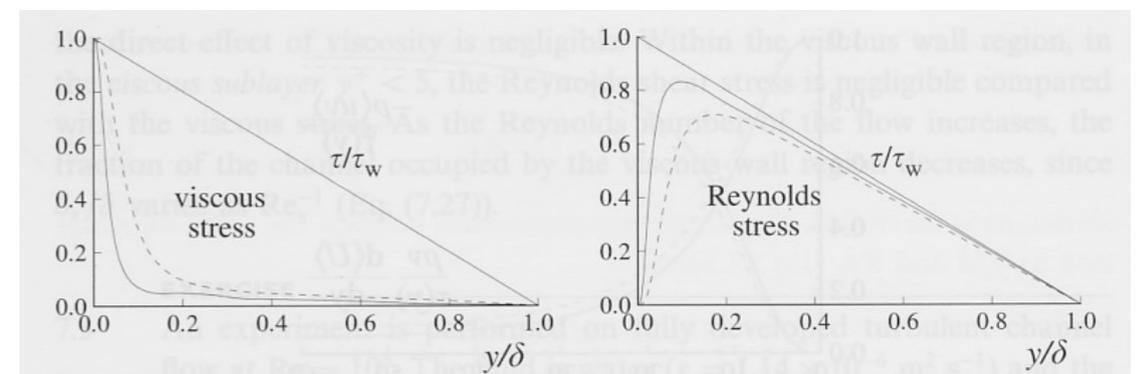
## Enhancement of transport

$$\frac{\text{(Turbulent viscosity)}}{\text{(Molecular viscosity)}} \quad \frac{\nu_T}{\nu} \sim \frac{u\ell}{\nu} \sim Re$$

Turbulent



DNS of turbulent and viscous stresses



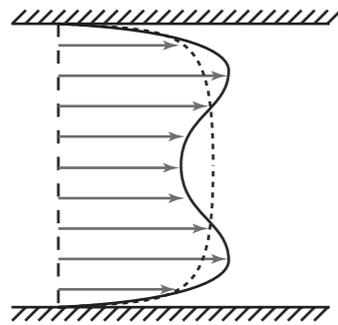
dashed line, Re=5,600; solid line Re=13,750

## Variation in space and time

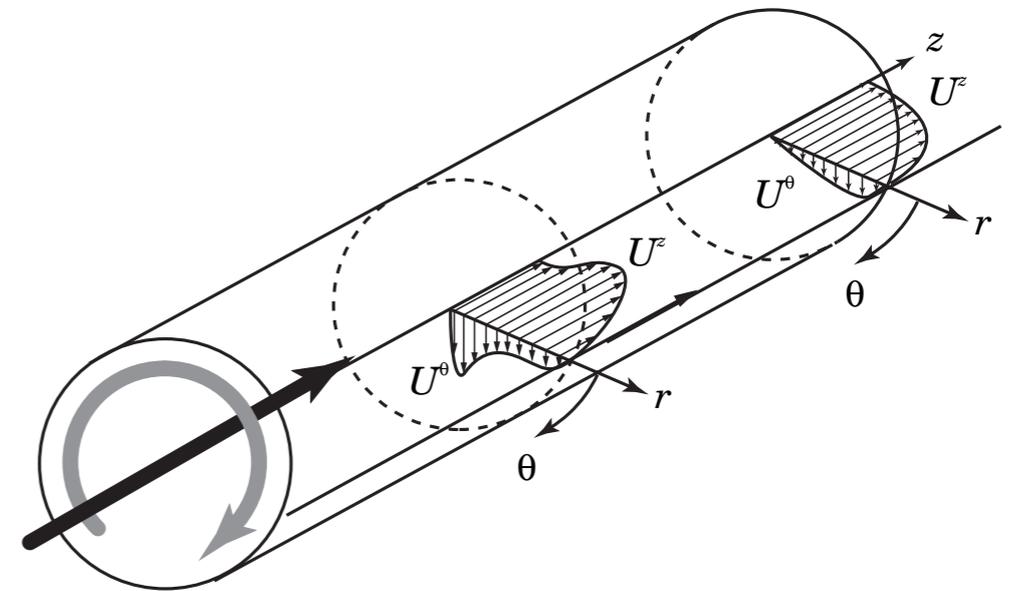
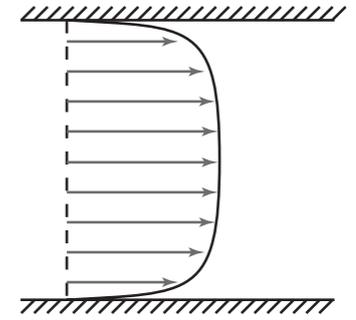
# Suppression of transport

## Swirling flow in a circular pipe

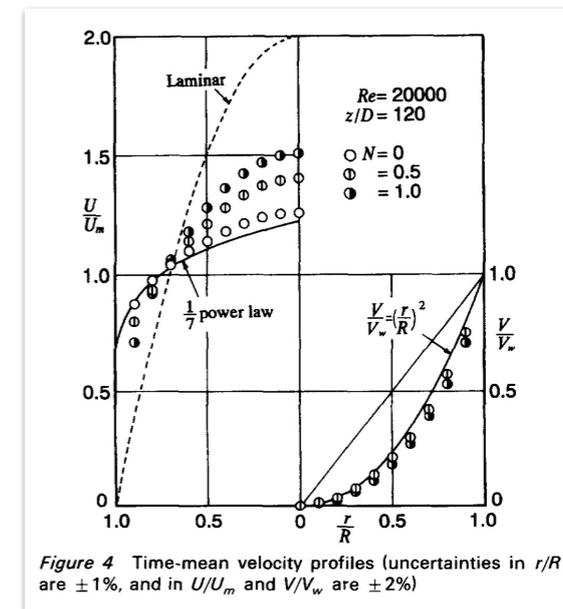
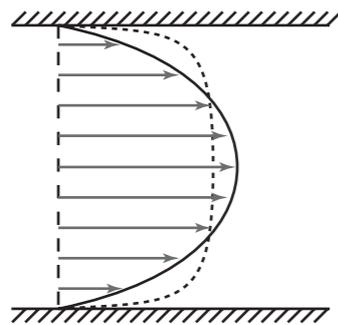
### Turbulent swirling pipe flow



Turbulent



### Axially rotating turbulent pipe flow

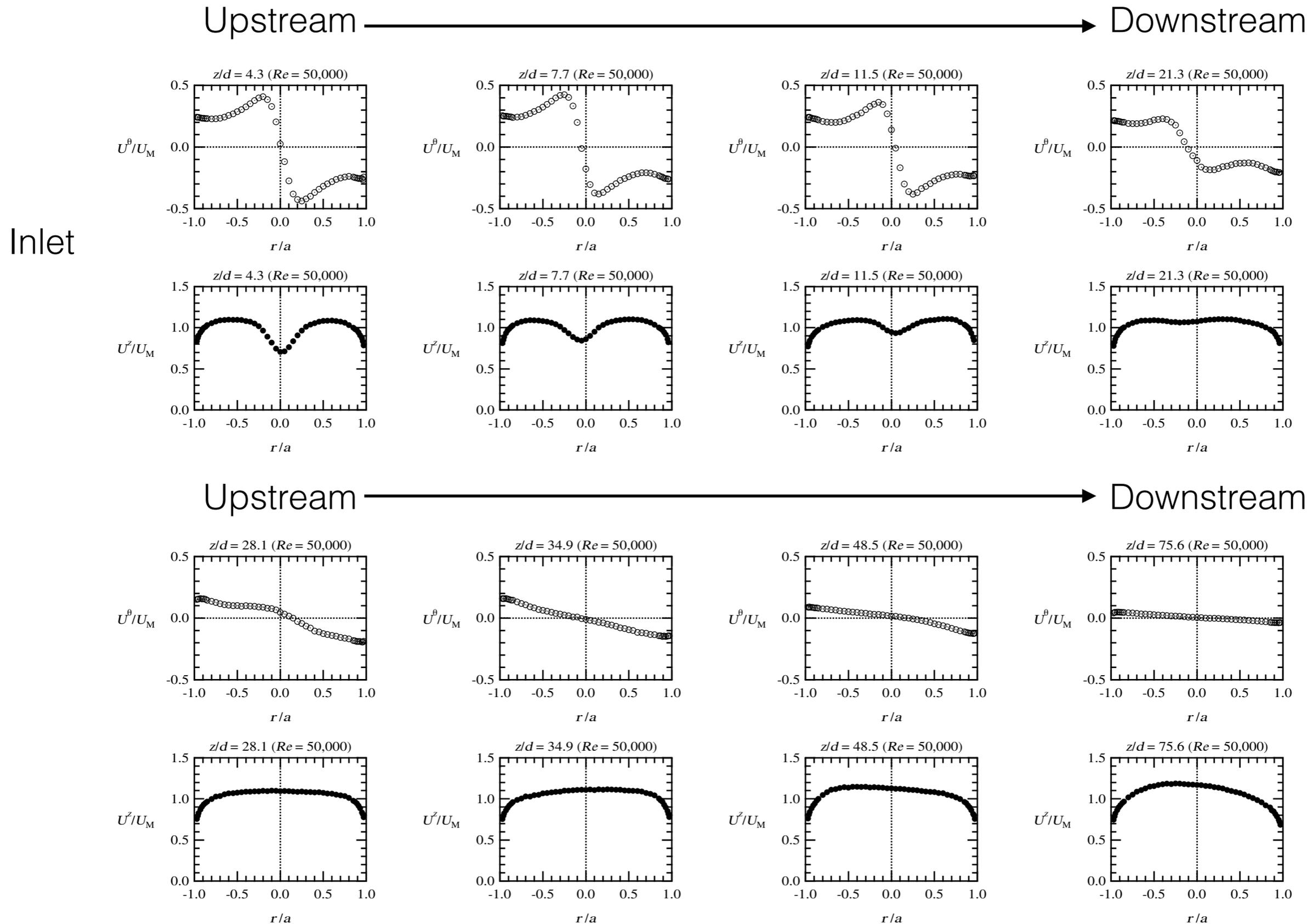


These flow properties cannot be reproduced by the standard eddy-viscosity representation at all. Too much dissipative.

(Imao et al. 1996)

# Axial evolution of a weak swirl

(Steenbergen, 1995)



# Equation of fluctuating velocity

$$\mathbf{u} = \mathbf{U} + \mathbf{u}', \quad \mathbf{U} = \langle \mathbf{u} \rangle, \quad \mathbf{u}' = \mathbf{u} - \langle \mathbf{u} \rangle$$

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial U_\alpha}{\partial x_a}}_{\text{turbulence-mean velocity interaction}} \underbrace{- u'_a \frac{\partial u'_\alpha}{\partial x_a} + \frac{\partial}{\partial x_a} \langle u'_a u'_\alpha \rangle}_{\text{turbulence-turbulence interaction}} - \frac{\partial p'}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

turbulence-mean velocity interaction  
turbulence-turbulence interaction



## Instability approach

Quasi-linear -> nonlinearity

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial U_\alpha}{\partial x_a}}_{\text{turbulence-mean velocity interaction}} - \frac{\partial p'^{(R)}}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

Linear in  $\mathbf{u}'$  and  $p'^{(R)}$ ; each (Fourier) mode evolves independently



## Closure approach

Homogeneous isotropic -> inhomogeneities

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial u'_\alpha}{\partial x_a} + \frac{\partial}{\partial x_a} \langle u'_a u'_\alpha \rangle}_{\text{turbulence-turbulence interaction}} - \frac{\partial p'^{(S)}}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

Homogeneous turbulence, no dependence on large-scale inhomogeneity

# Navier–Stokes equation in the wave-number space

Homogeneous and  
Isotropic Turbulence  
(HIT)

$$\frac{\partial \hat{u}_\alpha(\mathbf{k}; t)}{\partial t} = -ik_\alpha \iint d\mathbf{p} d\mathbf{q} \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) \hat{u}_\alpha(\mathbf{p}; t) \hat{u}_\alpha(\mathbf{q}; t) = ik_\alpha \hat{p}(\mathbf{k}; t) - \nu k^2 \hat{u}_\alpha(\mathbf{k}; t)$$

The dynamics of  $\mathbf{k}$  mode is governed by its interaction with all other modes



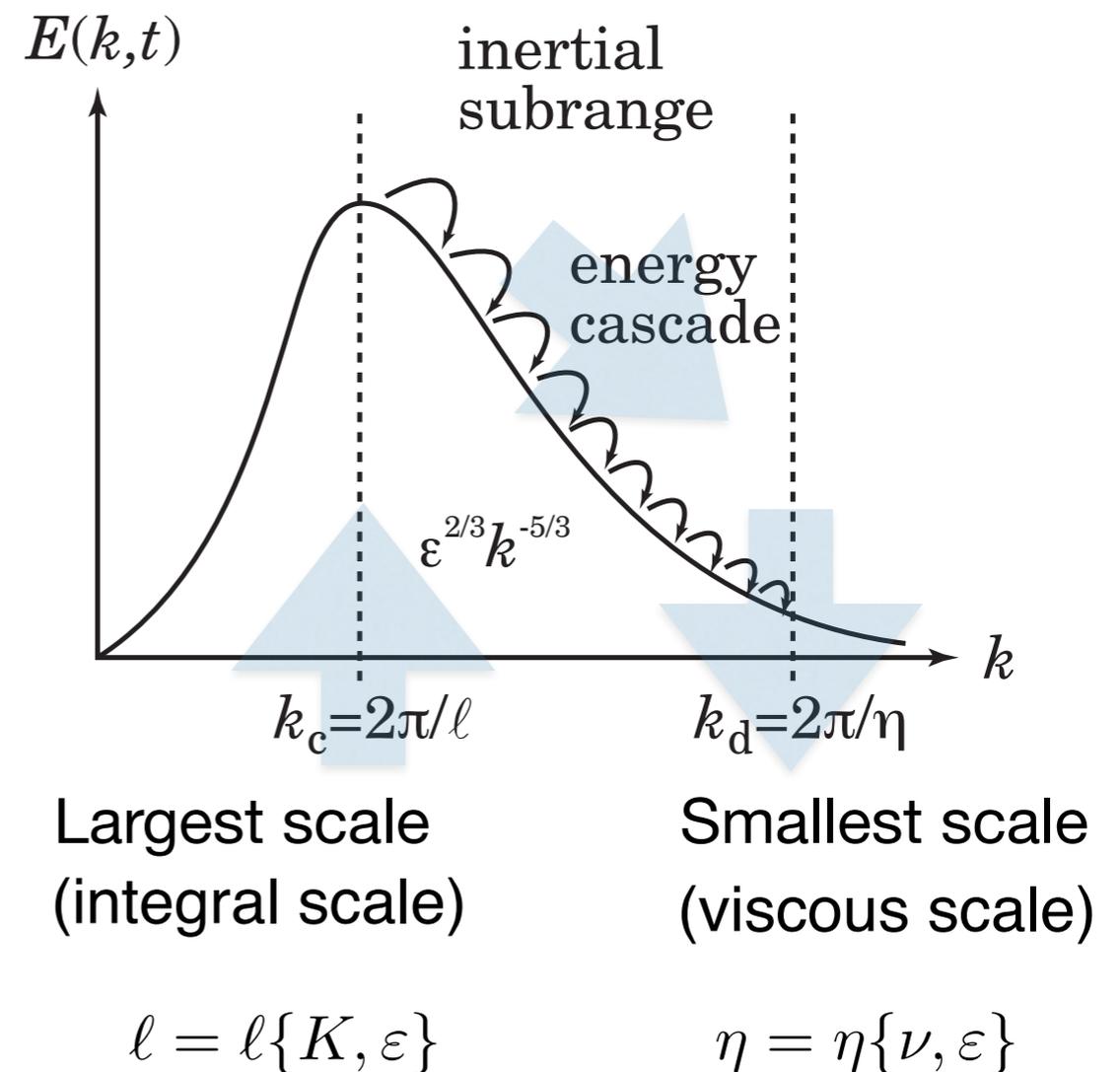
All the scales, from the largest to smallest scales, have to be solved

Energy injection

Energy flux

Energy dissipation

$\mathcal{E}$



## Kolmogorov microscale $\eta = \eta\{\nu, \varepsilon\}$

$$u_\eta \sim \frac{\nu}{\eta} \quad \left[ Re = \frac{u_\eta \eta}{\nu} \sim 1 \right]$$

$$\eta \sim \left( \frac{\nu^3}{\varepsilon} \right)^{1/4}$$



$$\varepsilon \sim \nu \frac{u_\eta^2}{\eta^2} \sim \nu \frac{(\nu/\eta)^2}{\eta^2} \sim \frac{\nu^3}{\eta^4}$$

$$u_\eta = (\varepsilon \nu)^{1/4}$$

$$u_\eta \sim \frac{\nu}{\eta} \sim \frac{\nu}{(\nu^3/\varepsilon)^{1/4}} \sim (\varepsilon \nu)^{1/4}$$

$$\tau_\eta = \left( \frac{\nu}{\varepsilon} \right)^{1/2}$$

$$\tau_\eta \sim \frac{\eta}{u_\eta} \sim \frac{(\nu^3/\varepsilon)^{1/4}}{(\varepsilon \nu)^{1/4}} \sim \left( \frac{\nu}{\varepsilon} \right)^{1/2}$$

## Integral scale

$$l = l\{K, \varepsilon\}$$

$$l \sim \frac{u_l^3}{\varepsilon} \sim \frac{K^{3/2}}{\varepsilon}$$



$$\varepsilon \sim \frac{u_l^2}{l/u_l} \sim \frac{u_l^3}{l} \quad Re = \frac{u_l l}{\nu}$$

$$u_l \sim K^{1/2}$$

$$\tau_l \sim \frac{K}{\varepsilon}$$

$$\varepsilon \sim \frac{K}{\tau_l}$$

# Scale differences

Length

$$\frac{\eta}{\ell} \sim Re^{-3/4}$$

$$\frac{\eta}{\ell} \sim \frac{(\nu^3/\varepsilon)^{1/4}}{K^{3/2}/\varepsilon} \sim \frac{(\varepsilon\nu)^{3/4}}{K^{3/2}} \sim \left(\frac{\varepsilon\nu}{K^2}\right)^{3/4} \sim \left(\frac{\nu}{u_\ell \ell}\right)^{3/4} \sim Re^{-3/4}$$

Velocity

$$\frac{u_\eta}{u_\ell} \sim Re^{-1/4}$$

$$\frac{u_\eta}{u_\ell} \sim \frac{(\varepsilon\nu)^{1/4}}{K^{1/2}} \sim \left(\frac{\varepsilon\nu}{K^2}\right)^{1/4} \sim \left(\frac{\nu}{u_\ell \ell}\right)^{1/4} \sim Re^{-1/4}$$

Time

$$\frac{\tau_\eta}{\tau_\ell} \sim Re^{-1/2}$$

$$\frac{\tau_\eta}{\tau_\ell} \sim \frac{(\nu/\varepsilon)^{1/2}}{K/\varepsilon} \sim \left(\frac{\varepsilon\nu}{K^2}\right)^{1/2} \sim \left(\frac{\nu}{u_\ell \ell}\right)^{1/2} \sim Re^{-1/2}$$

$$\left[ \nu_T \sim u_\ell \ell, \quad \frac{\nu_T}{\nu} \sim \frac{u_\ell \ell}{\nu} \sim Re \right]$$

# Largest and smallest scales in turbulence

Largest scale (integral scale)  $\ell = \ell\{K, \varepsilon\}$   $\ell \sim \frac{u^3}{\varepsilon} \sim \frac{K^{3/2}}{\varepsilon}$

Smallest scale (viscous scale)  $\eta = \eta\{\nu, \varepsilon\}$   $\eta \sim \left(\frac{\nu^3}{\varepsilon}\right)^{1/4}$

$$\frac{\text{(Largest scale)}}{\text{(Smallest scale)}} \quad \frac{\ell}{\eta} \sim \left(\frac{u\ell}{\nu}\right)^{3/4} = O\left(Re^{3/4}\right)$$

Required grid points

$$N_G = \left(\frac{\ell}{\eta}\right)^3 = O\left(Re^{9/4}\right)$$

	Re	$N_G$		Re	$N_G$
Walking	$O(10^4)$	$O(10^9)$	↑	Earth's outer core	$O(10^8)$ $O(10^{18})$
Cars	$O(10^6)$	$O(10^{13.5})$	↑	Solar convection zone	$O(10^{10})$ $O(10^{22.5})$
Airplanes	$O(10^8)$	$O(10^{18})$		Galaxies	$O(10^{11})$ $O(10^{25})$

**Direct numerical simulation (DNS) is just impossible in the foreseeable future**

# **Theoretical formulation**

# Approaches to turbulence

# Equation of fluctuating velocity

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial U_\alpha}{\partial x_a}}_{\text{turbulence-mean interaction}} \underbrace{- u'_a \frac{\partial u'_\alpha}{\partial x_a} + \frac{\partial}{\partial x_a} \langle u'_a u'_\alpha \rangle}_{\text{turbulence-turbulence interaction}} - \frac{\partial p'}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

turbulence–mean interaction    turbulence–turbulence interaction

→ **Instability approach** → **Quasi-linear theory (+ non-linearity)**

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial U_\alpha}{\partial x_a}}_{\text{turbulence-mean interaction}} - \frac{\partial p^{(R)}}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

Linear in  $\mathbf{u}'$  and  $p^{(R)}$ , each (Fourier) mode evolves independently

→ **Closure approach** → **Homogeneous turb. theory (+ multiple scale)**

$$\frac{\partial u'_\alpha}{\partial t} + U_a \frac{\partial u'_\alpha}{\partial x_a} = \underbrace{-u'_a \frac{\partial u'_\alpha}{\partial x_a} + \frac{\partial}{\partial x_a} \langle u'_a u'_\alpha \rangle}_{\text{turbulence-turbulence interaction}} - \frac{\partial p^{(S)}}{\partial x_\alpha} + \nu \frac{\partial^2 u'_\alpha}{\partial x_a^2}$$

Homogeneous turbulence, no dependence on large-scale inhomogeneity

Nonlinearity

# Renormalised perturbation expansion theory

**Kraichnan, R. H.** (1959) "The structure of isotropic turbulence at very high Reynolds number," J. Fluid Mech. **5**, 497

Velocity

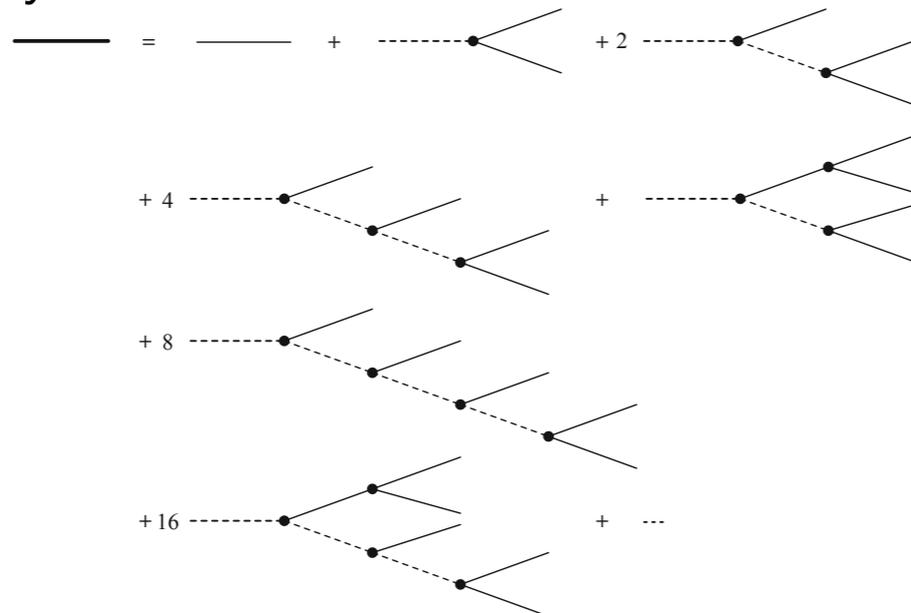
$$u_\alpha(\mathbf{k}; t) = u_\alpha^{(L)}(\mathbf{k}; t) + iM_{cab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p}d\mathbf{q} \\ \times \int_{-\infty}^t dt_1 G_{\alpha c}^{(L)}(\mathbf{k}; t, t_1) u_a(\mathbf{p}; t_1) u_b(\mathbf{q}; t_1)$$

Response function

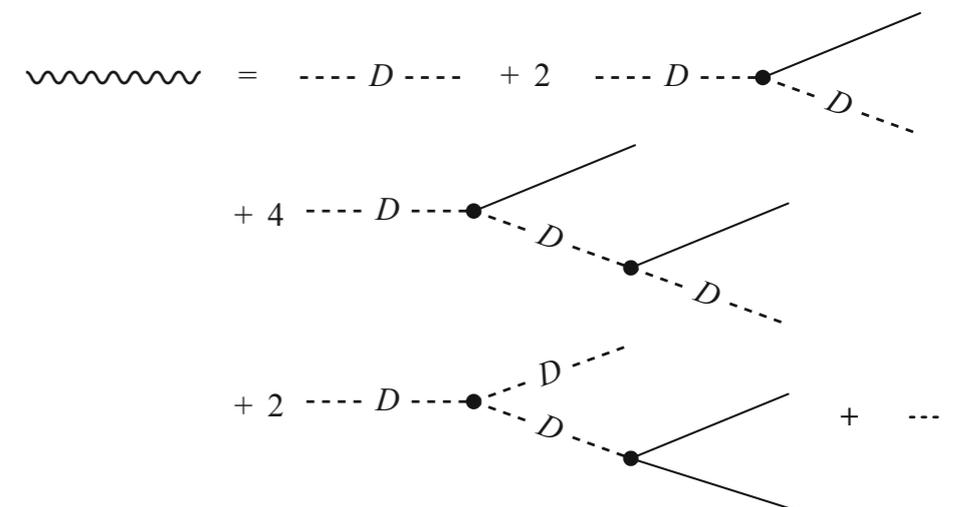
$$G'_{\alpha\beta}(\mathbf{k}; t, t') = G_{\alpha\beta}^{(L)}(\mathbf{k}; t, t') + i2M_{cab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p}d\mathbf{q} \\ \times \int_{t'}^t dt_1 u_a(\mathbf{p}; t_1) G_{\alpha\beta}^{(L)}(\mathbf{q}; t, t_1)$$

Perturbation expansion with respect to the non-linearity from a Gaussian random state at the infinite past

Velocity



Response function





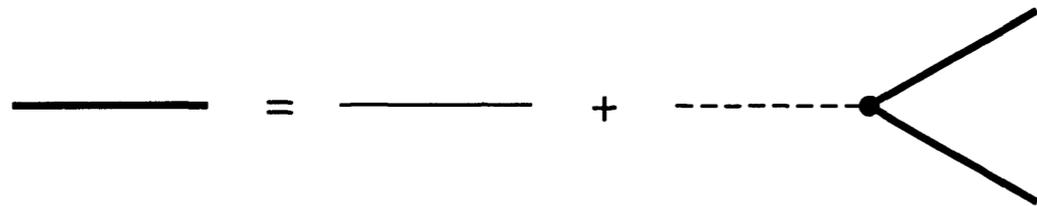
$$u_i(\mathbf{k}; t) = v_i(\mathbf{k}; t) + iM_{ij\ell}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p}d\mathbf{q} \\ \times \int_{-\infty}^t dt_1 g(k; t, t_1) u_j(\mathbf{p}; t_1) u_\ell(\mathbf{q}; t_1)$$

$$u_i(k; t) : \frac{i, k, t}{\text{---}}$$

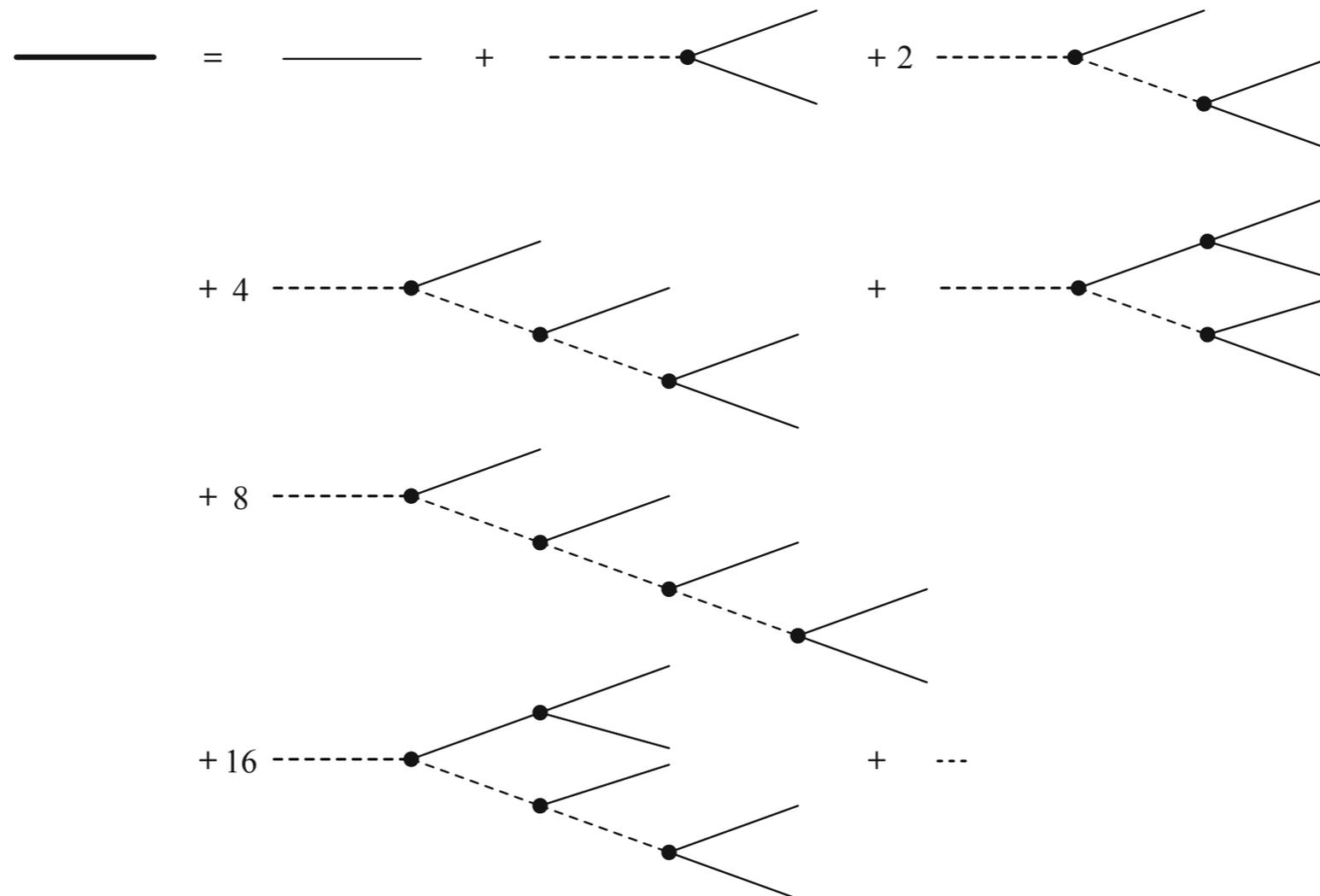
$$v_i(k; t) : \frac{i, k, t}{\text{---}}$$

$$g(k; t, t') : \frac{t \quad k \quad t'}{\text{-----}}$$

$$iM_{ij\ell}(k) : \frac{k}{i} \text{---} \bullet \begin{matrix} p & j \\ & \diagdown \\ & q & \ell \end{matrix}$$



Solve this equation in a perturbative manner with  $\mathbf{v}(\mathbf{k};t)$  as the leading term:



# Response function equation

The response generated by adding an infinitesimal disturbance  $\delta \mathbf{f}(\mathbf{k}; t)$  to

$$\begin{aligned} \frac{\partial}{\partial t} u_i(\mathbf{k}; t) &= -\nu k^2 u_i(\mathbf{k}; t) \\ &\quad + i M_{ij\ell}(\mathbf{k}) \iint u_j(\mathbf{p}; t) u_\ell(\mathbf{q}; t) \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \\ \delta u_i(\mathbf{k}; t) &= \int d\mathbf{k}' \int_{-\infty}^t G'_{ij}(\mathbf{k}, \mathbf{k}'; t, t_1) \delta f_j(\mathbf{k}'; t_1) dt_1 \end{aligned}$$

The **response function equation** is written as

$$\begin{aligned} \frac{\partial}{\partial t} G'_{ij}(\mathbf{k}, \mathbf{k}'; t, t') + \nu k^2 G'_{ij}(\mathbf{k}, \mathbf{k}'; t, t') - D_{ij}(\mathbf{k}') \delta(\mathbf{k} - \mathbf{k}') \delta(t - t') \\ = 2i M_{ilm}(\mathbf{k}) \iint u_\ell(\mathbf{p}; t) G'_{mj}(\mathbf{q}, \mathbf{k}'; t, t') \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \end{aligned}$$

Integrate this with respect to  $\mathbf{k}'$  with putting  $G'_{ij}(\mathbf{k}, \mathbf{k}'; t, t') = G'_{ij}(\mathbf{k}; t, t') \delta(\mathbf{k} - \mathbf{k}')$

$$\begin{aligned} \frac{\partial}{\partial t} G'_{ij}(\mathbf{k}; t, t') + \nu k^2 G'_{ij}(\mathbf{k}; t, t') - D_{ij}(\mathbf{k}) \delta(t - t') \\ = 2i M_{ilm}(\mathbf{k}) \iint u_\ell(\mathbf{p}; t) G'_{mj}(\mathbf{q}; t, t') \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \end{aligned}$$

Formally integrate this **with regarding the r.h.s. as known**

$$G'_{ij}(\mathbf{k}; t, t') = G_{Vij}(\mathbf{k}; t, t') + 2iM_{nlm}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p}d\mathbf{q} \\ \times \int_{t'}^t dt_1 G_{Vin}(\mathbf{k}; t, t_1) u_\ell(\mathbf{p}; t_1) G'_{mj}(\mathbf{q}; t_1, t')$$

$$\text{wavy} = \text{---}D\text{---} + 2 \text{---}D\text{---} \bullet \text{wavy}$$

$$G'_{ij}(k; t, t') : \begin{matrix} t & k & t \\ \text{wavy} & & \text{wavy} \\ i & & j \end{matrix}$$

$$G_{Vij}(k; t, t') : \begin{matrix} t & k & t \\ \text{---}D\text{---} & & \text{---}D\text{---} \\ i & & j \end{matrix}$$

$$G_{ij}(k; t, t') : \begin{matrix} t & k & t' \\ \text{wavy} & & \text{wavy} \\ i & & j \end{matrix}$$

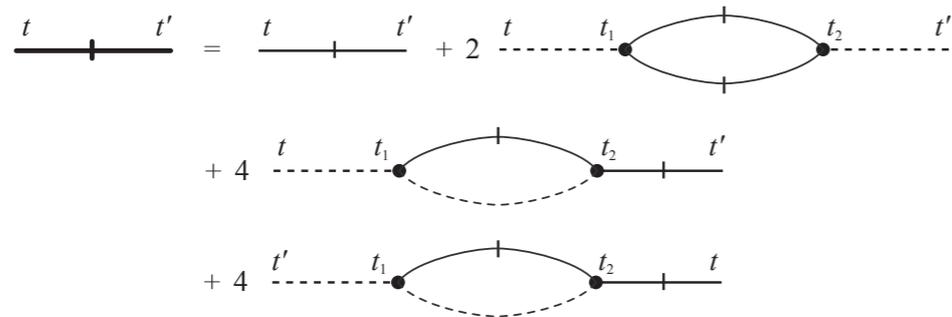
Perturbation expansion up to  $O(M^2)$

$$\text{wavy} = \text{---}D\text{---} + 2 \text{---}D\text{---} \bullet \text{---}D\text{---}$$

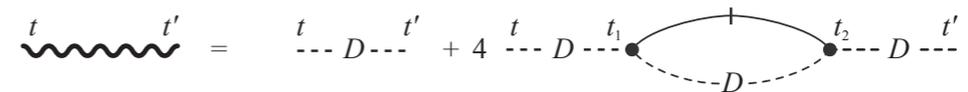
$$+ 4 \text{---}D\text{---} \bullet \text{---}D\text{---} \bullet \text{---}D\text{---}$$

$$+ 2 \text{---}D\text{---} \bullet \text{---}D\text{---} \bullet \text{---}D\text{---}$$

Correlation function  $Q_{\alpha\beta}(\mathbf{k}, \mathbf{k}'; t, t')$



Response function  $G_{\alpha\beta}(\mathbf{k}, \mathbf{k}'; t, t')$



**Renormalisation** a part of the infinite series with respect to the propagators is **summed up to the infinite orders**

$$f_{\text{ex}}(x) = 1 + x + x^2 + x^3 + \dots$$

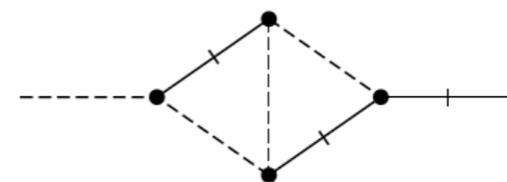
$$f_{\text{ex}}(x) = 1 + x(1 + x + x^2 + \dots)$$

No good

Truncation  $f_{\text{ex}}(x) = 1 + x(1 + x)$

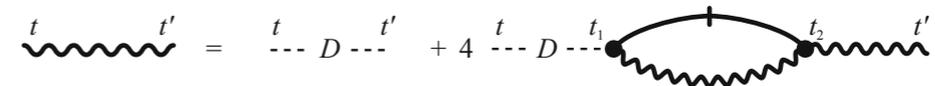
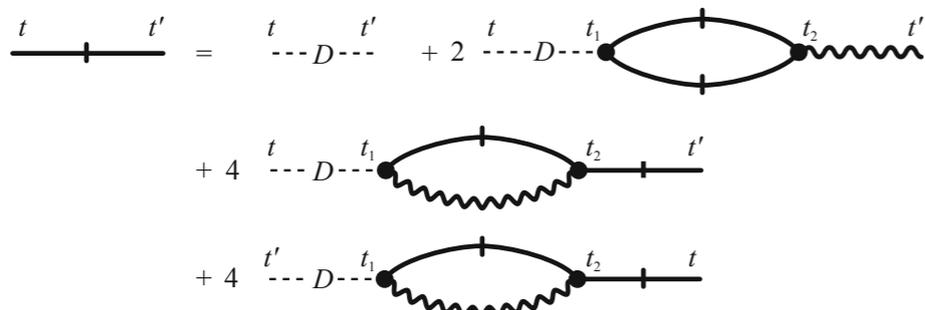
Renormalisation  $f_{\text{ex}}(x) = 1 + x f_{\text{ex}}(x)$

$$f_{\text{ex}}(x) = \frac{1}{1-x}$$



Higher-order vertex not included

**DIA = line (propagator) renormalisation (lowest-order in vertex)**



Inhomogeneity,  
anisotropy, and non-  
equilibrium properties



RHK with Akira Yoshizawa at IIS in 1996

*The difference between the stupidity and genius is that the genius has its limit. (Albert Einstein)*



***“Crazy”***

# Multiple-Scale Direct-Interaction Approximation

Mirror-symmetric case: Yoshizawa, Phys. Fluids **27**, 1377 (1984)

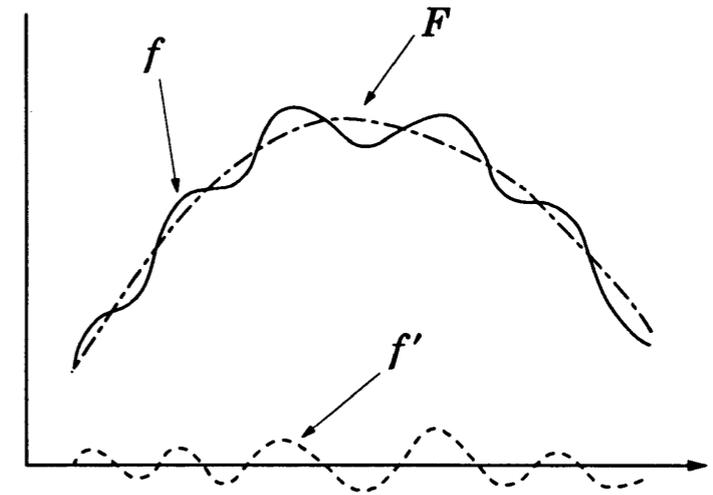
Non-mirror-symmetric case: Yokoi & Yoshizawa, Phys. Fluids A **5**, 464 (1993)

Fast and slow variables  $\xi = \mathbf{x}$ ,  $\mathbf{X} = \delta_x \mathbf{x}$ ;  $\tau = t$ ,  $T = \delta_t t$

Slow variables  $\mathbf{X}$  and  $T$  change only when  $\mathbf{x}$  and  $t$  change much.

$$f = F(\mathbf{X}; T) + f'(\xi, \mathbf{X}; \tau, T)$$

$$\nabla = \nabla_{\xi} + \delta_x \nabla_{\mathbf{X}}; \quad \frac{\partial}{\partial t} = \frac{\partial}{\partial \tau} + \delta_t \frac{\partial}{\partial T}$$



Velocity-fluctuation equation ( $\delta_x = \delta_t$ )

$$\frac{\partial u'_\alpha}{\partial \tau} + U_a \frac{\partial u'_\alpha}{\partial \xi_a} + \frac{\partial}{\partial \xi_a} u'_a u'_\alpha + \frac{\partial p'}{\partial \xi_a} - \nu \nabla_{\xi}^2 u'_\alpha$$

$$= \delta \left( -u'_a \frac{\partial U_\alpha}{\partial X_a} - \frac{D u'_\alpha}{DT} - \frac{\partial p'}{\partial X_\alpha} - \frac{\partial}{\partial X_a} \left( u'_a u'_\alpha - R_{a\alpha} + 2\nu \frac{\partial^2 u'_\alpha}{\partial X_a \partial \xi_a} \right) \right)$$

$$+ \delta^2 (\nu \nabla_X^2 u'_\alpha)$$

$$\frac{\partial u'_a}{\partial \xi_a} + \delta \frac{\partial u'_a}{\partial X_a} = 0$$

**Inhomogeneities, anisotropies,  
non-equilibrium properties**

$$\frac{D}{DT} = \frac{\partial}{\partial T} + \mathbf{U} \cdot \nabla_X$$

# Multiple-Scale DIA calculations

Scale parameter expansion  $f' = f'_0 + \delta f'_1 + \delta^2 f'_2 + \dots = \sum_n \delta^n f'_n$

Basic-field (lowest-order field) equation

$$\frac{\partial u_0^i(\mathbf{k}; \tau)}{\partial \tau} + \nu k^2 u_0^i(\mathbf{k}; \tau) - i M^{ijk}(\mathbf{k}) \iint d\mathbf{p} d\mathbf{q} \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) u_0^j(\mathbf{p}; \tau) u_0^k(\mathbf{q}; \tau) = 0$$

Projection operators

$$M^{ijk}(\mathbf{k}) = \frac{1}{2} [k^j D^{ik}(\mathbf{k}) + k^k D^{ij}(\mathbf{k})], \quad D^{ij}(\mathbf{k}) = \delta^{ij} - \frac{k^i k^j}{k^2}$$

Green's function

$$\begin{aligned} \frac{\partial G'_{\alpha\beta}(\mathbf{k}; \tau, \tau')}{\partial \tau} + \nu k^2 G'_{\alpha\beta}(\mathbf{k}; \tau, \tau') \\ - 2i M^{\alpha ab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} u'_{0a}(\mathbf{p}; \tau) G'_{b\beta}(\mathbf{q}; \tau, \tau') \\ = D_{\alpha\beta}(\mathbf{k}) \delta(\tau - \tau') \end{aligned}$$

## 1st-order field

$$\begin{aligned}
 & \frac{\partial u'_{1\alpha}(\mathbf{k}; \tau)}{\partial \tau} + \nu k^2 u'_{1\alpha}(\mathbf{k}; \tau) \\
 & - 2i M_{\alpha ab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} u'_{0a}(\mathbf{p}; \tau) u'_{S1b}(\mathbf{q}; \tau) \\
 = & -D_{\alpha b}(\mathbf{k}) u'_{0a}(\mathbf{k}; \tau) \frac{\partial U_b}{\partial X_a} - D_{\alpha a}(\mathbf{k}) \frac{D u'_{0a}(\mathbf{k}; \tau)}{DT_1} \\
 & + 2M_{\alpha ab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \frac{q_b}{q^2} u'_{0a}(\mathbf{p}; \tau) \frac{\partial u'_{0c}(\mathbf{q}; \tau)}{\partial X_{Ic}} \\
 & - D_{\alpha d}(\mathbf{k}) M_{abcd}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \frac{\partial}{\partial X_{Ic}} (u'_{0a}(\mathbf{p}; \tau) u'_{0b}(\mathbf{q}; \tau))
 \end{aligned}$$

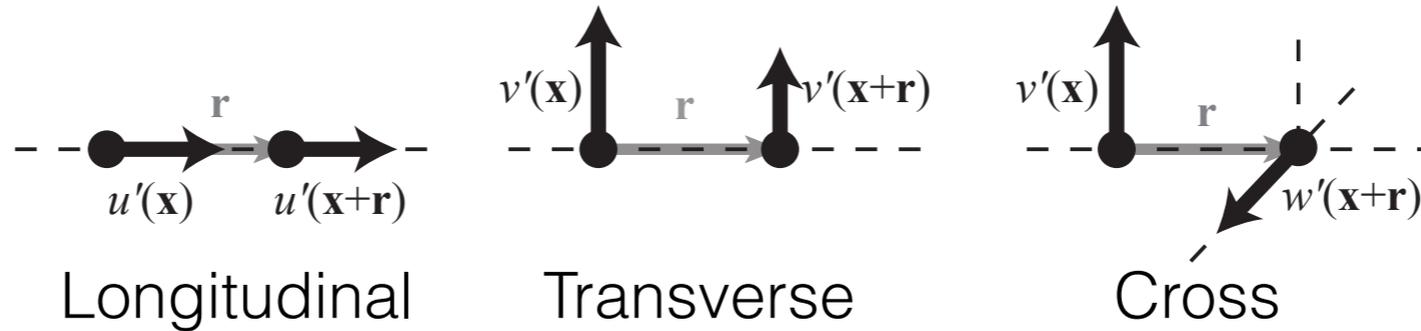
Inhomogeneities, anisotropies,  
and non-equilibrium properties

## Formal solution in terms of Green's function

$$\begin{aligned}
 u'_{1\alpha}(\mathbf{k}; \tau) = & -\frac{\partial U_b}{\partial X_a} \int_{-\infty}^{\tau} d\tau_1 G'_{\alpha b}(\mathbf{k}; \tau, \tau_1) u'_{0a}(\mathbf{k}; \tau_1) \\
 & - \int_{-\infty}^{\tau} d\tau_1 G'_{\alpha a}(\mathbf{k}; \tau, \tau_1) \frac{D u'_{0a}(\mathbf{k}; \tau_1)}{DT_1} \\
 & + 2M_{dab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \int_{-\infty}^{\tau} d\tau_1 G'_{\alpha d}(\mathbf{k}; \tau, \tau_1) \frac{q_b}{q^2} u'_{0a}(\mathbf{p}; \tau_1) \frac{\partial u'_{0c}(\mathbf{q}; \tau_1)}{\partial X_{Ic}} \\
 & - M_{abcd}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \int_{-\infty}^{\tau} d\tau_1 G'_{\alpha d}(\mathbf{k}; \tau, \tau_1) \frac{\partial}{\partial X_{Ic}} (u'_{0a}(\mathbf{p}; \tau_1) u'_{0b}(\mathbf{q}; \tau_1))
 \end{aligned}$$

# Statistical assumptions on the lowest-order (basic) fields

## Basic fields: homogeneous isotropic



Scalar (isotropic) quantity

Should be invariant under rotation of  $\mathbf{r}$ ,  $\mathbf{a}$ ,  $\mathbf{b}$

$$a^i \tilde{R}^{ij}(\mathbf{r}; t, t') b^j = A(\mathbf{r}; t, t') \mathbf{a} \cdot \mathbf{b} + B(\mathbf{r}; t, t') (\mathbf{r} \cdot \mathbf{a})(\mathbf{r} \cdot \mathbf{b}) + C(\mathbf{r}; t, t') \mathbf{r} \cdot (\mathbf{a} \times \mathbf{b})$$

$$\mathbf{a} = \mathbf{e}^i, \quad \mathbf{b} = \mathbf{e}^j \quad \longrightarrow \quad \tilde{R}^{ij}(\mathbf{r}; t, t') = A(\mathbf{r}; t, t') \delta^{ij} + B(\mathbf{r}; t, t') r^i r^j + C(\mathbf{r}; t, t') \epsilon^{ij\ell} r^\ell$$

$$\text{Equivalently,} \quad \tilde{R}^{ij}(\mathbf{r}; t, t') = f \frac{r^i r^j}{r^2} + g \left( \delta^{ij} - \frac{r^i r^j}{r^2} \right) + h \epsilon^{ij\ell} \frac{r^\ell}{r}$$

$$\frac{\langle \vartheta_B'^\alpha(\mathbf{k}; \tau) \chi_B'^\beta(\mathbf{k}'; \tau') \rangle}{\delta(\mathbf{k} + \mathbf{k}')}$$

$$= \underbrace{\Pi^{\alpha\beta}(\mathbf{k}) Q_{\vartheta\chi C}(k; \tau, \tau')}_{\text{Longitudinal}} + \underbrace{D^{\alpha\beta}(\mathbf{k}) Q_{\vartheta\chi S}(k; \tau, \tau')}_{\text{Transverse}} + \underbrace{\frac{i}{2} \frac{k^c}{k^2} \epsilon^{\alpha\beta c} H_{\vartheta\chi}(k; \tau, \tau')}_{\text{Cross}}$$

with solenoidal and dilatational projection operators

$$D^{\alpha\beta}(\mathbf{k}) = \delta^{\alpha\beta} - \frac{k^\alpha k^\beta}{k^2}, \quad \Pi^{\alpha\beta}(\mathbf{k}) = \frac{k^\alpha k^\beta}{k^2}$$

## Calculation of turbulent correlation by DIA

$$\begin{aligned}\langle f'(\mathbf{x}; t)g'(\mathbf{x}; \mathbf{t}) \rangle &= \int d\mathbf{k} \langle f'(\mathbf{k}; \tau)g'(\mathbf{k}; \tau) \rangle / \delta(\mathbf{0}) \\ &= \int d\mathbf{k} (\langle f'_0g'_0 \rangle + \langle f'_0g'_1 \rangle + \langle f'_1g'_0 \rangle + \dots) / \delta(\mathbf{0})\end{aligned}$$

# Mean-field equations in compressible MHD

Yokoi, N., J. Plasma Phys. **84**, 735840501 & 775840603 (2018a,b)

Density	$\frac{\partial \bar{\rho}}{\partial t} + \nabla \cdot (\bar{\rho} \mathbf{U}) = -\nabla \cdot \langle \rho' \mathbf{u}' \rangle$	Turb. mass flux
Momentum	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} U^\alpha + \frac{\partial}{\partial x^a} \bar{\rho} U^a U^\alpha \\ = -(\gamma_0 - 1) \frac{\partial}{\partial x^\alpha} \bar{\rho} Q + \frac{\partial}{\partial x^\alpha} \mu S^{a\alpha} + (\mathbf{J} \times \mathbf{B})^\alpha \\ - \frac{\partial}{\partial x^\alpha} \left( \underbrace{\bar{\rho} \langle u'^a u'^\alpha \rangle}_{\text{Reynolds stress}} - \frac{1}{\mu_0} \underbrace{\langle b'^a b'^\alpha \rangle}_{\text{Turb. Maxwell stress}} + U^a \underbrace{\langle \rho' u'^\alpha \rangle}_{\text{Turb. energy flux}} + U^\alpha \underbrace{\langle \rho' u'^a \rangle}_{\text{Turb. mass-energy correl.}} \right) + R_U^\alpha \end{aligned}$	
Internal energy	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} Q + \nabla \cdot (\bar{\rho} \mathbf{U} Q) = \nabla \cdot \left( \frac{\kappa}{C_V} \nabla Q \right) - \nabla \cdot (\bar{\rho} \langle q' \mathbf{u}' \rangle + Q \langle \rho' \mathbf{u}' \rangle + \mathbf{U} \langle \rho' q' \rangle) \\ - (\gamma_0 - 1) \left( \bar{\rho} Q \nabla \cdot \mathbf{U} + \underbrace{\bar{\rho} \langle q' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. energy dilatation}} + Q \underbrace{\langle \rho' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. mass dilatation}} \right) + R_Q \end{aligned}$	
Magnetic field	$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B} + \langle \mathbf{u}' \times \mathbf{b}' \rangle) + \eta \nabla^2 \mathbf{B}$	Turb. electromotive force

# Some main results of theoretical analysis

## Reynolds and turbulent Maxwell stress

$$\langle \mathbf{u}'\mathbf{u}' - \mathbf{b}'\mathbf{b}' \rangle_D = -\overset{\text{eddy viscosity}}{\nu_K} \mathbf{S} + \overset{\text{cross helicity}}{\nu_M} \mathcal{M} + \overset{\text{inhomogeneous helicity}}{\eta_H} \nabla H \Omega_* + \dots$$

D: deviatoric part

cross helicity

mean velocity strain  $\mathbf{S} = \nabla \mathbf{U} + (\nabla \mathbf{U})^\dagger$  mean magnetic-field strain  $\mathcal{M} = \nabla \mathbf{B} + (\nabla \mathbf{B})^\dagger$

$\Omega_*$ : absolute mean vorticity (mean vorticity + rotation)

## Turbulent electromotive force

$$\begin{aligned} \langle \mathbf{u}' \times \mathbf{b}' \rangle = & \overset{\text{Turb. Mag. Diffusivity}}{-(\beta + \zeta)} \nabla \times \mathbf{B} + \overset{\text{Alpha effect}}{\gamma} \nabla \times \mathbf{U} + \alpha \mathbf{B} + (\nabla \zeta) \times \mathbf{B} \\ & - \chi_\rho \nabla \bar{\rho} \times \mathbf{B} - \chi_Q \nabla Q \times \mathbf{B} - \chi_D \frac{D\mathbf{U}}{Dt} \times \mathbf{B} \quad \text{Compressibility} \end{aligned}$$

Cross-helicity effect      Magnetic pumping

## Turbulent mass flux

$$\langle \rho' \mathbf{u}' \rangle = -\kappa_\rho \nabla \bar{\rho} - \kappa_Q \nabla Q - \kappa_D \frac{D\mathbf{U}}{DT} - \kappa_B \mathbf{B}$$

## Turbulent internal-energy flux

$$\langle q' \mathbf{u}' \rangle = -\eta_Q \nabla Q - \eta_\rho \nabla \bar{\rho} - \eta_B \mathbf{B} \quad + \text{Non-equilibrium effects}$$

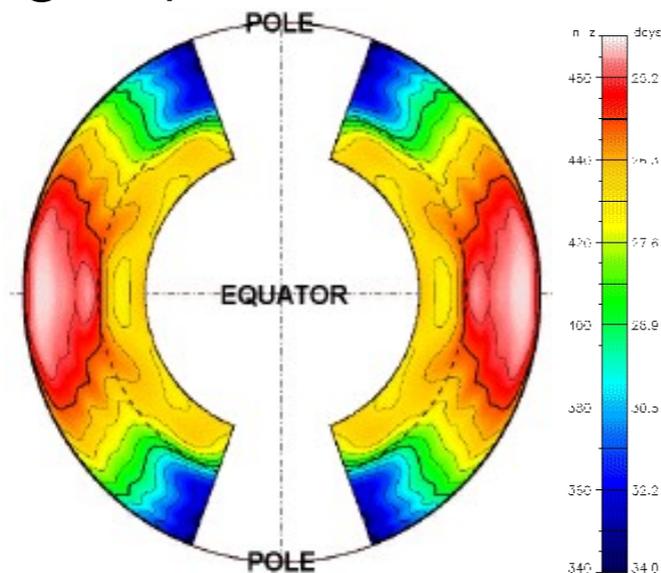
# **Illustrative examples**

# Differential Rotation

Helioseismology shows the internal structure of the Sun.

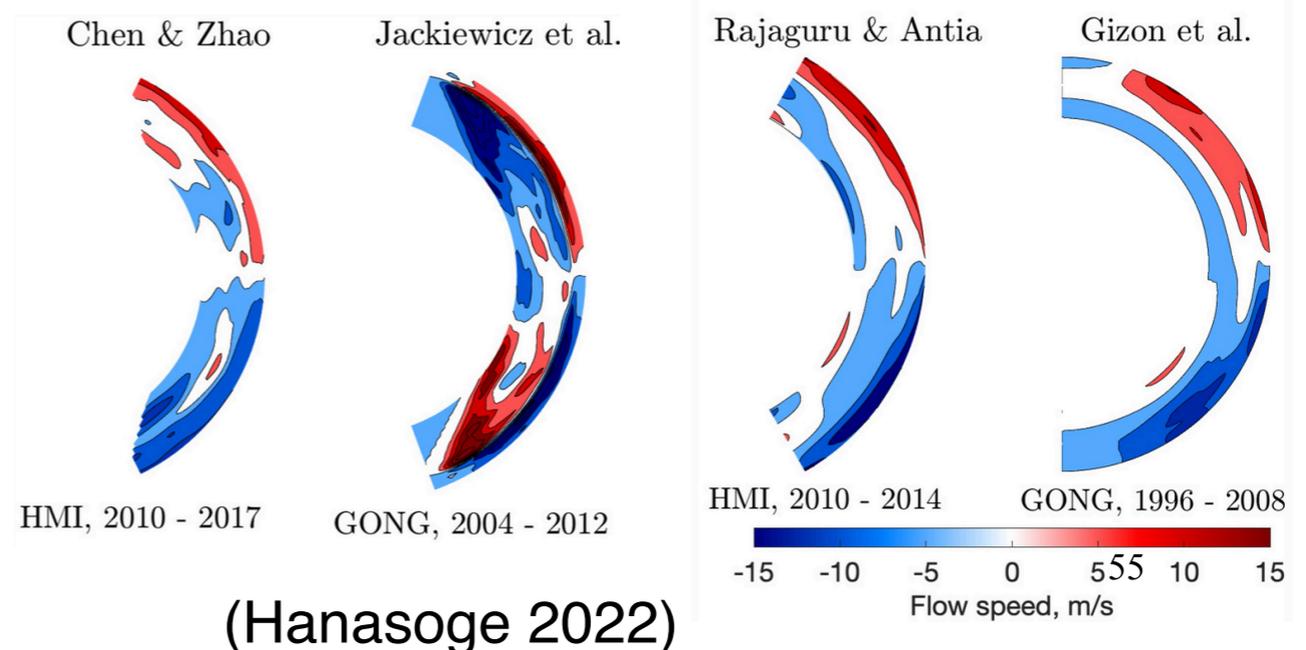
## *Azimuthal velocity*

- Surface **differential rotation** is maintained throughout the convection zone
- Solid body rotation in the radiative interior
- Thin matching zone of shear known as the **tachocline** at the base of the solar convection zone (just in the stable region)



## *Meridional circulation*

- Meridional flow **poleward at surface**
- Interior structure not settled
- **As stars spin faster** they tend to have **slower meridional flows** (from models)

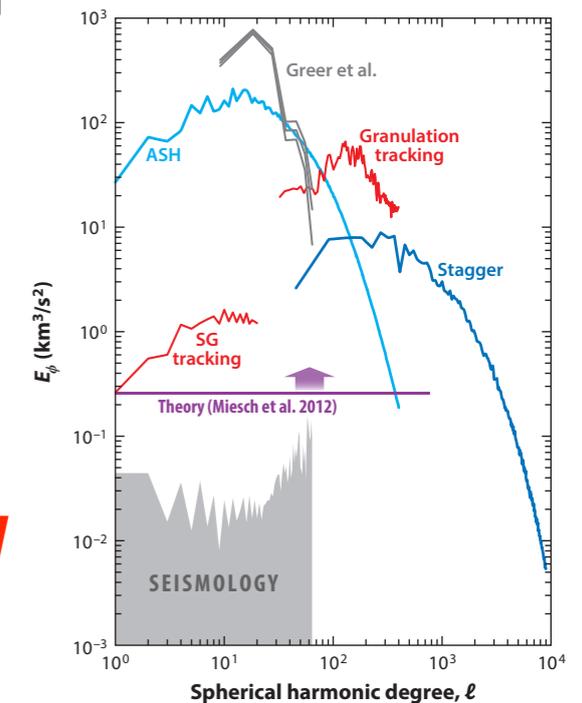


# Convection conundrum

The current numerical simulations do not capture some basic characteristics of the solar convection

(Schumacher & Sreenivasan 2020)

- **The convective velocity amplitude at large horizontal scales observed by helioseismic investigations is much smaller** than the one predicted by global convection simulations
- The differential rotation profile obtained by the global numerical simulations shows **the anti-prograde profile**, if the solar values of **luminosity (energy transfer rate) and rotation rate** are adopted in the simulation.
- Also, if the large-scale convection motions are actually small, **how such weak flows can transfer the solar luminosity and mean differential rotation rate** observed in the helioseismology?



(Hanasoge, Gizon & Sreenivasan 2016)

# Approaches to the conundrum

## Rotation effect

Amplitude and length-scale of convection could be smaller than we think, leading to a smaller  $Ro = U / 2\Omega L$  (Vasil et al 2022, but Käpylä 2023)

-> **Inhomogeneous helicity effect coupled with rotation (Yokoi & Yoshizawa 1993, Yokoi & Brandenburg 2016)**

## Coherent motion effect

Non-local heat transport via entropy rain (e.g., Brandenburg 2016, Anders et al 2019)

-> **Non-equilibrium effect associated with plumes and thermals (Yokoi, Masada & Takiwaki 2022)**

## Magnetic-field effect

Maxwell Stresses from small-scale dynamo acts to counteract Reynolds Stresses that lead to anti-solar rotation profile (e.g., Hotta et al 2022)

-> **Cross-helicity effect in angular momentum transport (Yokoi 2023)**

# Angular-momentum transport by inhomogeneous kinetic helicity

**Yokoi, N. & Yoshizawa, A.** “Statistical analysis of the effects of helicity in inhomogeneous turbulence,” **Phys. Fluids A** **5**, 464-477 (1993)

<https://doi.org/10.1063/1.858869>

**Yokoi, N. & Brandenburg, A.** “Global flow generation by inhomogeneous helicity,” **Phys. Rev. E**. **93**, 033125-1-14 (2016)

<https://doi.org/10.1103/PhysRevE.93.033125>

**Pouquet, A. & Yokoi, N.** “Helical fluid and (Hall)-MHD turbulence: a brief review,” **Phil. Trans. Roy. Soc. A** **380**, 20210081-1-18 (2022)

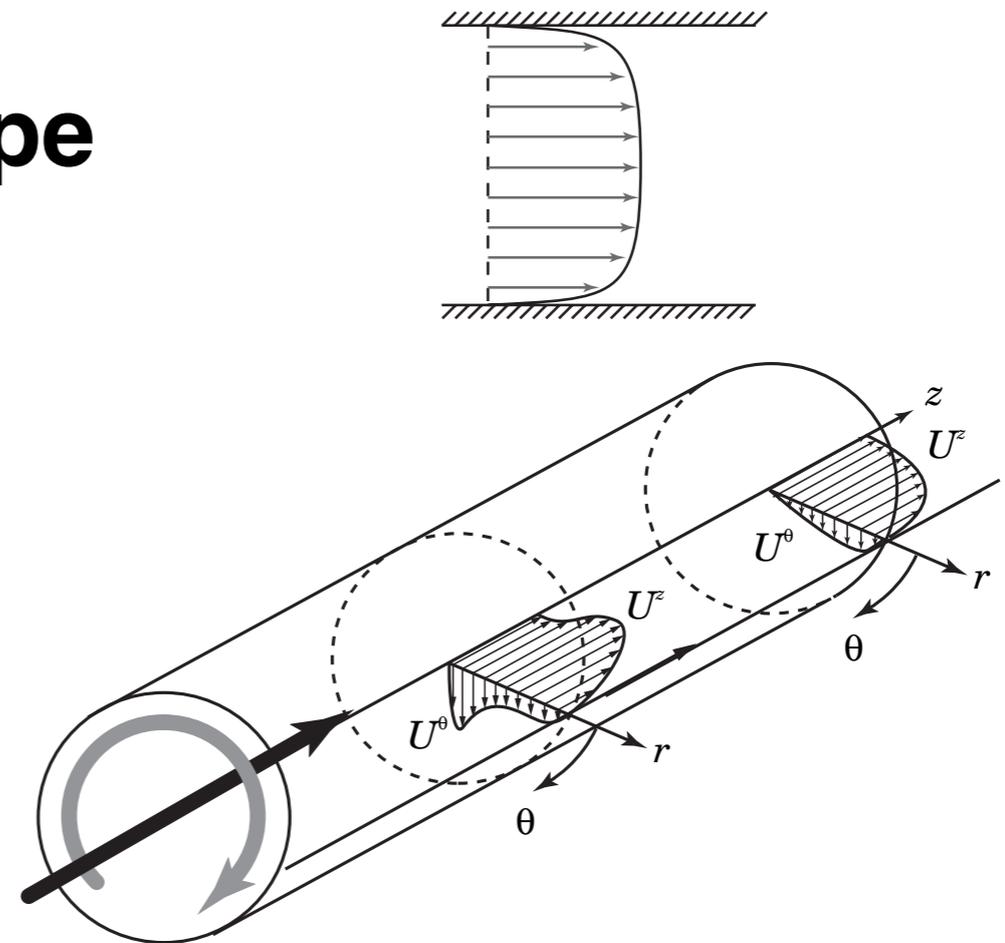
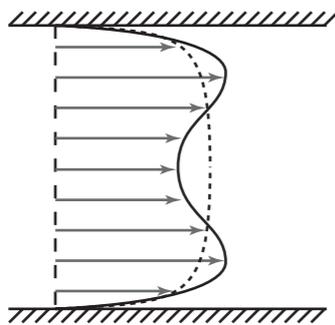
<https://doi.org/10.1098/rsta.2021.0087>

**Yokoi, N.** “Transports in helical fluid turbulence,” pp.25-50,  
in Kuzanyan, Yokoi, Georgoulis & Stepanov (eds.) **AGU Book: Helicities in Geophysics, Astrophysics and Beyond (Wiley, 2023)**

<https://doi.org/10.1002/9781119841715>

# Swirling flow in a circular pipe

## Turbulent swirling pipe flow



## Axially rotating turbulent pipe flow

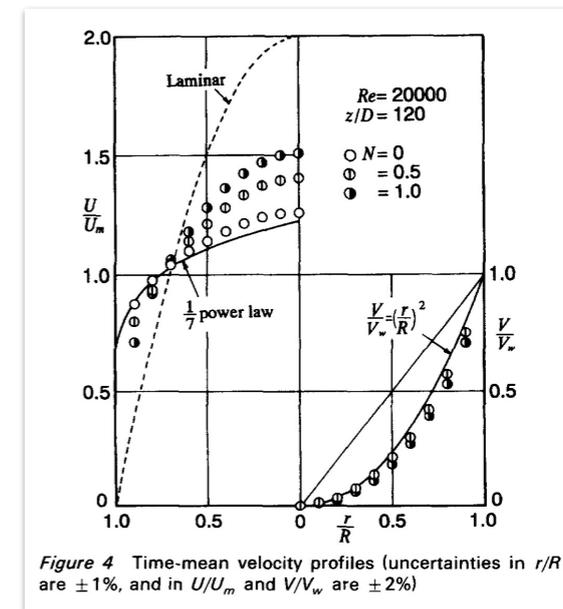
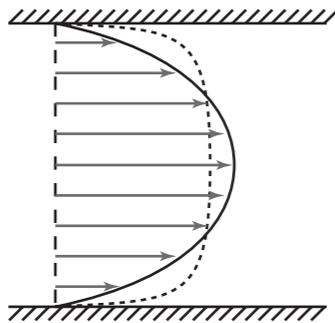


Figure 4 Time-mean velocity profiles (uncertainties in  $r/R$  are  $\pm 1\%$ , and in  $U/U_m$  and  $V/V_w$  are  $\pm 2\%$ )

These flow properties cannot be reproduced by the standard eddy-viscosity representation at all. Too much dissipative.

(Imao et al. 1996)

# Calculation of the Reynolds stress

$$R_{ij} = \langle u'_i(\boldsymbol{\xi}, \mathbf{X}; \tau, T) u'_j(\boldsymbol{\xi}, \mathbf{X}; \tau, T) \rangle = \int R_{ij}(\mathbf{k}, \mathbf{X}; \tau, T) d\mathbf{k}$$

$$\begin{aligned} \langle u'^{\alpha} u'^{\beta} \rangle &= \langle u'_B{}^{\alpha} u'_B{}^{\beta} \rangle + \langle u'_B{}^{\alpha} u'_{01}{}^{\beta} \rangle + \langle u'_{01}{}^{\alpha} u'_B{}^{\beta} \rangle + \dots \\ &+ \langle u'_B{}^{\alpha} u'_{10}{}^{\beta} \rangle + \langle u'_{10}{}^{\alpha} u'_B{}^{\beta} \rangle + \dots \end{aligned}$$

$$\langle u'^{\alpha} u'^{\beta} \rangle_D = -\nu_T \mathcal{S}^{\alpha\beta} + \left[ \Gamma^{\alpha} (\Omega^{\beta} + 2\omega_F^{\beta}) + \Gamma^{\beta} (\Omega^{\alpha} + 2\omega_F^{\alpha}) \right]_D$$

where

$$\mathcal{S}^{\alpha\beta} = \frac{\partial U^{\alpha}}{\partial x^{\beta}} + \frac{\partial U^{\beta}}{\partial x^{\alpha}} - \frac{2}{3} \nabla \cdot \mathbf{U} \delta^{\alpha\beta}$$

Eddy viscosity

$$\nu_T = \frac{7}{15} \int d\mathbf{k} \int_{-\infty}^t d\tau_1 G(k; \tau, \tau_1) Q(k; \tau, \tau_1)$$

mixing length  $\nu_T \sim \tau u^2 \sim ul$

Helicity-related coefficient

$$\boldsymbol{\Gamma} = \frac{1}{30} \int k^{-2} d\mathbf{k} \int_{-\infty}^t d\tau_1 G(k; \tau, \tau_1) \nabla H(k; \tau, \tau_1)$$

helicity inhomogeneity is essential

# Eddy viscosity + Helicity model

Reynolds stress

(Yokoi & Yoshizawa, PoF A5, 464, 1993)

$$\begin{aligned} \mathcal{R}_{\alpha\beta} &\equiv \langle u'_\alpha u'_\beta \rangle \\ &= \frac{2}{3} K \delta_{\alpha\beta} - \nu_T \left( \frac{\partial U_\alpha}{\partial x_\beta} + \frac{\partial U_\beta}{\partial x_\alpha} \right) + \eta \left[ \Omega_\alpha \frac{\partial H}{\partial x_\beta} + \Omega_\beta \frac{\partial H}{\partial x_\alpha} - \frac{2}{3} \delta_{\alpha\beta} (\boldsymbol{\Omega} \cdot \nabla) H \right] \end{aligned}$$

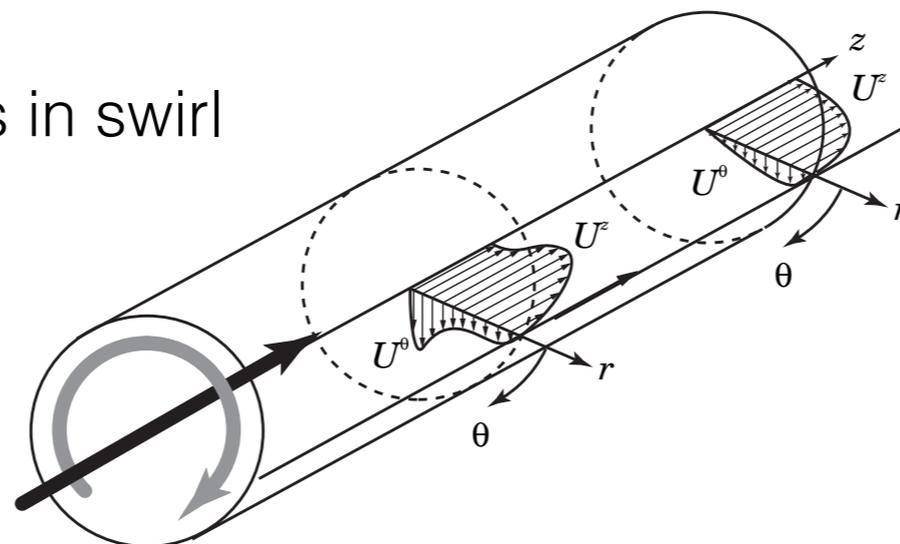
$$\nu_T = C_\nu \tau K, \quad \tau = K/\epsilon, \quad \eta = C_H \tau (K^3/\epsilon^2)$$

Turbulence quantities

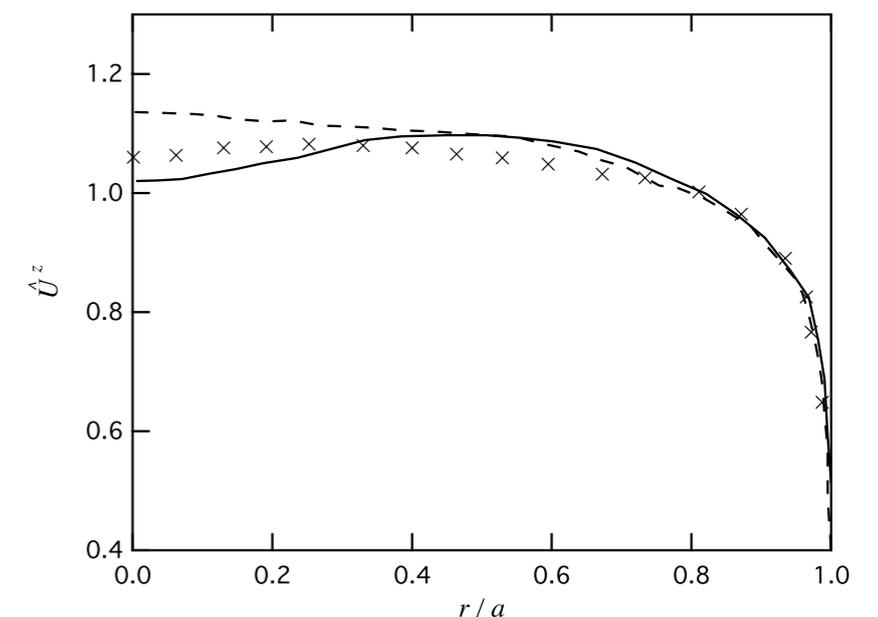
$$K \equiv \frac{1}{2} \langle \mathbf{u}'^2 \rangle, \quad \epsilon \equiv \nu \left\langle \frac{\partial u'_b}{\partial x_a} \frac{\partial u'_b}{\partial x_a} \right\rangle,$$

$$H \equiv \langle \mathbf{u}' \cdot \boldsymbol{\omega}' \rangle, \quad \epsilon_H \equiv 2\nu \left\langle \frac{\partial u'_b}{\partial x_a} \frac{\partial \omega'_b}{\partial x_a} \right\rangle$$

Velocity profiles in swirl

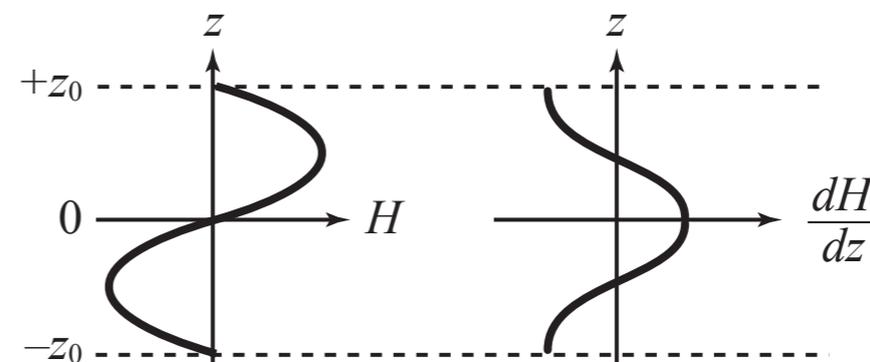
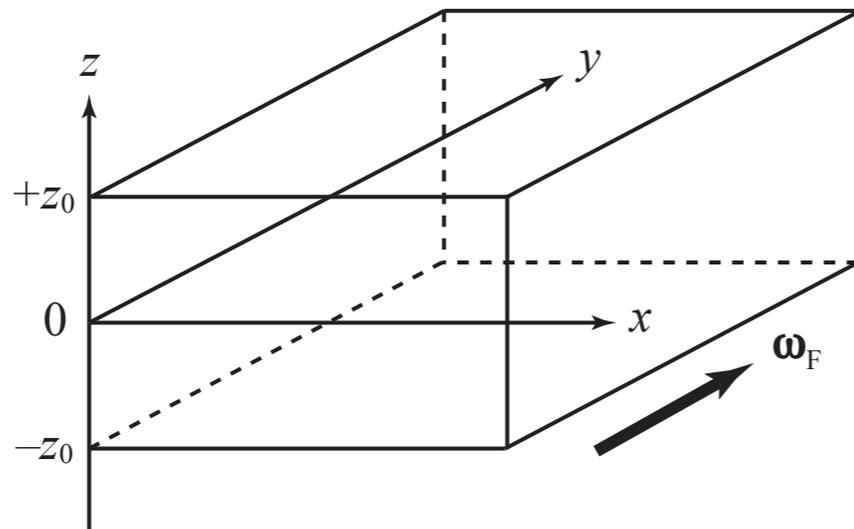


Helicity turbulence model



# DNS set-up

(Yokoi & Brandenburg, PRE, 2016)



Set-up of the turbulence and rotation  $\boldsymbol{\omega}_F$  (left), the schematic spatial profile of the turbulent helicity  $H (= \langle \mathbf{u}' \cdot \boldsymbol{\omega}' \rangle)$  (center) and its derivative  $dH/dz$  (right).

Rotation

$$\boldsymbol{\omega}_F = (\omega_F^x, \omega_F^y, \omega_F^z) = (0, \omega_F, 0)$$

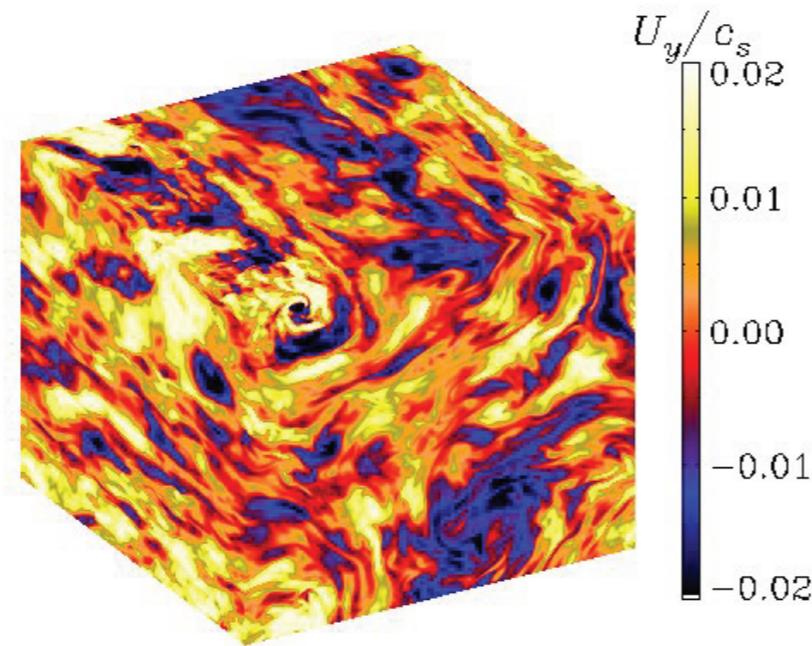
Inhomogeneous  
turbulent helicity

$$H(z) = H_0 \sin(\pi z / z_0)$$

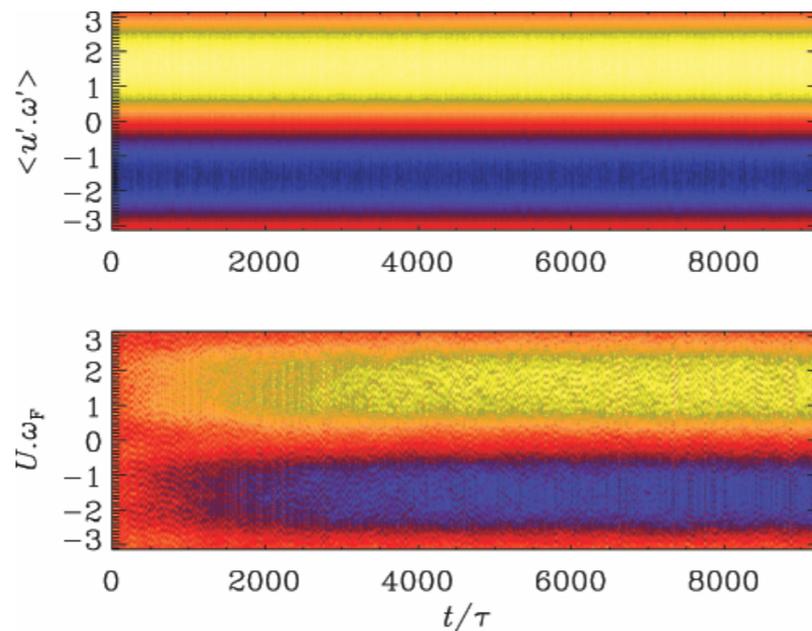
Run	$k_f/k_1$	Re	Co	$\eta/(\nu_T \tau^2)$
A	15	60	0.74	0.22
B1	5	150	2.6	0.27
B2	5	460	1.7	0.27
B3	5	980	1.6	0.51
C1	30	18	0.63	0.50
C2	30	80	0.55	0.03
C3	30	100	0.46	0.08

Summary of DNS results

# Global flow generation



Axial flow component  $U_y$  on the periphery of the domain



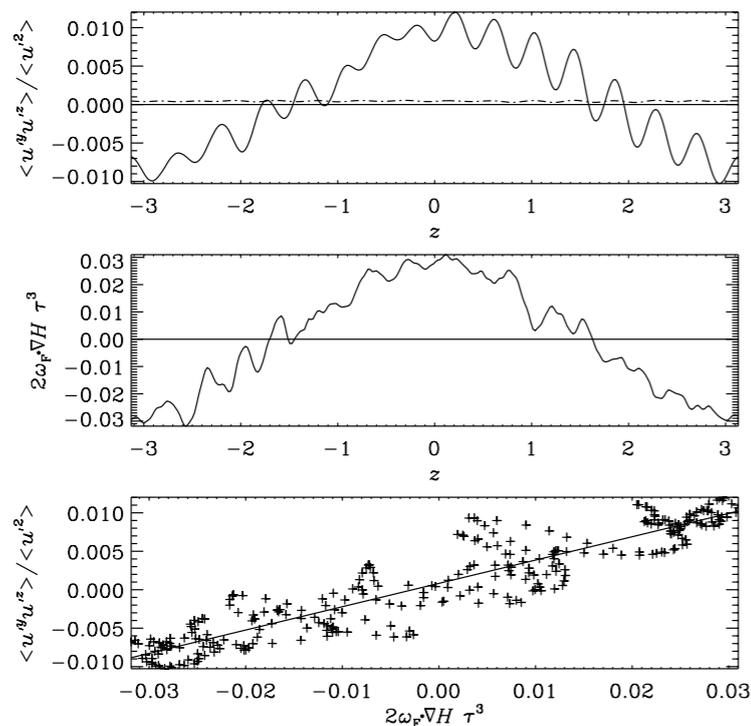
Turbulent helicity  $\langle \mathbf{u}' \cdot \boldsymbol{\omega}' \rangle$  (top) and mean-flow helicity  $\mathbf{U} \cdot 2\boldsymbol{\omega}_F$  (bottom)

# Reynolds stress

$$\langle u'^{\alpha} u'^{\beta} \rangle_D = -\nu_T \mathcal{S}^{\alpha\beta} + \left[ \Gamma^{\alpha} \left( \Omega^{\beta} + 2\omega_F^{\beta} \right) + \Gamma^{\beta} \left( \Omega^{\alpha} + 2\omega_F^{\alpha} \right) \right]_D$$

Early stage

$$\langle u'^y u'^z \rangle = \eta 2\omega_F^y \frac{\partial H}{\partial z}$$

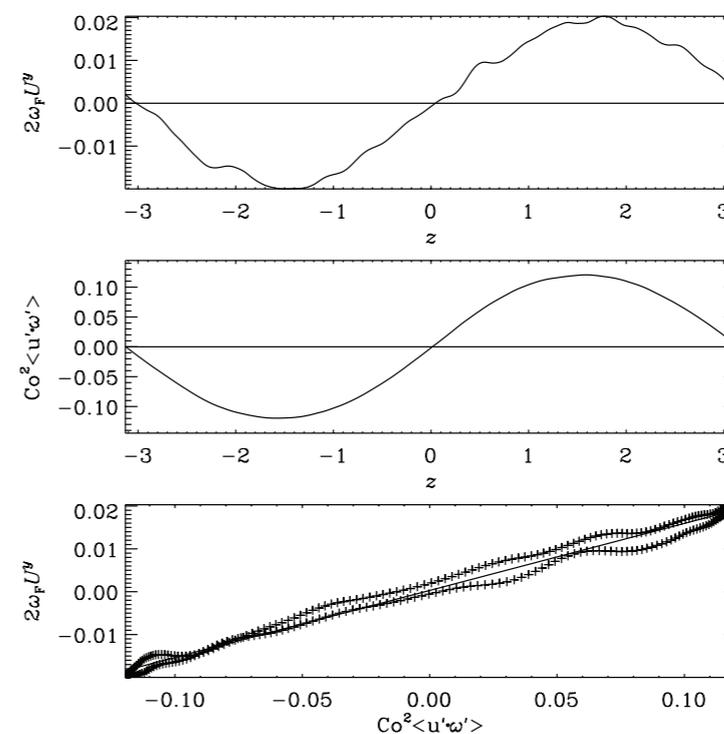


Reynolds stress  $\langle u'^y u'^z \rangle$  (top),  
 helicity-effect term  $(\nabla H)^z 2\omega_F^y$  (middle),  
 and their correlation (bottom).

Developed stage

$$\langle u'^y u'^z \rangle = -\nu_T \frac{\partial U^y}{\partial z} + \eta 2\omega_F^y \frac{\partial H}{\partial z}$$

$$U^y = (\eta/\nu_T) 2\omega_F^y H$$



Mean axial velocity  $U^y$  (top), turbulent  
 helicity multiplied by rotation  $2\omega_F H$   
 (middle), and their correlation (bottom).

# Physical origin

Reynolds stress  $\mathcal{R}^{ij} \equiv \langle u'^i u'^j \rangle$

Vortexmotive force  $\mathbf{V}_M \equiv \langle \mathbf{u}' \times \boldsymbol{\omega}' \rangle$

$$V_M^i = -\frac{\partial \mathcal{R}^{ij}}{\partial x^j} + \frac{\partial K}{\partial x^i}$$

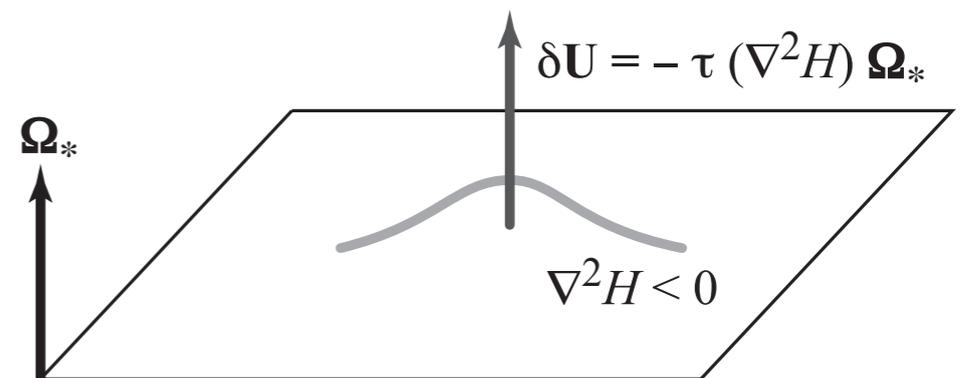
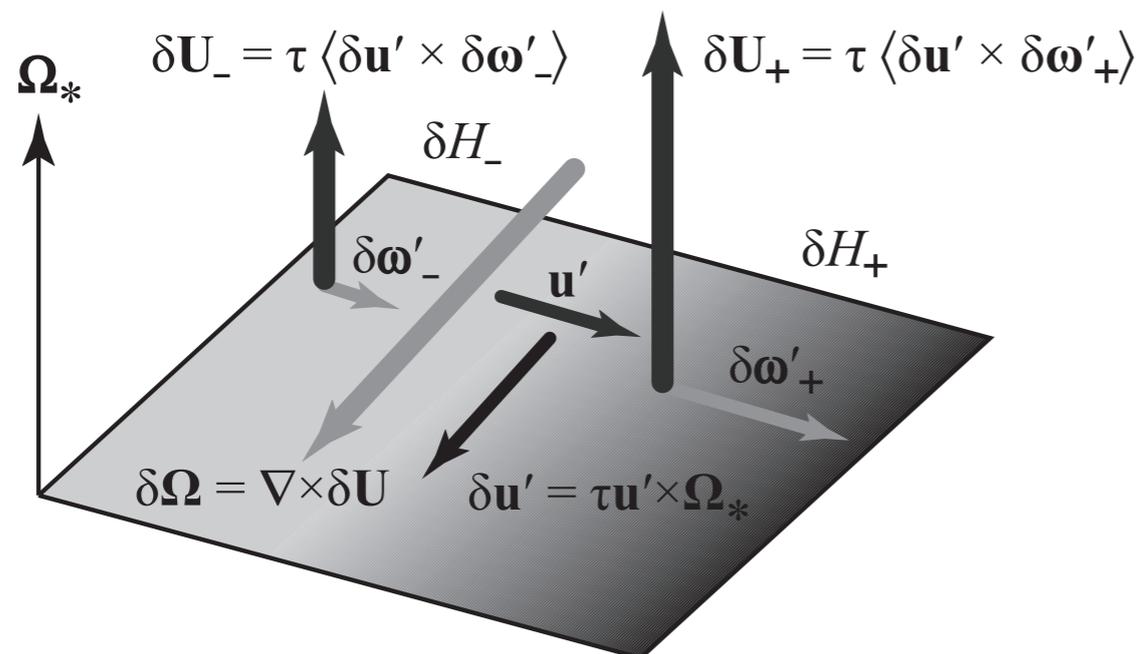
$$\frac{\partial \boldsymbol{\Omega}}{\partial t} = \nabla \times (\mathbf{U} \times \boldsymbol{\Omega}) + \nabla \times \mathbf{V}_M + \nu \nabla^2 \boldsymbol{\Omega}$$

$$\mathbf{V}_M = -D_\Gamma 2\boldsymbol{\omega}_F - \nu_T \nabla \times \boldsymbol{\Omega}$$

$$D_\Gamma = \nabla \cdot \boldsymbol{\Gamma} \propto \nabla^2 H$$

→  $\delta \mathbf{U} \sim -(\nabla^2 H) \boldsymbol{\Omega}_*$

$$\nabla^2 H \simeq -\frac{\delta H}{\ell^2} = -\frac{\langle \mathbf{u}' \cdot \delta \boldsymbol{\omega}' \rangle}{\ell^2}$$



Mean flow induction consuming  
eminently localised turbulent helicity

# Angular-momentum transport in the solar convection zone

Angular momentum around the rotation axis

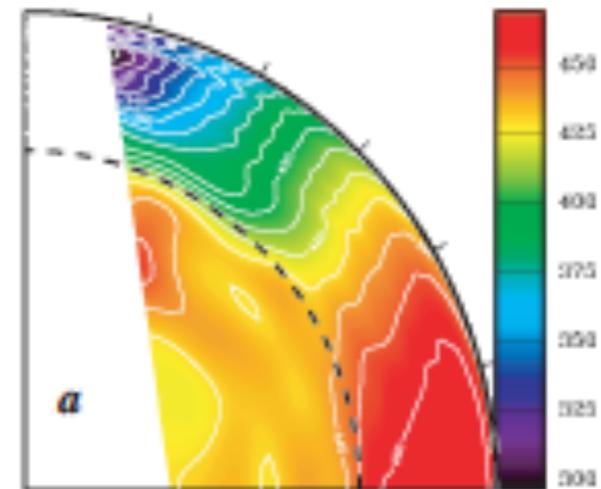
$$L = \Gamma r^2 \omega_F + \Gamma r U^\phi \quad \Gamma = \sin \theta$$

$$\frac{\partial}{\partial t} \rho L + \nabla \cdot (\rho \mathbf{F}_L) = 0$$

Vector flux of angular momentum  $\mathbf{F}_L$

$$F_L^r = L U^r + r \Gamma \mathcal{R}^{r\phi}$$

$$F_L^\theta = L U^\theta + r \Gamma \mathcal{R}^{\theta\phi}$$



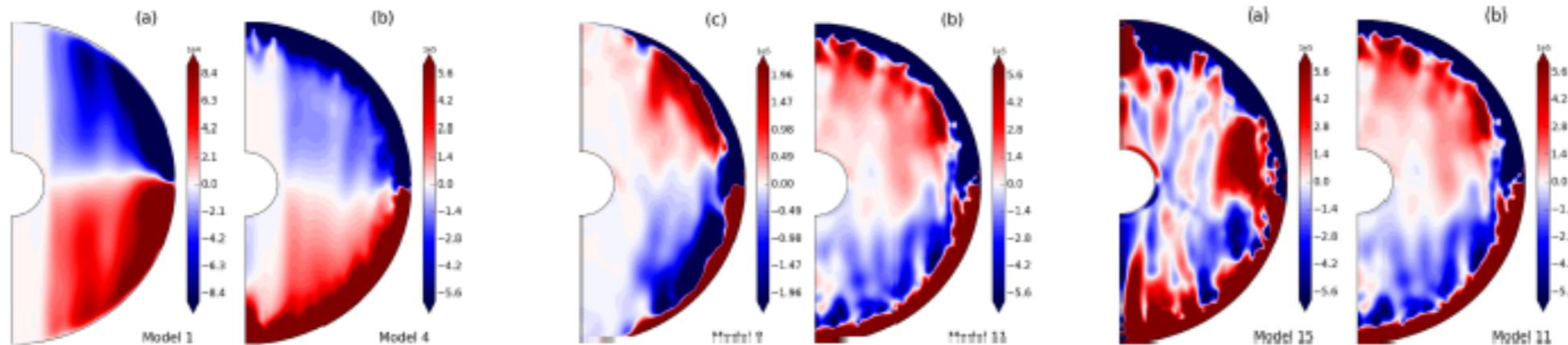
Miesch (2005) Liv. Rev. Sol. Phys. 2005-1

Helicity effect

$$\mathcal{R}_H^{r\phi} = + \frac{\partial H}{\partial r} \left( \frac{1}{r} \frac{\partial}{\partial r} r U^\theta - \frac{1}{r} \frac{\partial U^r}{\partial \theta} \right)$$

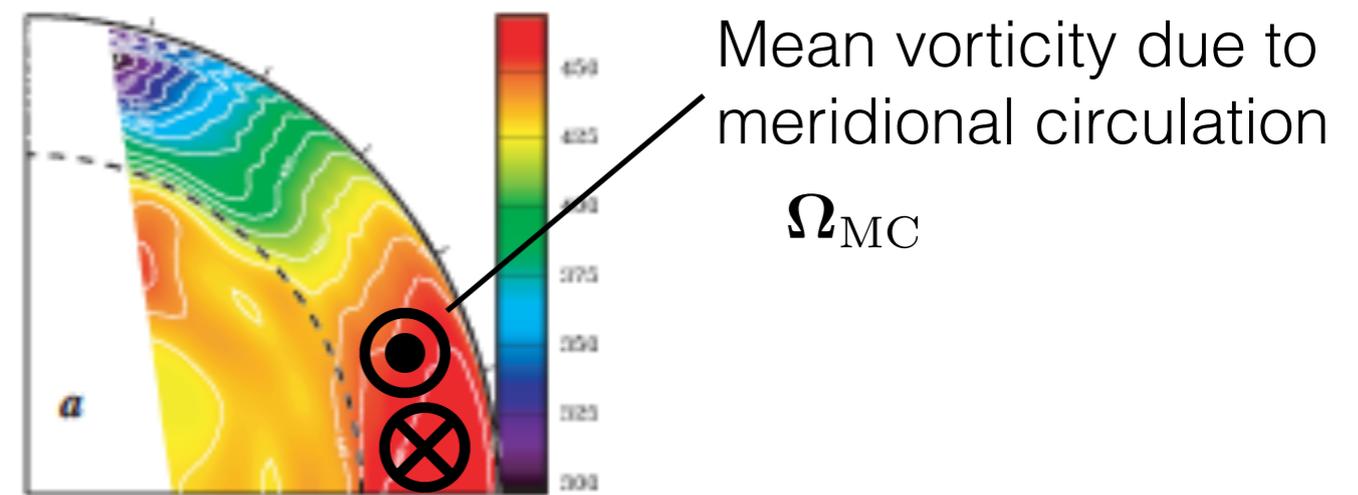
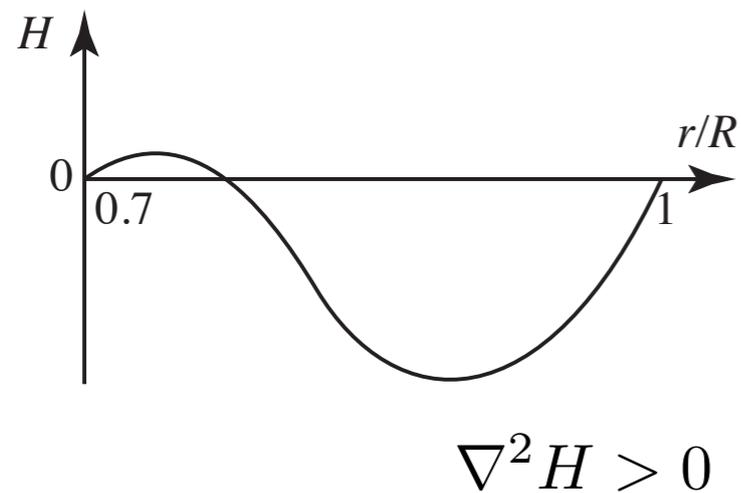
$$\mathcal{R}_H^{\theta\phi} = + \frac{1}{r} \frac{\partial H}{\partial \theta} \left( \frac{1}{r} \frac{\partial}{\partial r} r U^\theta - \frac{1}{r} \frac{\partial U^r}{\partial \theta} \right)$$

# Helicity effect in the stellar convection zone



Duarte, et al, (2016) MNRAS **456**, 1708

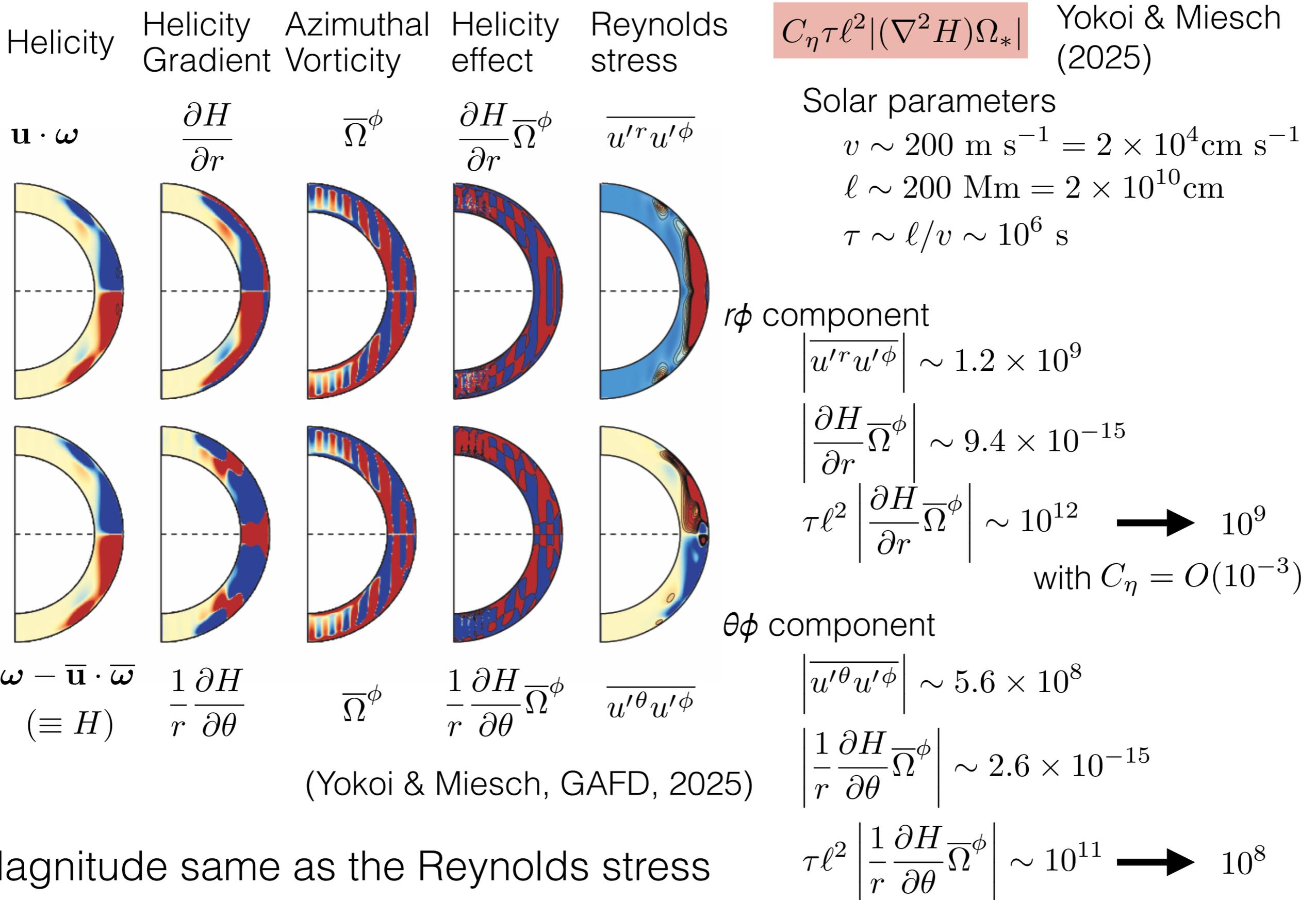
Schematic helicity distribution



$$\mathbf{U} \sim -\tau_U (\nabla^2 H) \boldsymbol{\Omega}_{MC}$$

Prograde mean velocity induction

# Helicity effect in the Reynolds stress



# Validation of the helicity SGS model by DNSs

Yokoi, Mininni, Pouquet, Rosenberg & Marino, Phys. Fluids (to be submitted)

# Problem of Constant Adjustment

Smagorinsky model  $\nu_S = (C_S \Delta)^2 S$

Smagorinsky constant needs to be adjusted such as

Isotropic flow  $C_S \simeq 0.18$

Mixing-layer flow  $C_S \simeq 0.15$

Channel flow  $C_S \simeq 0.1$

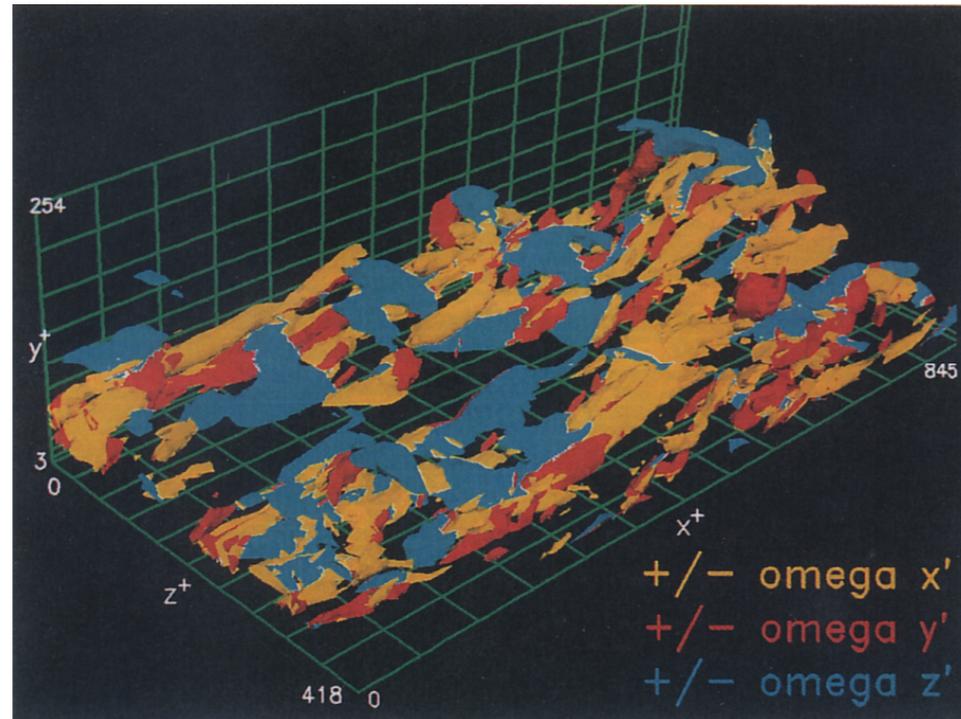
To alleviate

- Dynamic procedure to determine the coefficient  $C_S$
- Alternatives to the generic form

$$\mathcal{T}_{\alpha\beta} = C f_{\alpha\beta}(\nabla \bar{\mathbf{u}}; \Delta) \longrightarrow \mathcal{T}_{\alpha\beta} = C f_{\alpha\beta}(\nabla \bar{\mathbf{u}}; \Delta, \dots)$$

- Evolution equations of the SGS quantities

# Implication to SGS modelling



Fluctuating vorticity  
(Robinson, Kline & Spalart 1988)

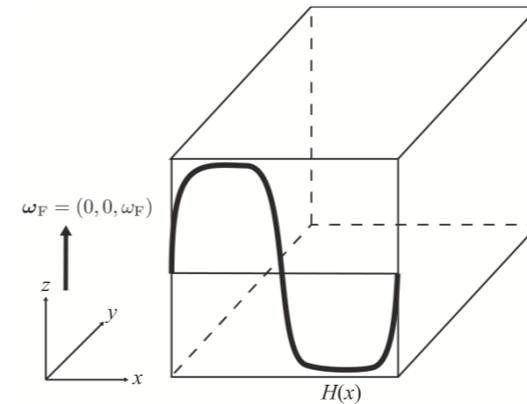
	Smagorinsky constant $C_S$	Dissipative nature	Vortical structures
Isotropic flow	0.18	more	weak
Mixing-layer flow	0.15	↕	intermediate
Channel flow	0.1	less	strong

Coherent vortical structures (streetwise vorticity) may be related to the less dissipative nature

# SGS stress

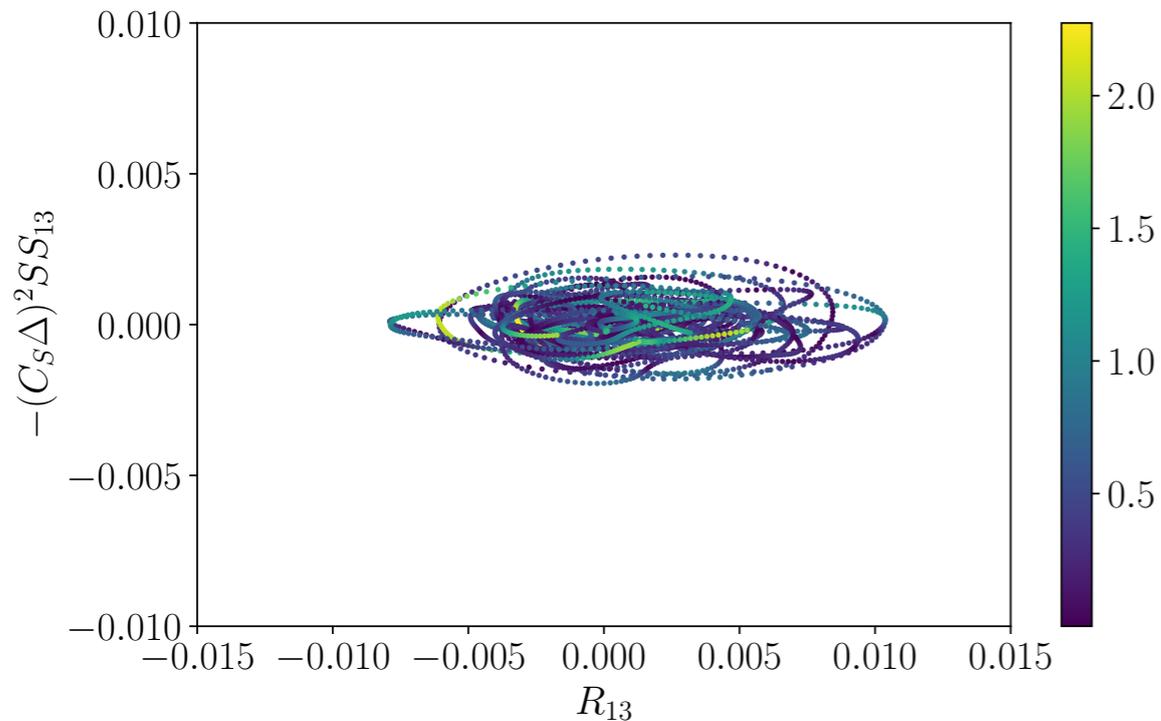
Yokoi, N., Mininni, P., Pouquet, A., et al. Phys. Fluids (2025)

$$\begin{aligned}\tau_{SD} &\equiv \overline{\mathbf{u}\mathbf{u}} - \overline{\mathbf{u}}\overline{\mathbf{u}} \\ &= -\nu_S \overline{\mathbf{s}} + \eta_S \left[ \nabla H_S \overline{\boldsymbol{\omega}} + (\nabla H_S \overline{\boldsymbol{\omega}})^\dagger \right]_D\end{aligned}$$



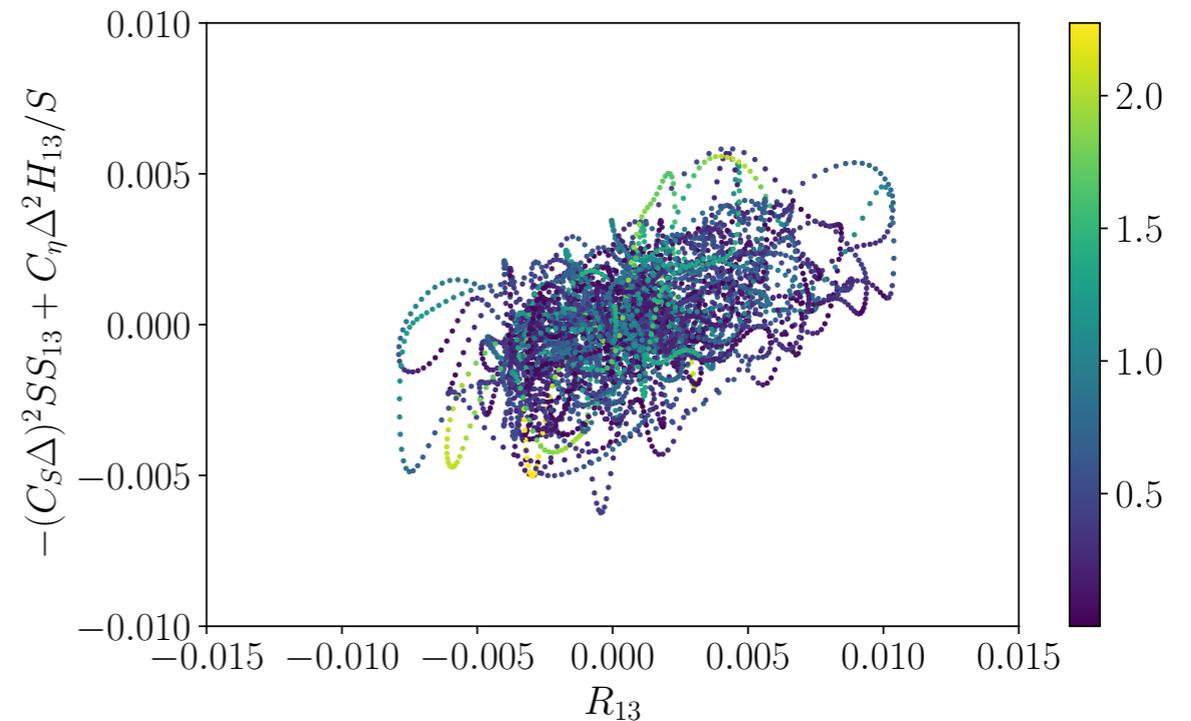
HIT  
with a Gaussian filter  
at  $kc=14$

Smagorinsky model



$$R_{13} = -(C_S \Delta)^2 S S_{13}$$

Helicity SGS model



$$R_{13} = -(C_S \Delta)^2 S S_{13} + C_\eta \Delta^2 H_{13}/S$$

SGS helicity correction improves the SGS stress evaluation

# Non-equilibrium effect in convective transport

**Yokoi, N., Masada, Y. & Takiwaki, T.** “Modelling stellar convective transport with plumes: I. Non-equilibrium turbulence effect in double-averaging formulation,” **Mon. Not. Roy. Astron. Soc. 516, 2718–2735 (2022)**

<https://doi.org/10.1093/mnras/stac1181>

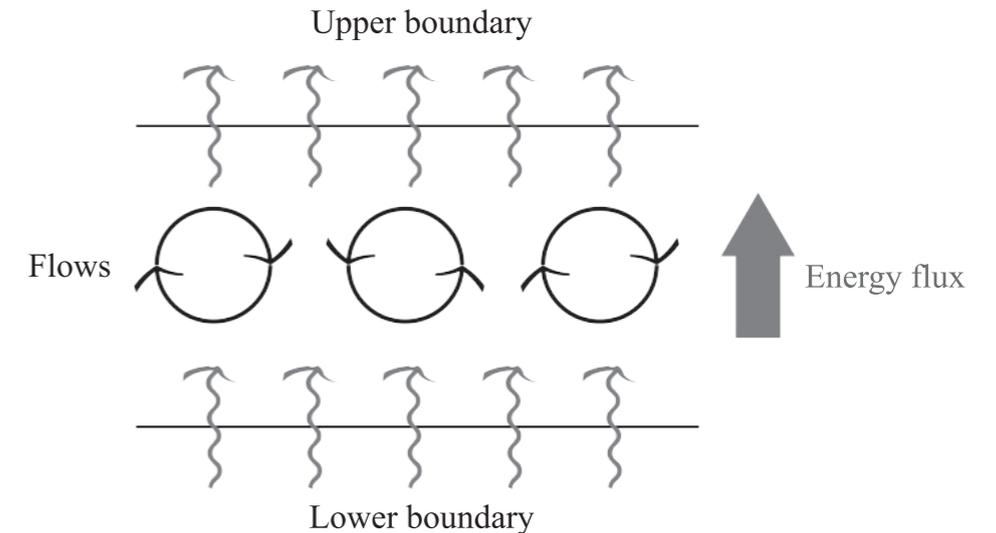
**Yokoi, N.** “Non-Equilibrium Turbulent Transport in Convective Plumes Obtained from Closure Theory,” **Atmosphere 14, 1013-1-22 (2023)**

<https://doi.org/10.3390/atmos14061013>

# Non-equilibrium properties

## Non-equilibrium open system

Non-equilibrium state is sustained by the energy flux through the boundaries with mass and flow



## Deviation from the local equilibrium

Kolmogorov homogeneous isotropic turbulence:

**Local equilibrium between the energy production/injection and dissipation**

**Non-equilibrium property due to the time variation of fluctuations along the large- or meso-scale flows**

$$\frac{\partial}{\partial t} = \frac{\partial}{\partial \tau} + \delta \frac{\partial}{\partial T} \longrightarrow \frac{D\mathbf{u}'}{DT} = \frac{\partial \mathbf{u}'}{\partial T} + (\mathbf{U} \cdot \nabla_{\mathbf{x}})\mathbf{u}'$$

**Irreversibility** Asymmetry with respect to the exchange of time variables

$$H_{ub}(\tau, \tau_1) \neq H_{ub}(\tau_1, \tau)$$

# Plumes in stellar convection

# Plumes and turbulent transport

## Entrainment with plumes

Turner 1973      Importance of the non-equilibrium effects

Linden 2000      Entrainment assumption and scaling

## Surface cooling diving plumes

Spruit 1997      Dominant role of the cooling driven plumes

Rieutord & Zhan 1995      Long-lasting effect of diving plumes

Rast 1998      Beyond hydrostatic pressure and entrainment assumption

## Two layer polytropic gas configuration

Cossette & Rast 2016      More effective mixing due to radiative cooling

## Entropy rain

Brandenburg 2016      Turbulence modelling implementing plume effects

## Entrainment model in the supernovae explosion

Murphy & Meakin 2011      Usefulness of entrainment model

Anett + 2015      Chaotic model with roll of Lorentz

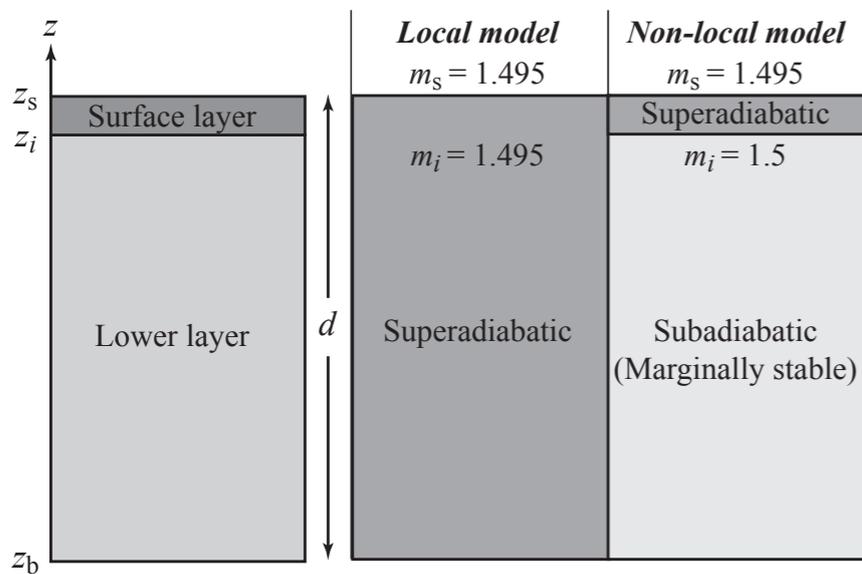
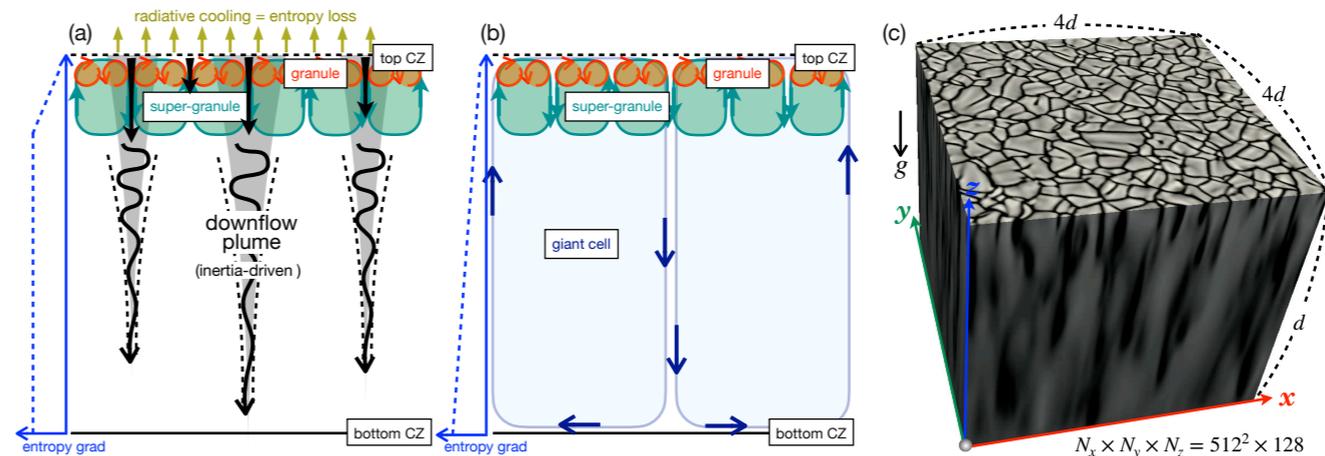
# Setup for Stellar Convection

Two-layer polytropic gas configurations

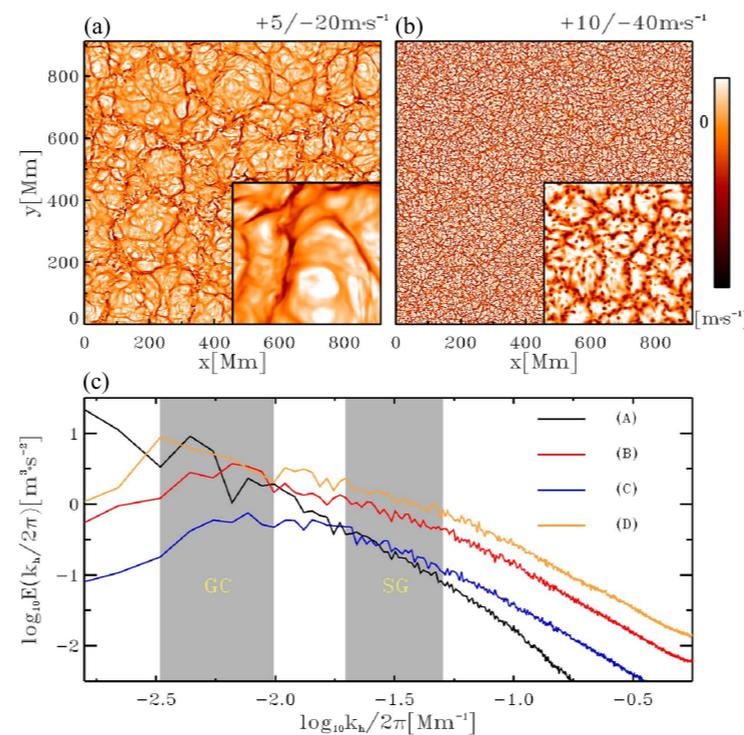
Locally driven case

Surface cooling diving plume case

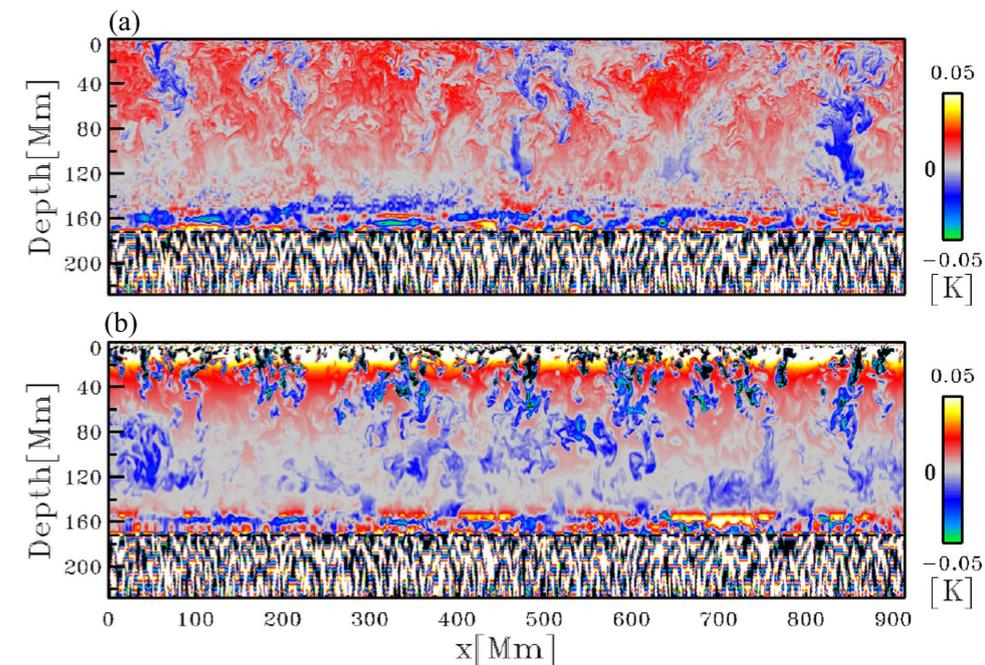
(Cossette & Rast 2016,  
ApJL, 829, L17)



Horizontal cross section of instantaneous vertical velocity and horizontal velocity power spectra

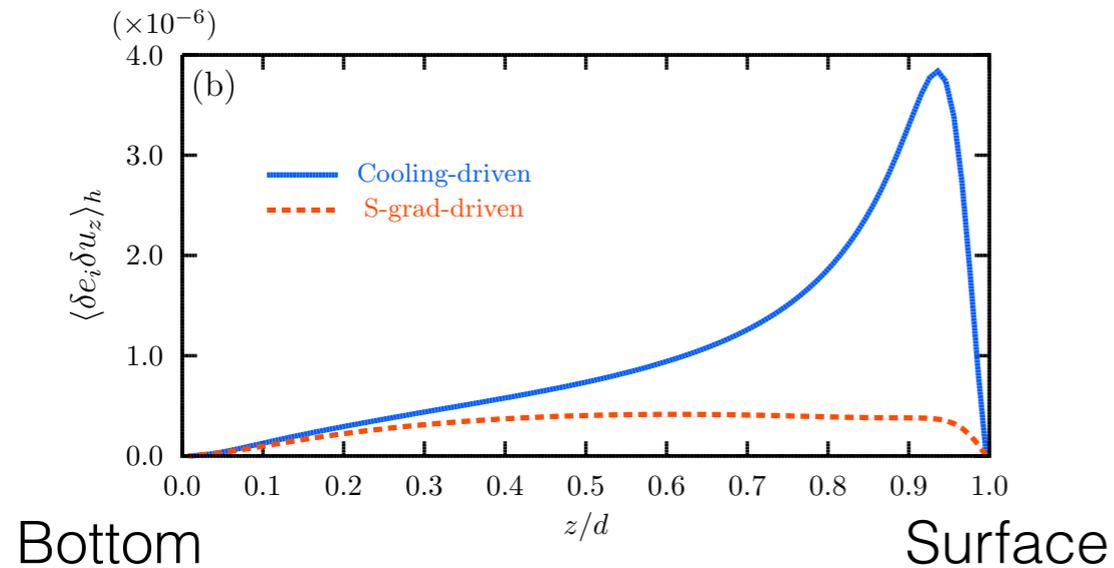


Vertical cross section of instantaneous temperature fluctuations



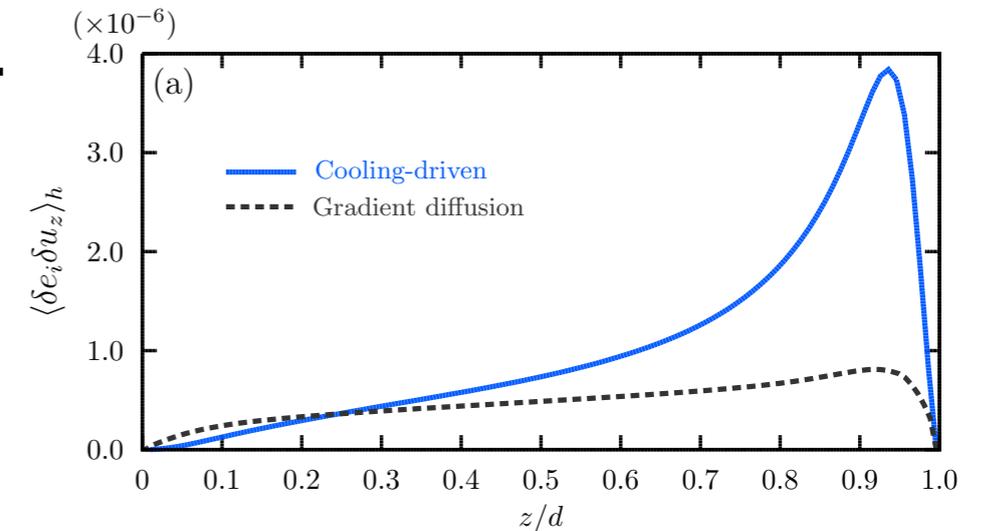
# Spatial distributions of turbulent energy flux $\langle e' u'^z \rangle$

DNS

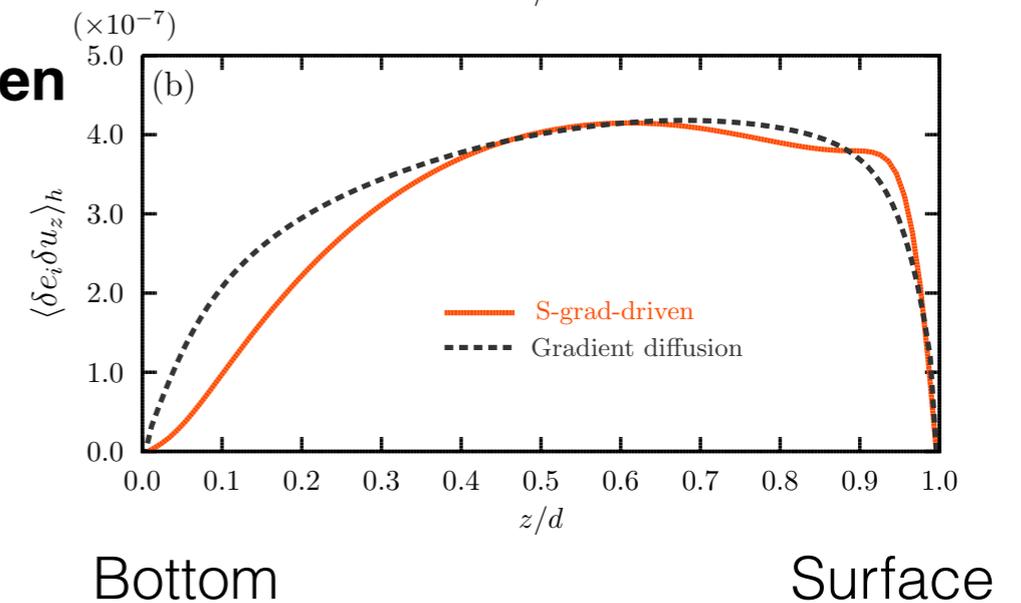


Gradient diffusion model with mixing-length theory (MLT)

Non-locally-driven



Locally-driven



**Turbulent transport of the surface cooling diving plume (non-local transport) cannot be properly described by the gradient-diffusion model with MLT.**

Non-equilibrium effects

## From the Multiple-scale DIA calculations

1st-order field

$$\begin{aligned}
 & \frac{\partial u'_{1\alpha}(\mathbf{k}; \tau)}{\partial \tau} + \nu k^2 u'_{1\alpha}(\mathbf{k}; \tau) \\
 & - 2i M_{\alpha ab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} u'_{0a}(\mathbf{p}; \tau) u'_{S1b}(\mathbf{q}; \tau) \\
 = & - D_{\alpha b}(\mathbf{k}) u'_{0a}(\mathbf{k}; \tau) \frac{\partial U_b}{\partial X_a} - D_{\alpha a}(\mathbf{k}) \frac{D u'_{0a}(\mathbf{k}; \tau)}{D T_I} \\
 & + 2 M_{\alpha ab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \frac{q_b}{q^2} u'_{0a}(\mathbf{p}; \tau) \frac{\partial u'_{0c}(\mathbf{q}; \tau)}{\partial X_{Ic}} \\
 & - D_{\alpha d}(\mathbf{k}) M_{abcd}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \frac{\partial}{\partial X_{Ic}} (u'_{0a}(\mathbf{p}; \tau) u'_{0b}(\mathbf{q}; \tau))
 \end{aligned}$$

Formal solution  
in terms of the  
response  
function G

$$\begin{aligned}
 u_1'^{\alpha}(\mathbf{k}; \tau) = & - \frac{\partial U^b}{\partial X^a} \int_{-\infty}^{\tau} d\tau_1 G'^{\alpha b}(\mathbf{k}; \tau, \tau_1) u_0'^a(\mathbf{k}; \tau_1) \\
 & - \int_{-\infty}^{\tau} d\tau_1 G'^{\alpha a}(\mathbf{k}; \tau, \tau_1) \frac{D u_0'^a(\mathbf{k}; \tau_1)}{D T_I} \\
 & + 2 M^{dab}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \int_{-\infty}^{\tau} d\tau_1 G'^{\alpha d}(\mathbf{k}; \tau, \tau_1) \\
 & \quad \times \frac{q^b}{q^2} u_0'^a(\mathbf{p}; \tau_1) \frac{\partial u_0'^c(\mathbf{q}; \tau_1)}{\partial X_{Ic}} \\
 & - M^{abcd}(\mathbf{k}) \iint \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) d\mathbf{p} d\mathbf{q} \int_{-\infty}^{\tau} d\tau_1 G'^{\alpha d}(\mathbf{k}; \tau, \tau_1) \\
 & \quad \times \frac{\partial}{\partial X_{Ic}} (u_0'^a(\mathbf{p}; \tau_1) u_0'^b(\mathbf{q}; \tau_1))
 \end{aligned}$$

# Non-equilibrium Effects

Turbulent energy  $K \equiv \langle \mathbf{u}'^2 \rangle / 2 = \frac{1}{2} \int d\mathbf{k} \langle u'^l(\mathbf{k}; \tau) u'^l(\mathbf{k}'; \tau) \rangle / \delta(\mathbf{k} + \mathbf{k}')$

$$\langle u'^l u'^l \rangle = \langle u_0'^l u_0'^l \rangle + \delta \langle u_1'^l u_0'^l \rangle + \delta \langle u_0'^l u_1'^l \rangle + \dots$$

Length scale (energy containing scale)  $\ell_C$     Dissipation rate  $\varepsilon = \nu \left\langle \left( \frac{\partial u'^l}{\partial x^l} \right)^2 \right\rangle$

$$K = C_{K1} \varepsilon^{2/3} \ell_C^{2/3} - C_{K2} \varepsilon^{-3/2} \ell_C^{4/3} \frac{D\varepsilon}{Dt} - C_{K3} \varepsilon^{1/3} \ell_C^{1/3} \frac{D\ell_C}{Dt}$$

**Equilibrium effect**

**Non-equilibrium effect**

Solve this by iterations with respect to  $\ell_C$

$$\ell_C = C_{\ell 1} K^{3/2} \varepsilon^{-1} + C_{\ell 2} K^{3/2} \varepsilon^{-2} \frac{DK}{Dt} - C_{\ell 3} K^{5/2} \varepsilon^{-3} \frac{D\varepsilon}{Dt}$$

$$\ell_C = \ell_E \left( 1 - C'_N \frac{1}{K} \frac{D}{Dt} \frac{K^2}{\varepsilon} \right) \quad \text{Equilibrium length scale} \quad \ell_E = K^{3/2} / \varepsilon$$

$$\nu_T = \begin{cases} \nu_{TE} \left( 1 + C_N \frac{1}{K} \frac{D}{Dt} \frac{K^2}{\varepsilon} \right)^{-1} & \text{for } \frac{D}{Dt} \frac{K^2}{\varepsilon} > 0 \\ \nu_{TE} \left( 1 - C_N \frac{1}{K} \frac{D}{Dt} \frac{K^2}{\varepsilon} \right) & \text{for } \frac{D}{Dt} \frac{K^2}{\varepsilon} < 0 \end{cases}$$

# Relevance of the non-equilibrium effect in homogeneous shear turbulence

(Yoshizawa & Nisizima 1993, Phys. Fluids A5, 3302)

Homogeneous shear flow

$$\mathbf{U} = (U_x, U_y, U_z) = (Sy, 0, 0)$$

$K$ - $\varepsilon$  model

in homogeneous shear turbulence

$$\frac{\partial K}{\partial t} = P_K - \varepsilon$$

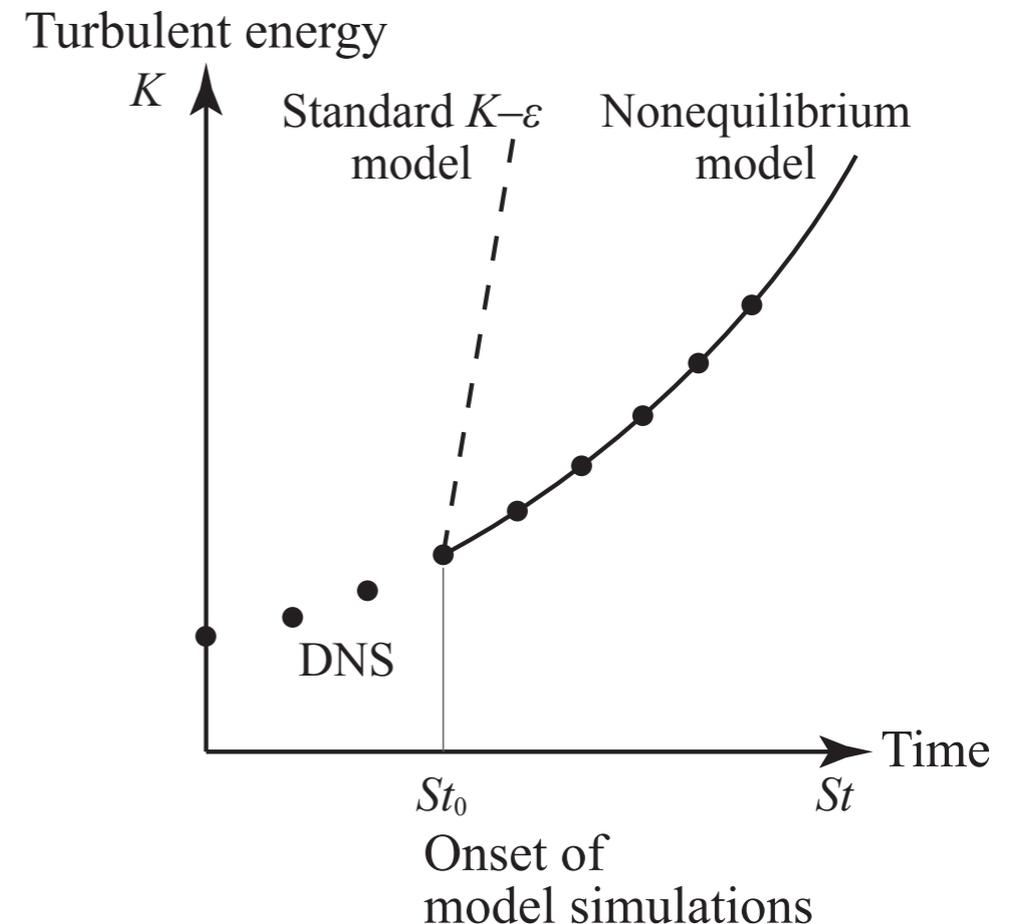
$$\frac{\partial \varepsilon}{\partial t} = C_{\varepsilon 1} \frac{\varepsilon}{K} P_K - C_{\varepsilon 2} \frac{\varepsilon}{K} \varepsilon$$

with  $P_K = -\langle u'_x u'_y \rangle \frac{dU_x}{dy} = +\nu_T S^2$

**Standard  $K$ - $\varepsilon$  model with Equilibrium eddy viscosity**

$$\nu_T = \nu_E = C_\nu \frac{K^2}{\varepsilon}$$

with  $C_\nu = 0.09$



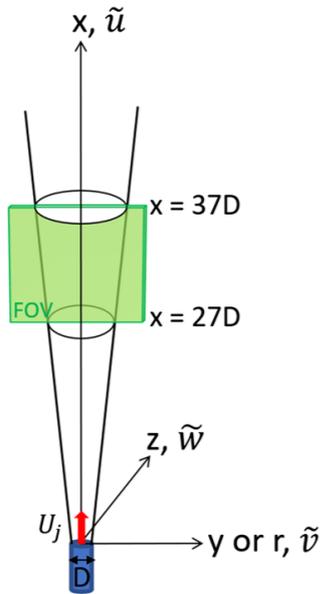
**Non-equilibrium eddy viscosity**

$$\nu_T = \nu_E \left( 1 + C_N \frac{1}{K} \frac{D}{Dt} \frac{K^2}{\varepsilon^2} \right)^{-1}$$

# Non-equilibrium effects in experiment I

## Round jet

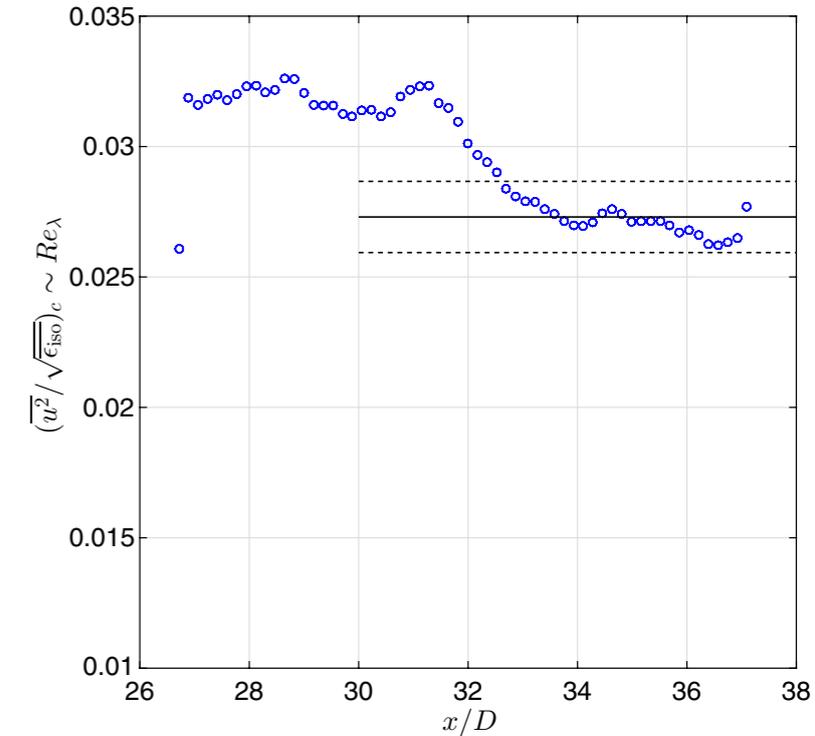
(Lai & Socolofsky 2019, Environ. Fluid Mech. **19**, 349)



Turbulent axial velocity fluctuation normalised by the square root of the dissipation rate at the jet centre

$$\frac{K^2}{\varepsilon} \propto \left[ \frac{\langle (u'^x)^2 \rangle}{\sqrt{\varepsilon_{\text{iso}}}} \right]^2$$

$$\frac{D}{Dt} \frac{K^2}{\varepsilon} < 0 \quad (\text{in turbulent round jets})$$



$x/D$	$\sqrt{\langle (u'^x)^2 \rangle} / U_c$	$\sqrt{\langle (u'^y)^2 \rangle} / U_c$	$\sqrt{\langle (u'^z)^2 \rangle} / U_c$	$K / U_c^2$	$\langle (u'^x)^2 \rangle / \sqrt{\varepsilon_{\text{iso}}}$
31	0.26	0.20	0.19	0.1437	$3.25 \times 10^{-2}$
37	0.26	0.20	0.19	0.1437	$2.65 \times 10^{-2}$

(-20%)

$$\Lambda \simeq C_N \frac{1}{K} (\mathbf{U} \cdot \nabla) \frac{K^2}{\varepsilon} \simeq C_N \left( \frac{U_c}{\sqrt{\langle (u'^x)^2 \rangle}} \right)^2 \frac{1}{U_c} \frac{\Delta \{ [\langle (u'^x)^2 \rangle / \sqrt{\langle \varepsilon_{\text{iso}} \rangle}]^2 \}}{D \Delta(x/D)} \simeq -1.7$$

$$\nu_{\text{NE}} = \nu_{\text{E}} \left( 1 - C_N \frac{1}{K} \frac{D}{Dt} \frac{K^2}{\varepsilon} \right) = \nu_{\text{E}} (1 - \Lambda) > \nu_{\text{E}} \quad \text{Eddy-viscosity-related constant} \quad C_\mu = 0.09 \rightarrow 0.135$$

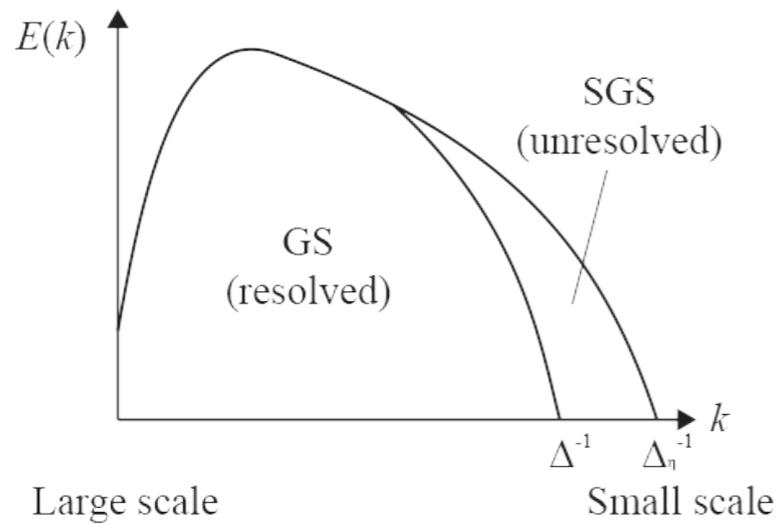


Transport enhancement

# Double averaging approach

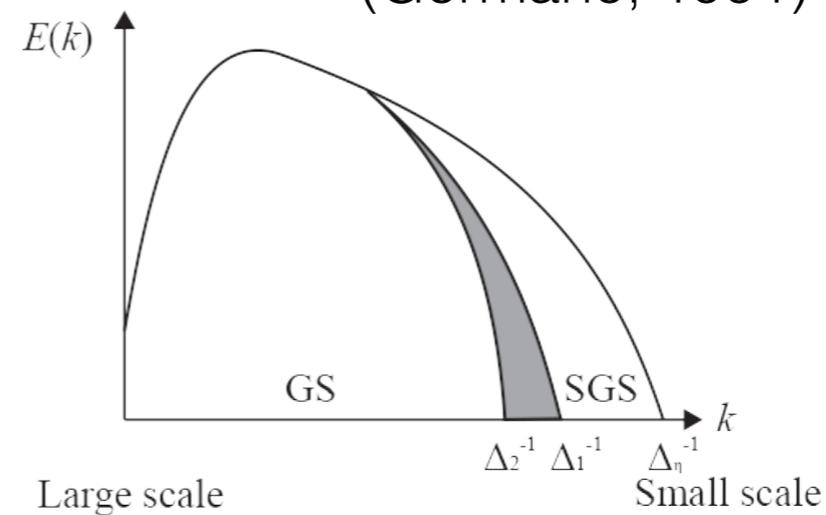
# Double-averaging procedure

## Subgrid-scale (SGS) modelling



## Double-filtering approach

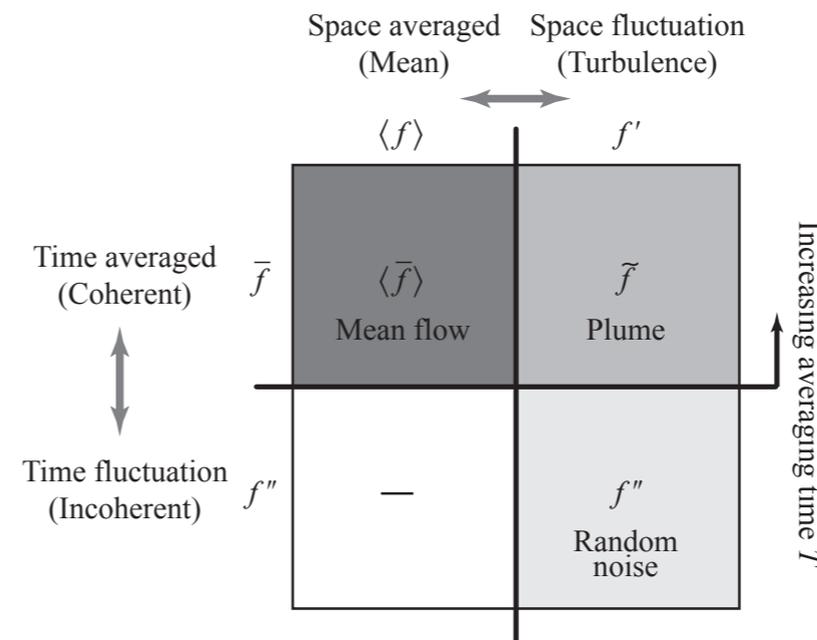
e.g., Dynamic Smagorinsky model (Germano, 1991)



## Time-Space double filtering

## Space-domain filtering

### Time-domain filtering



# Time-space double averaging Space-domain filtering

## Filtering both in space and time domains

### Time average

$$\bar{f}(\mathbf{x}; t) = \int f(\mathbf{x}; s) G(t - s) ds$$

Time filter

$$G(t - s) = \begin{cases} 1/T & (|t - s| \leq T/2) \\ 0 & (\text{otherwise}) \end{cases}$$

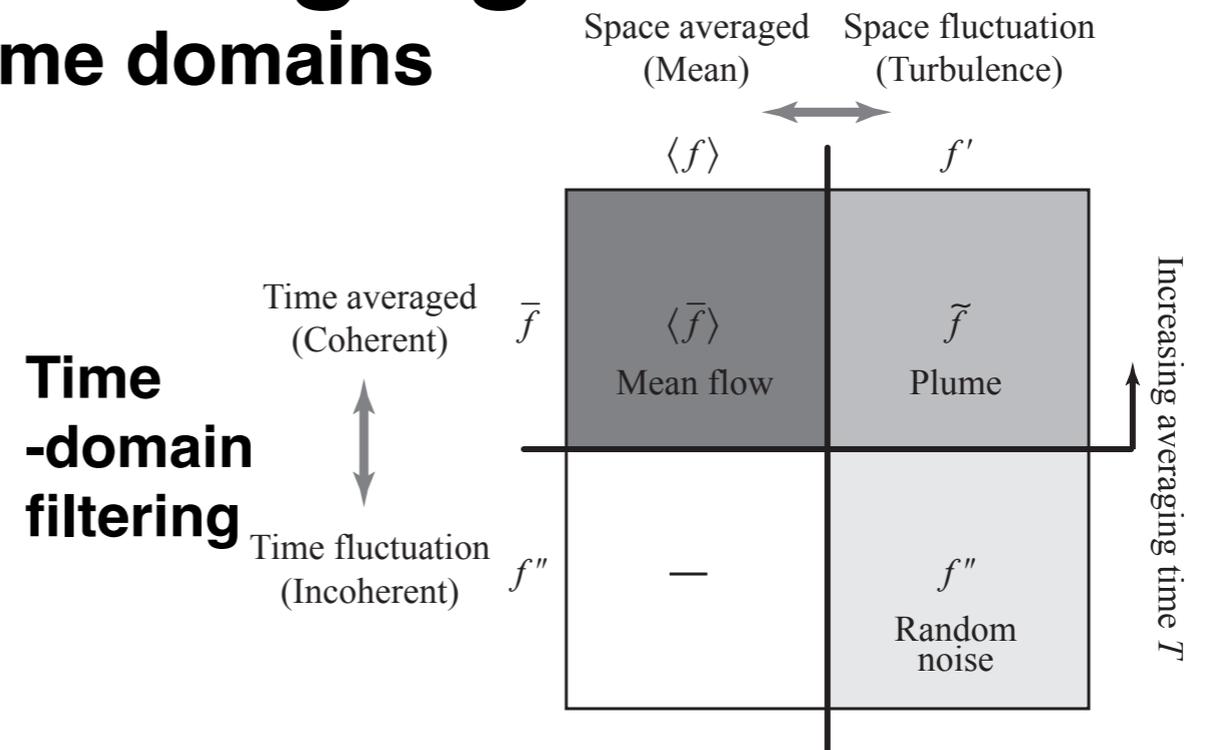
$T$ : Time average window

$$\tau \ll T \ll \Xi.$$

$\tau$ : Turn-over time       $\Xi$ : Mean-field evolution time

### Space average

$$\langle f \rangle(z; t) = \frac{1}{\Delta x \Delta y} \int_S f(x, y, z; t) dx dy$$



$$f = \underbrace{\langle \bar{f} \rangle}_{\text{coherent fluctuation}} + \underbrace{\tilde{f}}_{\bar{f} - \langle \bar{f} \rangle} + \underbrace{f''}_{f - \bar{f}} + \text{incoherent fluctuation}$$

# Coherent and incoherent fluctuations

$$f = \langle \bar{f} \rangle + \underbrace{\bar{f}}_{\substack{\tilde{f} \\ \bar{f} - \langle \bar{f} \rangle}} + \underbrace{f''}_{f - \bar{f}}$$

coherent fluctuation      incoherent fluctuation

Dispersion/Coherent fluctuation

$$\tilde{f} = \bar{f} - \langle \bar{f} \rangle$$

$$f' = \tilde{f} + f''$$

Rules of averaging

$$\langle \langle f \rangle \rangle = \langle f \rangle, \quad \overline{\bar{f}} = \bar{f} \quad \langle f' \rangle = 0, \quad \langle \tilde{f} \rangle = 0, \quad \langle f'' \rangle = 0, \quad \overline{f''} = 0$$

$$\langle \bar{f} \rangle = \langle f \rangle, \quad \overline{\langle f \rangle} = \langle f \rangle$$

$$\langle \langle f \rangle \langle g \rangle \rangle = \langle f \rangle \langle g \rangle, \quad \overline{\bar{f} \bar{g}} = \bar{f} \bar{g}$$

$$\langle \bar{f} \bar{g} \rangle = \langle \bar{f} \rangle \langle \bar{g} \rangle = \langle f \rangle \langle g \rangle$$

$$\langle \tilde{f} g'' \rangle = 0, \quad \langle f'' \tilde{g} \rangle = 0,$$

$$\overline{\tilde{f} g} = \tilde{f} \bar{g}, \quad \overline{\tilde{f} g''} = 0$$

$$\begin{aligned} \overline{\tilde{f} g} &= \overline{(\bar{f} - \langle \bar{f} \rangle) g} = \overline{\bar{f} g} - \overline{\langle \bar{f} \rangle g} = \bar{f} \bar{g} - \langle \bar{f} \rangle \bar{g} \\ &= (\bar{f} - \langle \bar{f} \rangle) \bar{g} = \tilde{f} \bar{g} \end{aligned}$$

# Turbulent energy equations

Coherent/dispersion  
fluctuation

$$\begin{aligned}
 & \left( \frac{\partial}{\partial t} + \langle u \rangle^\ell \frac{\partial}{\partial x^\ell} \right) \left\langle \frac{1}{2} (\tilde{u}^a)^2 \right\rangle \\
 &= - \langle \tilde{u}^a \tilde{u}^\ell \rangle \frac{\partial \langle u \rangle^a}{\partial x^\ell} - \frac{1}{\rho_0} \left\langle \tilde{p} \left( \frac{\partial \tilde{u}^a}{\partial x^a} \right) \right\rangle \\
 & - \nu \left\langle \left( \frac{\partial \tilde{u}^a}{\partial x^\ell} \right)^2 \right\rangle + \frac{g}{\Theta_0} \langle \tilde{\theta} \tilde{u}^\ell \rangle \delta^{\ell 3} \\
 & + \frac{\partial}{\partial x^\ell} \left\langle -\tilde{u}^\ell \frac{1}{2} (\tilde{u}^a)^2 + \tilde{p} \tilde{u}^\ell + \nu \frac{\partial}{\partial x^\ell} \frac{1}{2} (\tilde{u}^a)^2 \right\rangle \\
 & + \left\langle \widetilde{u''^\ell u''^a} \frac{\partial \tilde{u}^a}{\partial x^\ell} \right\rangle - \frac{\partial}{\partial x^\ell} \left\langle \widetilde{u''^\ell u''^a \tilde{u}^a} \right\rangle
 \end{aligned}$$

Incoherent/random  
fluctuation

$$\begin{aligned}
 & \left( \frac{\partial}{\partial t} + \langle u \rangle^\ell \frac{\partial}{\partial x^\ell} \right) \left\langle \frac{1}{2} (u''^a)^2 \right\rangle \\
 &= - \langle u''^a u''^\ell \rangle \frac{\partial \langle u \rangle^a}{\partial x^\ell} - \frac{1}{\rho_0} \left\langle p'' \left( \frac{\partial u''^a}{\partial x^a} \right) \right\rangle \\
 & - \nu \left\langle \frac{\partial u''^a}{\partial x^\ell} \frac{\partial u''^a}{\partial x^\ell} \right\rangle + \frac{g}{\Theta_0} \langle \theta'' u''^\ell \rangle \delta^{\ell 3} \\
 & + \frac{\partial}{\partial x^\ell} \left\langle -u''^\ell \frac{1}{2} (u''^a)^2 + p'' u''^\ell + \nu \frac{\partial}{\partial x^\ell} \frac{1}{2} (u''^a)^2 \right\rangle \\
 & - \left\langle \widetilde{u''^\ell u''^a} \frac{\partial \tilde{u}^a}{\partial x^\ell} \right\rangle - \frac{\partial}{\partial x^\ell} \left\langle \frac{1}{2} \widetilde{(u''^a)^2 \tilde{u}^\ell} \right\rangle
 \end{aligned}$$

$$\begin{aligned}
 \widetilde{u''^\alpha u''^\beta} &= \overline{u''^\alpha u''^\beta} - \langle \overline{u''^\alpha u''^\beta} \rangle \\
 &= \overline{u''^\alpha u''^\beta} - \langle u''^\alpha u''^\beta \rangle
 \end{aligned}$$

Dispersion part of the  
incoherent/random fluctuation correlation

# Modelling of transport with plumes

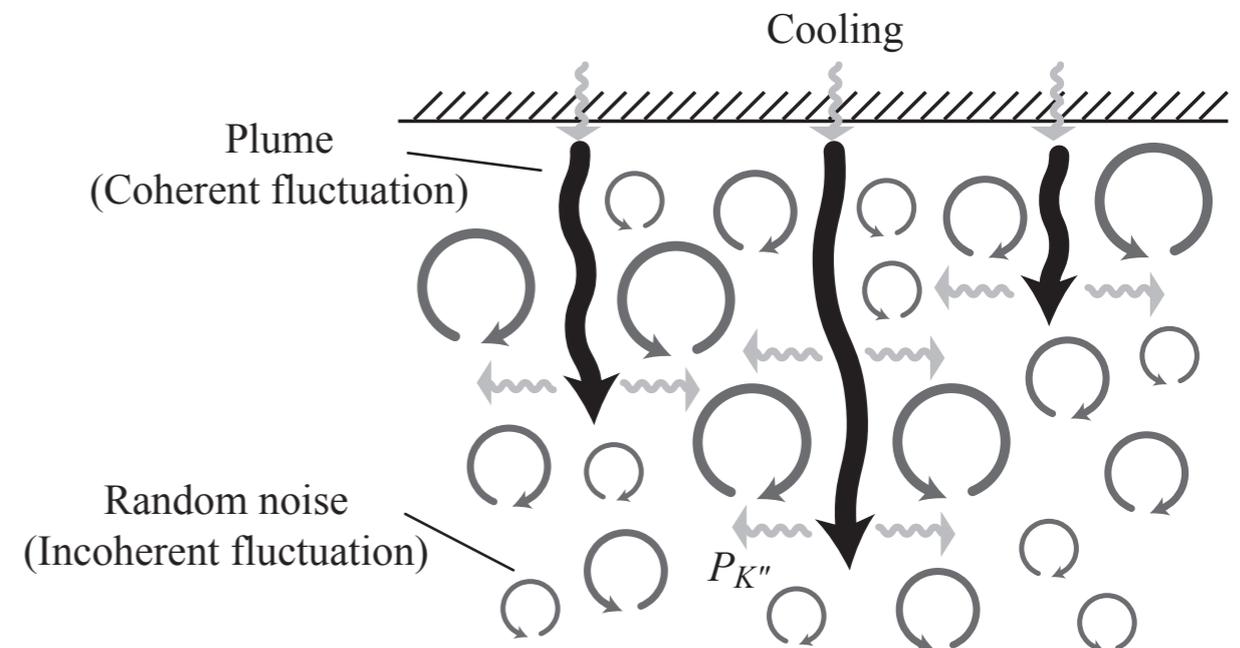
# Scenario: Interaction between coherent and incoherent fluctuations

Source of incoherent fluctuation energy

Sink of coherent fluctuation energy

$$P_{K''} = - \left\langle \widetilde{u''^l u''^m} \frac{\partial \widetilde{u}^m}{\partial x^l} \right\rangle = -P_{\tilde{K}}$$

In the presence of **the coherent velocity shear**, energy transfer between the coherent and incoherent fluctuation components occurs mediated by **the dispersion part of the random fluctuation velocity correlation**.



Due to **the non-equilibrium effect**, timescale of coherent fluctuations is altered, leading to **an enhancement of energy transfer to random fluctuations** if  $\Lambda < 0$ .

$$\widetilde{u''^l u''^m} \simeq -\tilde{\nu} \frac{\partial \widetilde{u}^m}{\partial x^l} + \text{N.E.}$$

$$P_{K''} = -\widetilde{u''^l u''^m} \frac{\partial \widetilde{u}^m}{\partial x^l} \simeq +\tilde{\nu}(1 - \Lambda) \left( \frac{\partial \widetilde{u}^m}{\partial x^l} \right)^2 > 0$$

$\Lambda < 0$  Enhancement of energy transfer to random fluctuation

# Turbulence modeling in the time-space double averaging

Turbulent internal-energy flux

$$\begin{aligned}\langle \overline{e' \mathbf{u}'} \rangle &= \langle \overline{\tilde{e} \tilde{\mathbf{u}}} \rangle + \langle \overline{e'' \mathbf{u}''} \rangle \\ &= - \left\langle \tilde{\kappa}_E \left( 1 - \tilde{C} \tilde{\tau} \frac{1}{\tilde{\mathbf{u}}^2} \frac{D \tilde{\mathbf{u}}^2}{Dt} \right) \nabla E \right\rangle - \left\langle \kappa_E'' \left( 1 - C'' \tau'' \frac{1}{\mathbf{u}''^2} \frac{D \mathbf{u}''^2}{Dt} \right) \nabla E \right\rangle\end{aligned}$$

$$\langle \overline{e' \mathbf{u}'} \rangle = -\kappa_{NE} \nabla E$$

$$\langle \overline{\mathbf{u}'^2} \rangle = \langle \overline{\tilde{\mathbf{u}}^2} \rangle + \langle \overline{\mathbf{u}''^2} \rangle$$

$$\langle \overline{\tilde{\mathbf{u}}^2} \rangle \simeq \langle \overline{\mathbf{u}''^2} \rangle \simeq \langle \overline{\mathbf{u}'^2} \rangle / 2$$

Coherent timescale

Incoherent timescale

$$\tilde{\tau} \gg \tau''$$

$$\tilde{\varepsilon} = \frac{\langle \overline{\tilde{\mathbf{u}}^2} \rangle}{\tilde{\tau}} \ll \frac{\langle \overline{\mathbf{u}''^2} \rangle}{\tau''} = \varepsilon''$$

Turbulent diffusivity with non-equilibrium effect

$$\kappa_{NE} = \begin{cases} \kappa_E \left[ 1 - C_{\tilde{\tau}} \frac{\tilde{\tau}}{\langle \overline{\mathbf{u}'^2} \rangle} \tilde{\Lambda}_D \right] & \text{for } \tilde{\Lambda}_D < 0 \\ \kappa_E \left[ 1 + C_{\tilde{\tau}} \frac{\tilde{\tau}}{\langle \overline{\mathbf{u}'^2} \rangle} \tilde{\Lambda}_D \right]^{-1} & \text{for } \tilde{\Lambda}_D > 0 \end{cases}$$

Non-equilibrium property  
along the plume motion

$$\tilde{\Lambda}_D = \left\langle (\tilde{\mathbf{u}} \cdot \nabla) \overline{\mathbf{u}'^2} \right\rangle$$

# Application to stellar convection

# Set-up of the convection simulation

Polytropic gas  $p = \rho^{1+\frac{1}{m}}$

Eq. of State  $p = (\gamma - 1)\rho e$   $e = \frac{p}{(\gamma - 1)\rho} = \frac{1}{\gamma - 1}\rho^{1/m}$   $p = \rho^{\frac{m+1}{m}} = (\gamma - 1)^{m+1}e^{m+1}$

Hydrostatic balance

$$-\frac{1}{\rho} \frac{\partial p}{\partial z} - g = 0 \quad \longrightarrow \quad \rho = \rho_s \left( \frac{e}{e_s} \right)^m \quad p = p_s \left( \frac{e}{e_s} \right)^{m+1}$$

$$e = e_s + \frac{g}{(\gamma - 1)(m + 1)}(z_s - z)$$

Determining the spatial distribution of the density and pressure

## Convection instability condition

$$\nabla - \nabla_{\text{ad}} > 0 \quad \text{where} \quad \nabla = \frac{\partial \ln \theta}{\partial \ln p} = \frac{p}{\theta} \frac{\partial \theta}{\partial p} \quad \nabla_{\text{ad}} = \left( \frac{\partial \ln \theta}{\partial \ln p} \right)_{\text{ad}} = \left( \frac{p}{\theta} \frac{\partial \theta}{\partial p} \right)_{\text{ad}}$$

For a polytropic gas  $\frac{\partial p}{\partial \theta} = (m + 1) \frac{p}{\theta}$   $\nabla = \frac{\partial \ln \theta}{\partial \ln p} = \frac{p}{\theta} \frac{\partial \theta}{\partial p} = \frac{1}{m + 1}$

In the adiabatic case  $e = \frac{1}{\gamma - 1} \rho^{\gamma-1} = \frac{1}{\gamma - 1} p^{1-\frac{1}{\gamma}}$   $\nabla_{\text{ad}} = \frac{\partial \ln e}{\partial \ln p} = \frac{\partial \ln \theta}{\partial \ln p} = 1 - \frac{1}{\gamma} = \frac{\gamma - 1}{\gamma}$

### Instability condition

$$m < \frac{\gamma}{\gamma - 1} - 1 \quad \gamma = 5/3 \quad \longrightarrow$$

$m < 1.5$  Unstable

$m = 1.5$  Marginally stable

# Two-layer polytropic gas convection

(Cossette & Rast 2016, ApJL, 829, L17)

$$\begin{cases} e(z_i \leq z \leq z_s) = e_s + \frac{g(z_s - z)}{(\gamma - 1)(m_s + 1)} \\ e(z_b \leq z \leq z_i) = e_i + \frac{g(z_i - z)}{(\gamma - 1)(m_i + 1)} \end{cases}$$

Control parameter:  $e_s$

**specific internal energy at the surface ( $z_s$ )**

Local model:  $m_s = m_i = 1.495$

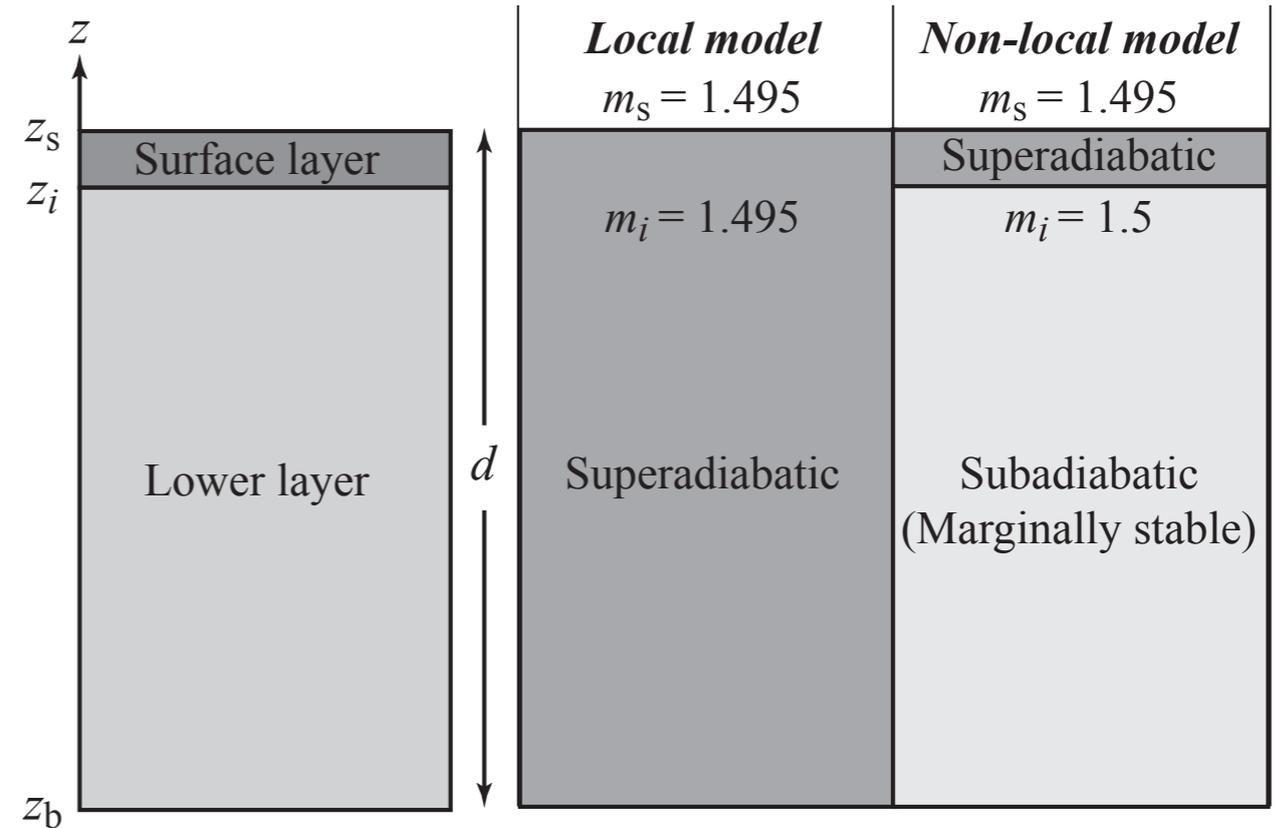
Non-local model:  $m_s = 1.495, m_i = 1.5 \quad z_i/d = 0.95$

$$\nu = \eta = 1 \times 10^{-4}, \quad \kappa = 1 \times 10^{-4} \quad (\text{conductivity})$$

Non-dimensional parameters  $Pr = \frac{\nu}{\chi}, \quad Pm = \frac{\nu}{\eta}, \quad Ra = \frac{gd^4\delta}{\chi\nu H_\rho}$  with  $\chi = \frac{\kappa}{\langle\rho\rangle}$

$$Pr \simeq 1, \quad Pm = 1, \quad Ra \simeq 4.2 \times 10^6$$

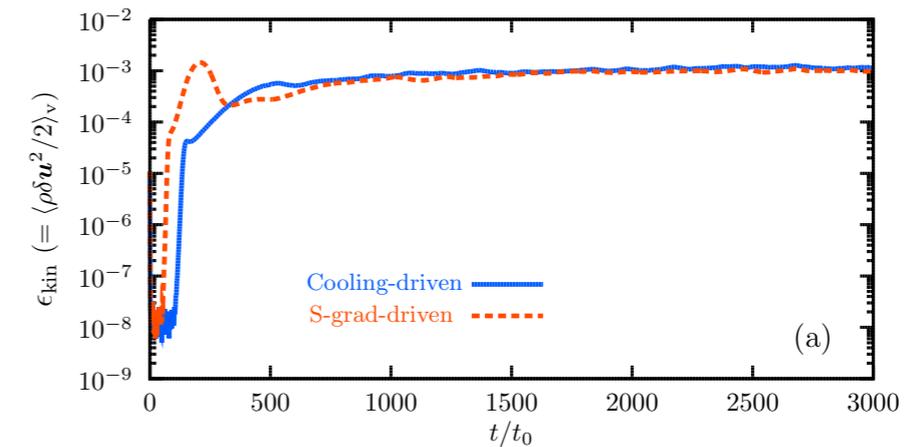
$$\rho_{\text{bottom}}/\rho_{\text{top}} = 100$$



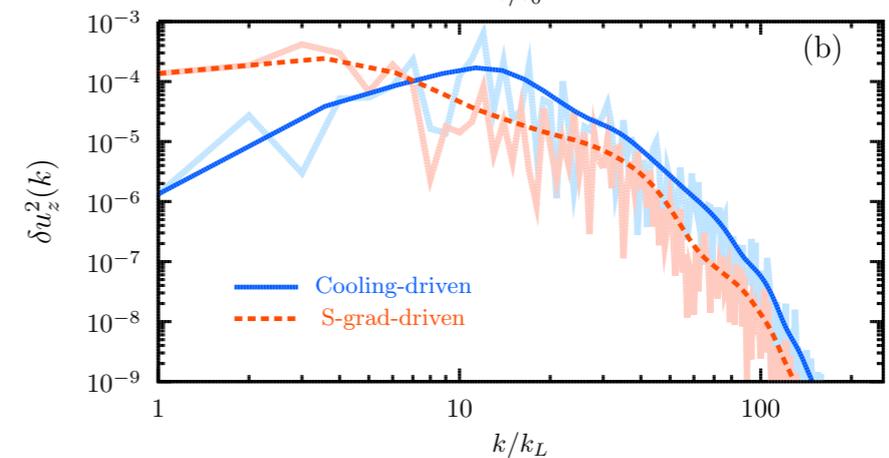
throughout the convection zone near surface region

# Results

Time evolution of kinetic energy



Energy spectra

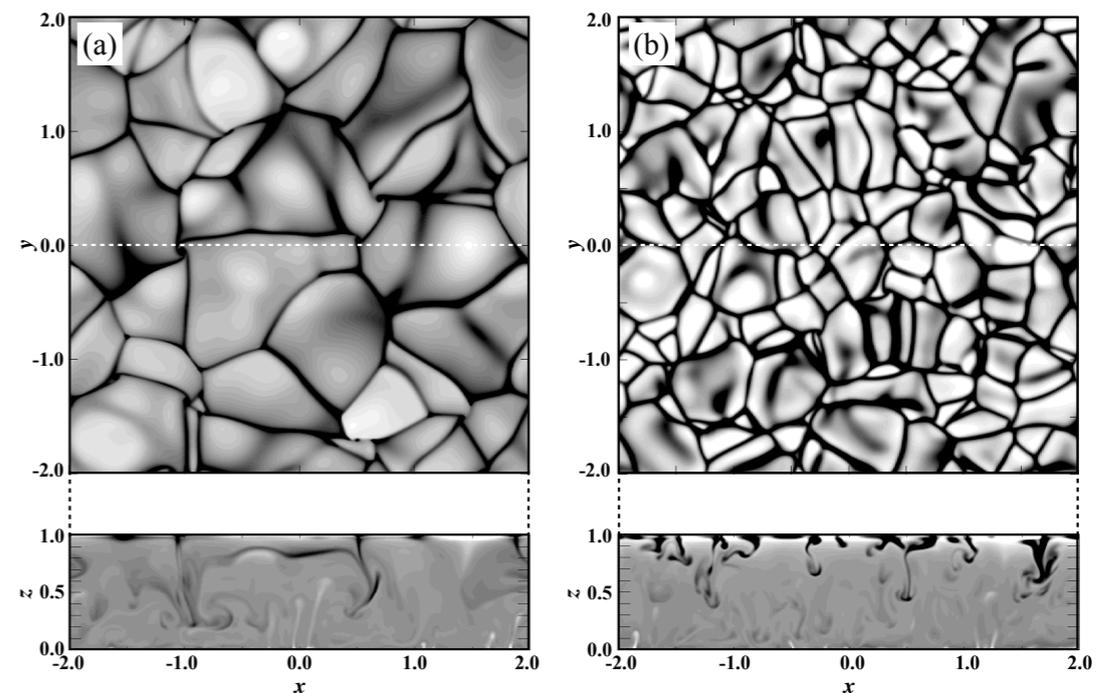


Entropy distributions

Horizontal cross-sections of the entropy fluctuation at the top surface

Vertical cross-sections of the entropy fluctuation from the horizontal mean

Locally driven case Non-locally driven case

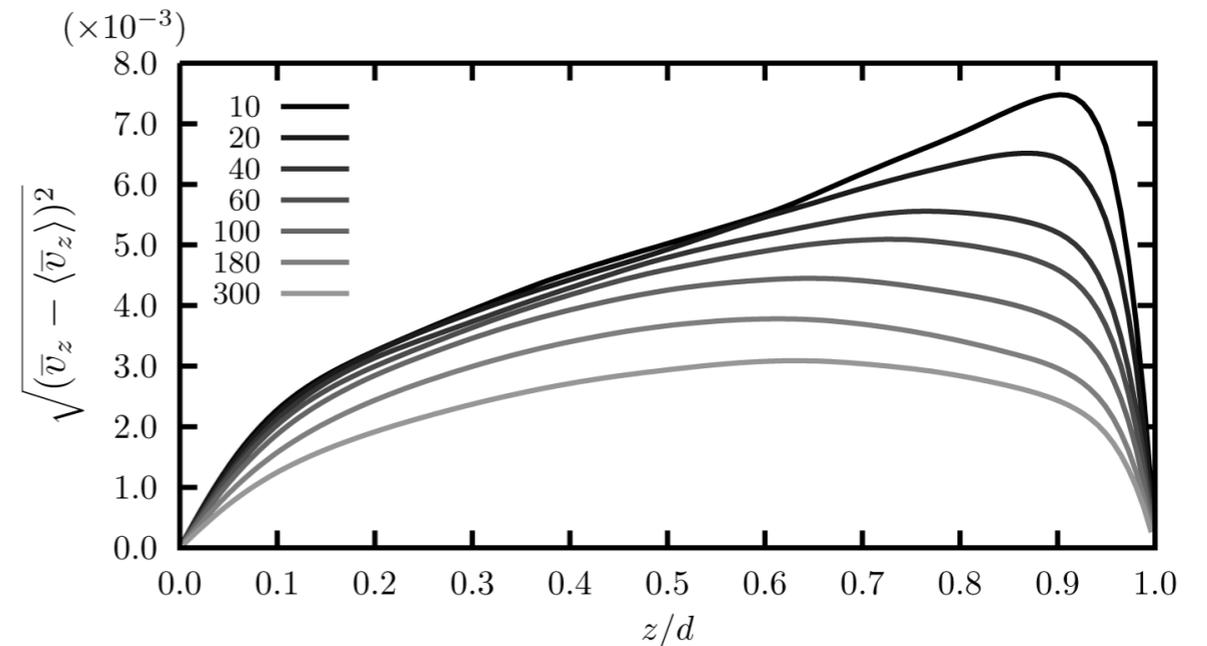
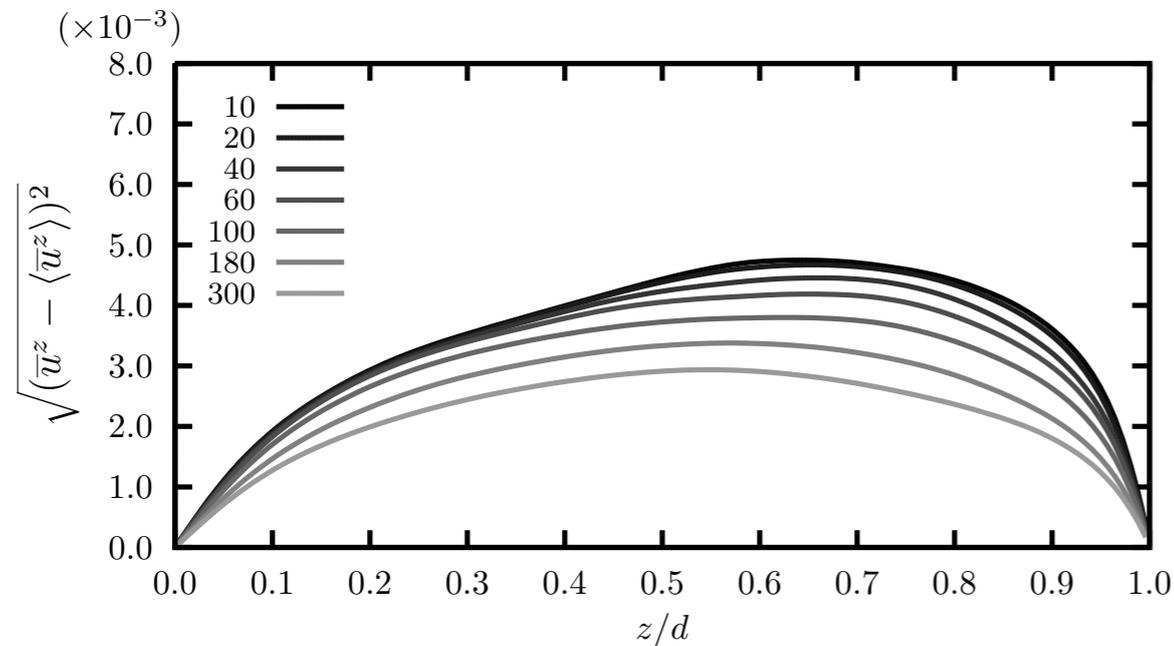


# Coherent fluctuations

$$\sqrt{\langle (\tilde{u}^z)^2 \rangle} = \sqrt{\langle (\bar{u}^z - \langle \bar{u}^z \rangle)^2 \rangle}$$

Local (superadiabatic driven)

Non-local (surface cooling driven)



## Lifetime of plume and averaging time

Non-dimensionalised horizontally averaged velocity is about

$$v^z = \sqrt{\langle (u^z)^2 \rangle} \sim 0.01$$

Non-dimensionalised full depth of the convection zone

$$d = 1$$

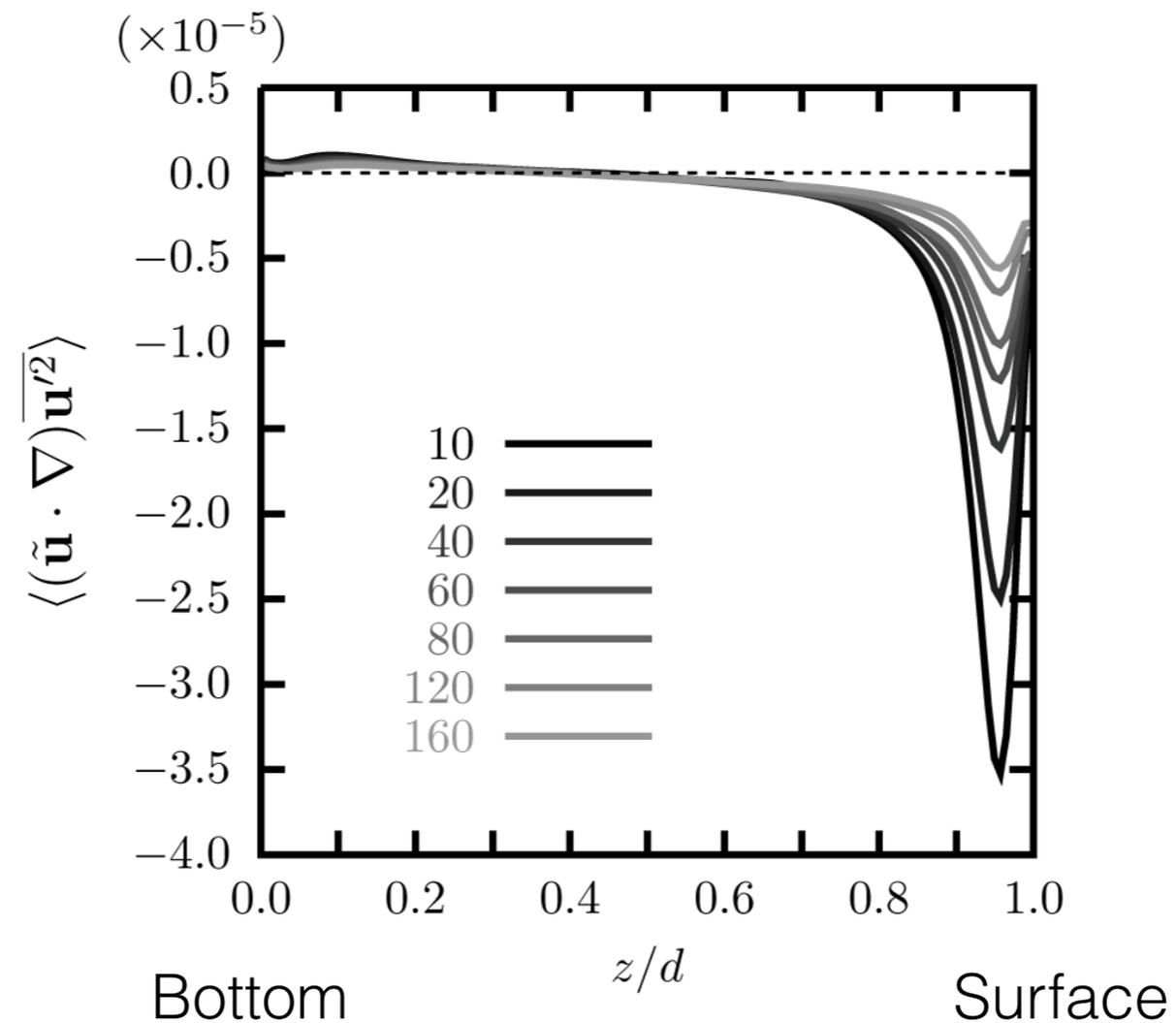
Lifetime of the plume may be estimated (from crossing time) as

$$\frac{d/2}{v^z} \sim \frac{0.5}{0.01} \sim 50$$

→ Averaging time  $T \lesssim 25$

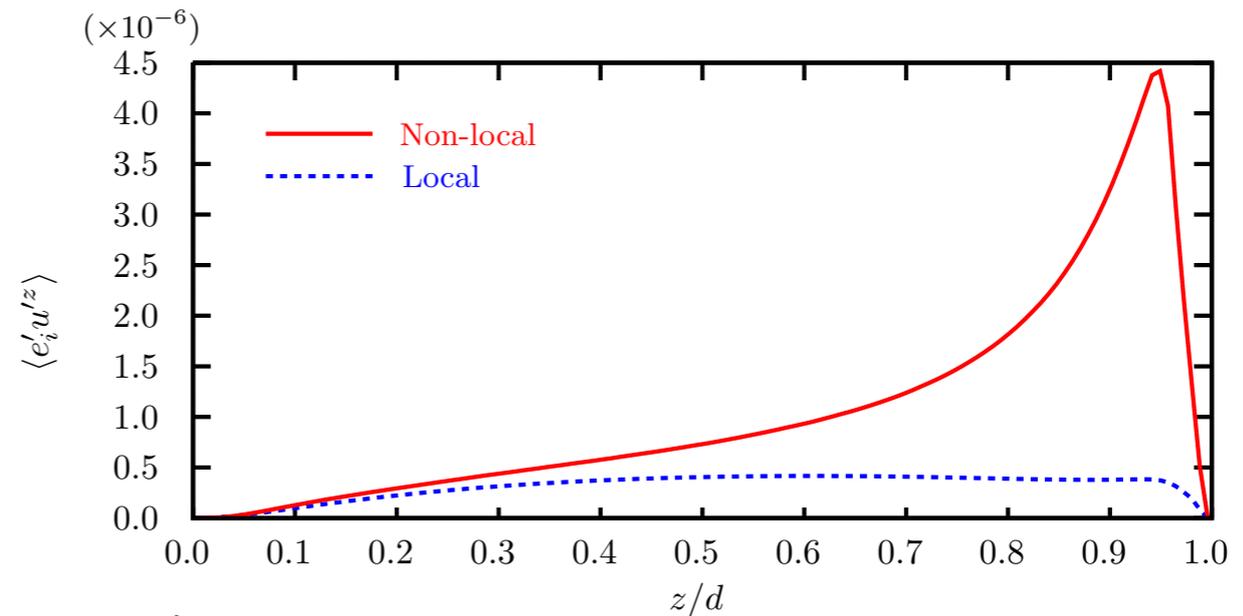
# Non-equilibrium property along the plume motion

$$\tilde{\Lambda}_D = \left\langle (\tilde{\mathbf{u}} \cdot \nabla) \overline{\mathbf{u}'^2} \right\rangle$$

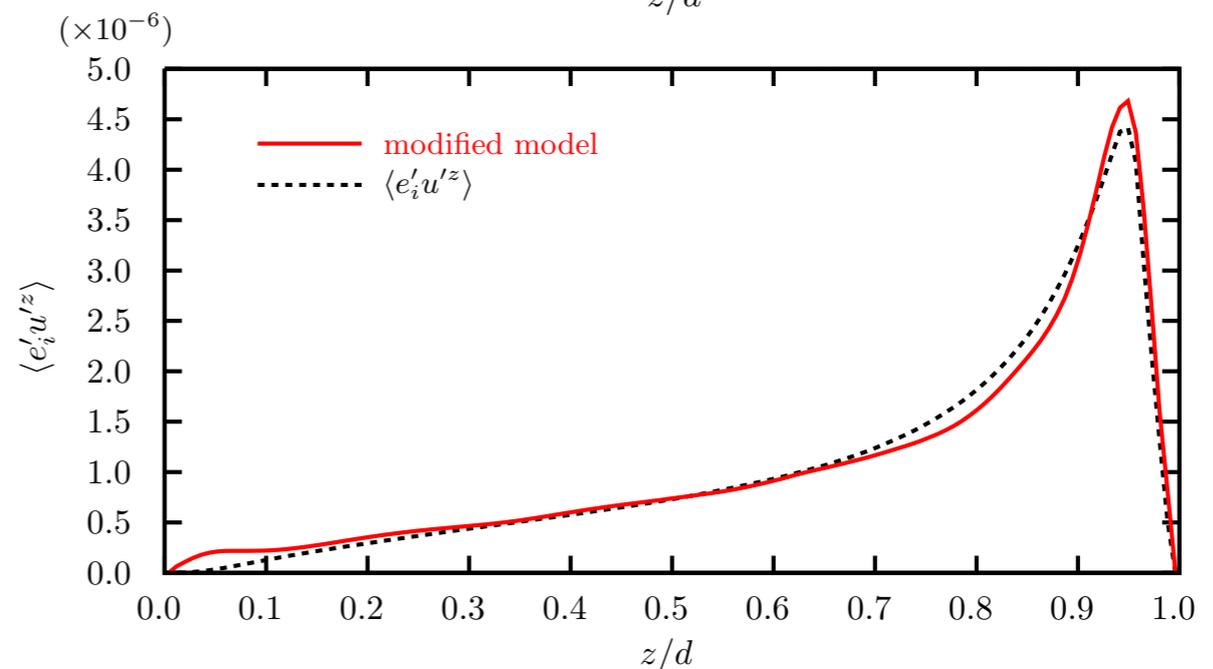


# Turbulent energy flux $\langle e' u'^z \rangle$

by DNS



by Non-equilibrium model



A similar result is obtained for the turbulent mass flux  $\langle \rho' u'^z \rangle$

The **non-equilibrium effect** in the time–space **double averaging** framework is a promising model approach to the **stellar convection with plume**.

# Summary of non-equilibrium effects on convective plumes

- Transport due to plume motion is incorporated into turbulence model through the **non-equilibrium effect through the time variation of fluctuations along the plume motion**
- Turbulence modelling in the **time–space double averaging framework**
- Turbulent convective transport of **surface cooling diving plumes (non-local transport) is properly described by the non-equilibrium model**

## Wish list

## Physics of plume formation

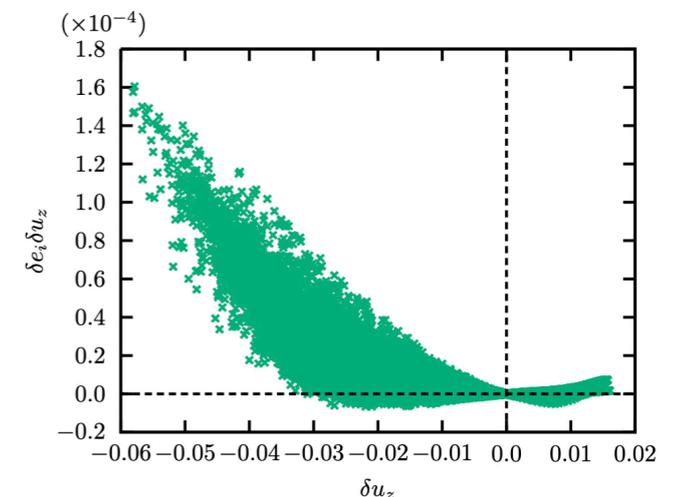
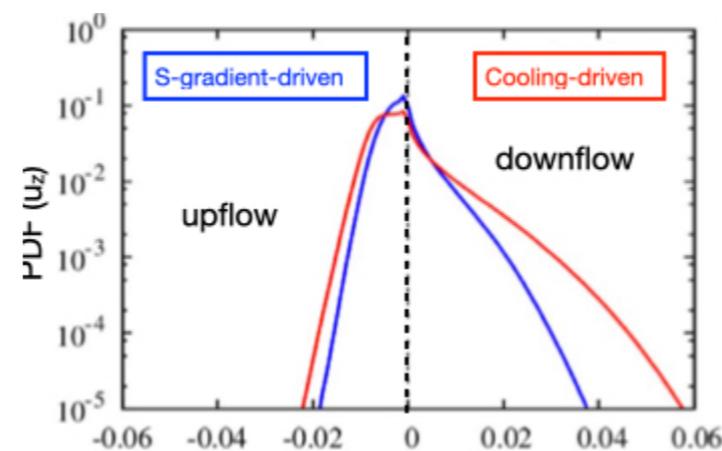
Statistical properties (probability distributions)

(Masada, Private communication)

Dynamical properties (length, aspect ratio, shape)

What determines the non-equilibrium properties of turbulence

Entropy production rate etc.



# Non-equilibrium effect in dynamos

## Cross-interaction responses

**Yokoi, N.** “Unappreciated cross-helicity effects in plasma physics: Anti-diffusion effects in dynamo and momentum transport,” **Rev. Mod. Plasma Phys. 7, 33-1-98 (2023)**

<https://doi.org/10.1007/s41614-023-00133-4>

**Mizerski, K., Yokoi, N. & Brandenburg, A.** “Cross-helicity effect on  $\alpha$ -type dynamo in non-equilibrium turbulence,” **J. Plasma Phys. 89, 905890412 (2023)**

<https://doi.org/10.1017/S0022377823000545>

# Global flow generation by cross helicity

Reynolds and turbulent Maxwell stress

Theoretical Result

$$\langle \mathbf{u}'\mathbf{u}' - \mathbf{b}'\mathbf{b}' \rangle_D = -\nu_K \mathcal{S} + \nu_M \mathcal{M} + \eta_H \Omega_* \nabla H + \dots$$

eddy viscosity
inhomogeneous helicity  
cross helicity

D: deviatoric part

$\mathcal{S}$ : mean velocity strain  $\mathcal{S} = \nabla \mathbf{U} + (\nabla \mathbf{U})^\dagger$

$\mathcal{M}$ : mean magnetic-field strain  $\mathcal{M} = \nabla \mathbf{B} + (\nabla \mathbf{B})^\dagger$

$\Omega_*$ : absolute mean vorticity (mean vorticity + rotation)

cf.  $\langle \mathbf{u}' \times \mathbf{b}' \rangle = -\eta_T \nabla \times \mathbf{B} + \gamma \nabla \times \mathbf{U} + \alpha \mathbf{B} + \dots$

**Turbulent cross helicity** coupled with **mean magnetic-field strain** may contribute to **transport suppression** and/or **global flow generation** against the<sup>101</sup> eddy-viscosity effect

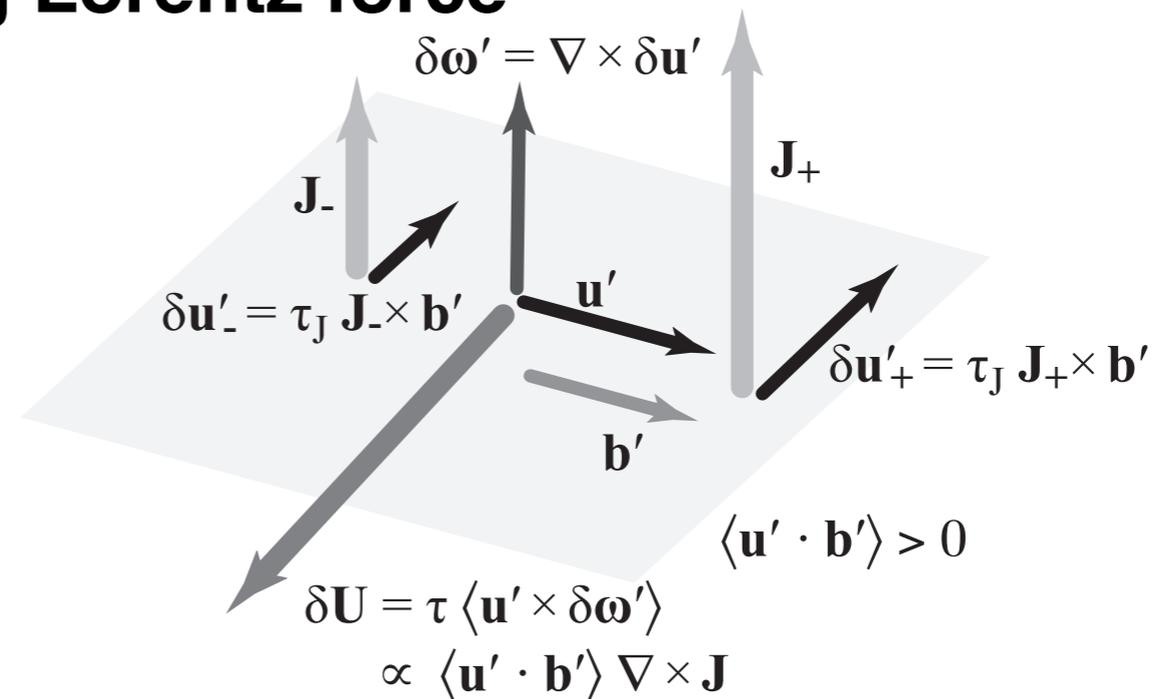
# Physical interpretation of large-scale flow generation by cross helicity

Velocity fluctuation induced by **fluctuating Lorentz force**

$$\delta \mathbf{u}' = \tau_J \mathbf{J} \times \mathbf{b}'$$

Associated vorticity

$$\begin{aligned} \delta \boldsymbol{\omega}' &= \nabla \times \delta \mathbf{u}' \\ &= \tau_J \nabla \times (\mathbf{J} \times \mathbf{b}') \\ &\simeq \tau_J (\mathbf{b}' \cdot \nabla) \mathbf{J} \end{aligned}$$



Non-trivial mean electric-current distribution (Inhomogeneous  $\mathbf{J}$ ) is required

$$\delta \mathbf{U} = \tau \langle \mathbf{u}' \times \delta \boldsymbol{\omega}' \rangle \propto \langle \mathbf{u}' \cdot \mathbf{b}' \rangle \nabla \times \mathbf{J} \quad \text{in the direction of } \nabla \times \mathbf{J}$$

## Theoretical formulation (e.g., incompressible MHD)

$$\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} = (\mathbf{b} \cdot \nabla) \mathbf{b} - \nabla p_M - 2\boldsymbol{\omega}_F \times \mathbf{u} + \nu \nabla^2 \mathbf{u}$$

$$\frac{\partial \mathbf{b}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{b} = (\mathbf{b} \cdot \nabla) \mathbf{u} + \eta \nabla^2 \mathbf{b}$$

$$\nabla \cdot \mathbf{u} = \nabla \cdot \mathbf{b} = 0$$

Multiple-scale analysis  $\boldsymbol{\xi}(= \mathbf{x}), \mathbf{X}(= \delta \mathbf{x}), \tau(= t), T(= \delta t)$

$$\nabla_{\mathbf{x}} = \nabla_{\boldsymbol{\xi}} + \delta \nabla_{\mathbf{X}}; \quad \frac{\partial}{\partial t} = \frac{\partial}{\partial \tau} + \delta \frac{\partial}{\partial T}$$

$$f(\mathbf{x}; t) = F(\mathbf{X}; T) + f'(\boldsymbol{\xi}, \mathbf{X}; \tau, T)$$

## Two-scale equations in configuration space

$$\frac{\partial u'^i}{\partial \tau} + U^j \frac{\partial u'^i}{\partial \xi^j} + \frac{\partial}{\partial \xi^j} (u'^j u'^i - b'^j b'^i) + \frac{\partial p'_M}{\partial \xi^i} - \nu \frac{\partial^2 u'^i}{\partial \xi^j \partial \xi^j} - B^j \frac{\partial b'^i}{\partial \xi^j} \quad \frac{\partial u'^i}{\partial \xi^i} + \delta \frac{\partial u'^i}{\partial X^i} = 0,$$

$$= \delta \left[ b'^j \frac{\partial B^i}{\partial X^j} - u'^j \left( \frac{\partial U^i}{\partial X^j} + \epsilon^{jik} \Omega_0^k \right) + B^j \frac{\partial b'^i}{\partial X^j} - \frac{\overline{D} u'^i}{DT} - \frac{\partial}{\partial X^j} (u'^j u'^i - b'^j b'^i - \langle u'^j u'^i - b'^j b'^i \rangle) - \frac{\partial p'_M}{\partial X^i} \right]$$

$$\frac{\partial b'^i}{\partial \tau} + U^j \frac{\partial b'^i}{\partial \xi^j} + \frac{\partial}{\partial \xi^j} (u'^j b'^i - b'^j u'^i) - \eta \frac{\partial^2 b'^i}{\partial \xi^j \partial \xi^j} - B^j \frac{\partial u'^i}{\partial \xi^j} \quad \frac{\partial b'^i}{\partial \xi^i} + \delta \frac{\partial b'^i}{\partial X^i} = 0$$

$$= \delta \left[ u'^j \frac{\partial B^i}{\partial X^j} - b'^j \left( \frac{\partial U^i}{\partial X^j} + \epsilon^{jik} \Omega_0^k \right) + B^j \frac{\partial u'^i}{\partial X^j} - \frac{\overline{D} b'^i}{DT} - \frac{\partial}{\partial X^j} (u'^j b'^i - b'^j u'^i - \langle u'^j b'^i - b'^j u'^i \rangle) \right]$$

inhomogeneities,  
anisotropies,  
non-equilibrium properties

## Scale parameter expansion

$$f^i(\mathbf{k}; \tau) = \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \delta^n f_{nm}^i(\mathbf{k}; \tau) - \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \delta^{n+1} i \frac{k^i}{k^2} \frac{\partial}{\partial X_I^j} f_{nm}^j(\mathbf{k}; \tau)$$

## Basic-field (lowest-order field) equations

$$\begin{aligned} \frac{\partial u_{00}^i(\mathbf{k}; \tau)}{\partial \tau} + \nu k^2 u_{00}^i(\mathbf{k}; \tau) \\ - i M^{ijk}(\mathbf{k}) \iint d\mathbf{p} d\mathbf{q} \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) \left[ u_{00}^j(\mathbf{p}; \tau) u_{00}^k(\mathbf{q}; \tau) - b_{00}^j(\mathbf{p}; \tau) b_{00}^k(\mathbf{q}; \tau) \right] = 0 \end{aligned}$$

$$\begin{aligned} \frac{\partial b_{00}^i(\mathbf{k}; \tau)}{\partial \tau} + \eta k^2 b_{00}^i(\mathbf{k}; \tau) \\ - i N^{ijk}(\mathbf{k}) \iint d\mathbf{p} d\mathbf{q} \delta(\mathbf{k} - \mathbf{p} - \mathbf{q}) \left[ u_{00}^j(\mathbf{p}; \tau) b_{00}^k(\mathbf{q}; \tau) - b_{00}^j(\mathbf{p}; \tau) u_{00}^k(\mathbf{q}; \tau) \right] = 0 \end{aligned}$$

## Projection operators

$$M^{ijk}(\mathbf{k}) = \frac{1}{2} \left[ k^j D^{ik}(\mathbf{k}) + k^k D^{ij}(\mathbf{k}) \right], \quad D^{ij}(\mathbf{k}) = \delta^{ij} - \frac{k^i k^j}{k^2}$$

$$N^{ijk}(\mathbf{k}) = k^j \delta^{ik} - k^k \delta^{ij}$$

$f_{01}(\mathbf{k}; \tau)$  equation

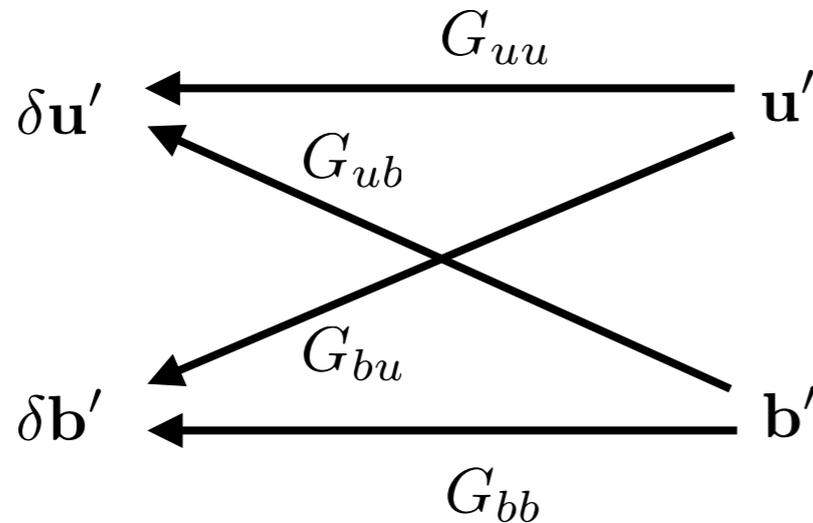
$$\begin{aligned}
 & \begin{pmatrix} \frac{\partial}{\partial \tau} + \nu k^2 & 0 \\ 0 & \frac{\partial}{\partial \tau} + \eta k^2 \end{pmatrix} \begin{pmatrix} u_{01}^i(\mathbf{k}; \tau) \\ b_{01}^i(\mathbf{k}; \tau) \end{pmatrix} \\
 & + i \begin{pmatrix} -2M^{ikm}(\mathbf{k}) \int_{\Delta} u_{00}^k(\mathbf{p}; \tau) & 2M^{ikm}(\mathbf{k}) \int_{\Delta} b_{00}^k(\mathbf{p}; \tau) \\ N^{ikm}(\mathbf{k}) \int_{\Delta} b_{00}^k(\mathbf{p}; \tau) & -N^{ikm}(\mathbf{k}) \int_{\Delta} u_{00}^k(\mathbf{p}; \tau) \end{pmatrix} \begin{pmatrix} u_{01}^m(\mathbf{q}; \tau) \\ b_{01}^m(\mathbf{q}; \tau) \end{pmatrix} \\
 & = -ik^j B^j \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} u_{00}^i(\mathbf{k}) \\ b_{00}^i(\mathbf{k}) \end{pmatrix} \quad \text{with} \quad \int_{\Delta} = \iint d\mathbf{p}d\mathbf{q} \delta(\mathbf{k} - \mathbf{p} - \mathbf{q})
 \end{aligned}$$

$f_{10}(\mathbf{k}; \tau)$  equation

$$\begin{aligned}
 & \begin{pmatrix} \frac{\partial}{\partial \tau} + \nu k^2 & 0 \\ 0 & \frac{\partial}{\partial \tau} + \eta k^2 \end{pmatrix} \begin{pmatrix} u_{10}^i(\mathbf{k}; \tau) \\ b_{10}^i(\mathbf{k}; \tau) \end{pmatrix} \\
 & + i \begin{pmatrix} -2M^{ikm}(\mathbf{k}) \int_{\Delta} u_{00}^k(\mathbf{p}; \tau) & 2M^{ikm}(\mathbf{k}) \int_{\Delta} b_{00}^k(\mathbf{p}; \tau) \\ N^{ikm}(\mathbf{k}) \int_{\Delta} b_{00}^k(\mathbf{p}; \tau) & -N^{ikm}(\mathbf{k}) \int_{\Delta} u_{00}^k(\mathbf{p}; \tau) \end{pmatrix} \begin{pmatrix} u_{10}^m(\mathbf{q}; \tau) \\ b_{10}^m(\mathbf{q}; \tau) \end{pmatrix} \\
 & = B^k \frac{\partial}{\partial X_I^k} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} u_{00}^i(\mathbf{k}) \\ b_{00}^i(\mathbf{k}) \end{pmatrix} - D^{jk}(\mathbf{k}) \frac{\bar{D}}{DT_I} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} u_{00}^k(\mathbf{k}) \\ b_{00}^k(\mathbf{k}) \end{pmatrix} \\
 & + \begin{pmatrix} -D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) & D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} \\ -D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} & D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) \end{pmatrix} \begin{pmatrix} u_{00}^m(\mathbf{k}) \\ b_{00}^m(\mathbf{k}) \end{pmatrix}
 \end{aligned}$$

# Response functions in MHD

Cross-interaction responses in incompressible MHD



cf. Hydrodynamic case



Green's function equations

$$\begin{pmatrix} \frac{\partial}{\partial \tau} + \nu k^2 & 0 \\ 0 & \frac{\partial}{\partial \tau} + \eta k^2 \end{pmatrix} \begin{pmatrix} G_{uu}^{ij} & G_{bu}^{ij} \\ G_{ub}^{ij} & G_{bb}^{ij} \end{pmatrix} + i \begin{pmatrix} -2M^{ikm} \int_{\Delta} u_{00}^k & 2M^{ikm} \int_{\Delta} b_{00}^k \\ N^{ikm} \int_{\Delta} b_{00}^k & -N^{ikm} \int_{\Delta} u_{00}^k \end{pmatrix} \begin{pmatrix} G_{uu}^{mj} & G_{bu}^{mj} \\ G_{ub}^{mj} & G_{bb}^{mj} \end{pmatrix} = \delta^{ij} \delta(\tau - \tau') \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

## $f_{10}(\mathbf{k}; \tau)$ solution

$$\begin{aligned}
 u_{10}^i(\mathbf{k}; \tau) = & \int_{-\infty}^{\tau} d\tau_1 G_{uu}^{ij}(\mathbf{k}; \tau, \tau_1) \left[ -D^{jk}(\mathbf{k}) \frac{\overline{D}u_{00}^k(\mathbf{k}; \tau_1)}{DT_I} - D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) u_{00}^m(\mathbf{k}; \tau_1) \right. \\
 & \left. + B^k \frac{\partial b_{00}^i(\mathbf{k}; \tau_1)}{\partial X_I^k} + D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} b_{00}^m(\mathbf{k}; \tau_1) \right] \\
 & + \int_{-\infty}^{\tau} d\tau_1 G_{ub}^{ij}(\mathbf{k}; \tau, \tau_1) \left[ B^k \frac{\partial u_{00}^i(\mathbf{k}; \tau_1)}{\partial X_I^k} - D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} u_{00}^m(\mathbf{k}; \tau_1) \right. \\
 & \left. - D^{jk}(\mathbf{k}) \frac{\overline{D}b_{00}^k(\mathbf{k}; \tau_1)}{DT_I} + D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) b_{00}^m(\mathbf{k}; \tau_1) \right]
 \end{aligned}$$

$$\begin{aligned}
 b_{10}^i(\mathbf{k}; \tau) = & \int_{-\infty}^{\tau} d\tau_1 G_{bu}^{ij}(\mathbf{k}; \tau, \tau_1) \left[ -D^{jk}(\mathbf{k}) \frac{\overline{D}u_{00}^k(\mathbf{k}; \tau_1)}{DT_I} - D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) u_{00}^m(\mathbf{k}; \tau_1) \right. \\
 & \left. + B^k \frac{\partial b_{00}^i(\mathbf{k}; \tau_1)}{\partial X_I^k} + D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} b_{00}^m(\mathbf{k}; \tau_1) \right] \\
 & + \int_{-\infty}^{\tau} d\tau_1 G_{bb}^{ij}(\mathbf{k}; \tau, \tau_1) \left[ B^k \frac{\partial u_{00}^i(\mathbf{k}; \tau_1)}{\partial X_I^k} - D^{jk}(\mathbf{k}) \frac{\partial B^k}{\partial X^m} u_{00}^m(\mathbf{k}; \tau_1) \right. \\
 & \left. - D^{jk}(\mathbf{k}) \frac{\overline{D}b_{00}^k(\mathbf{k}; \tau_1)}{DT_I} + D^{jk}(\mathbf{k}) \left( \frac{\partial U^k}{\partial X^m} + \epsilon^{mkn} \Omega_0^n \right) b_{00}^m(\mathbf{k}; \tau_1) \right]
 \end{aligned}$$

# Turbulent electromotive force (EMF)

$$E_M^i \equiv \epsilon^{ijk} \langle u'^j b'^k \rangle = \epsilon^{ijk} \int d\mathbf{k} \langle u^j(\mathbf{k}; \tau) b^k(\mathbf{k}'; \tau) \rangle / \delta(\mathbf{k} + \mathbf{k}').$$

$$\langle u^j b^k \rangle = \langle u_{00}^j b_{00}^k \rangle + \langle u_{01}^j b_{00}^k \rangle + \langle u_{00}^j b_{01}^k \rangle + \delta \langle u_{10}^j b_{00}^k \rangle + \delta \langle u_{00}^j b_{10}^k \rangle + \dots$$

α effect	Turb. Mag. Diffusivity	Turb. Pumping	Cross-helicity effect
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$$\langle \mathbf{u}' \times \mathbf{b}' \rangle = \alpha \mathbf{B} - (\beta + \zeta) \nabla \times \mathbf{B} - (\nabla \zeta) \times \mathbf{B} + \gamma \nabla \times \mathbf{U}$$

$$I\{A, B\} = \int d\mathbf{k} \int_{-\infty}^{\tau} d\tau_1 A(k; \tau, \tau_1) B(k; \tau, \tau_1)$$

Cross interaction mediated by  $G_{ub}$  and  $G_{bu}$

$$\alpha = \frac{1}{3} [-I\{G_{bb}, H_{uu}\} + I\{G_{uu}, H_{bb}\} - I\{G_{bu}, H_{ub}\} + I\{G_{ub}, H_{bu}\}]$$

$$\beta = \frac{1}{3} [I\{G_{bb}, Q_{uu}\} + I\{G_{uu}, Q_{bb}\} - I\{G_{bu}, Q_{ub}\} - I\{G_{ub}, Q_{bu}\}]$$

$$\zeta = \frac{1}{3} [I\{G_{bb}, Q_{uu}\} - I\{G_{uu}, Q_{bb}\} + I\{G_{bu}, Q_{ub}\} - I\{G_{ub}, Q_{bu}\}]$$

$$\gamma = \frac{1}{3} [I\{G_{bb}, Q_{ub}\} + I\{G_{uu}, Q_{bu}\} - I\{G_{bu}, Q_{uu}\} - I\{G_{ub}, Q_{bb}\}]$$

# Dynamo due to the cross-interaction responses

Torsional cross correlations

$$\begin{aligned} \alpha_X &= \frac{1}{3} [-I\{G_{bu}, H_{ub}\} + I\{G_{ub}, H_{bu}\}] \longrightarrow \alpha_X = -\tau_{bu} \langle \mathbf{u}' \cdot \mathbf{j}' \rangle + \tau_{ub} \langle \boldsymbol{\omega}' \cdot \mathbf{b}' \rangle \\ \beta_X &= \frac{1}{3} [-I\{G_{bu}, Q_{ub}\} - I\{G_{ub}, Q_{bu}\}] \longrightarrow \beta_X = -(\tau_{bu} + \tau_{ub}) \langle \mathbf{u}' \cdot \mathbf{b}' \rangle \\ \zeta_X &= \frac{1}{3} [+I\{G_{bu}, Q_{ub}\} - I\{G_{ub}, Q_{bu}\}] \longrightarrow \zeta_X = (\tau_{bu} - \tau_{ub}) \langle \mathbf{u}' \cdot \mathbf{b}' \rangle \\ \gamma_X &= \frac{1}{3} [-I\{G_{bu}, Q_{uu}\} - I\{G_{ub}, Q_{bb}\}] \longrightarrow \gamma_X = -\tau_{bu} \langle \mathbf{u}'^2 \rangle - \tau_{ub} \langle \mathbf{b}'^2 \rangle \end{aligned}$$

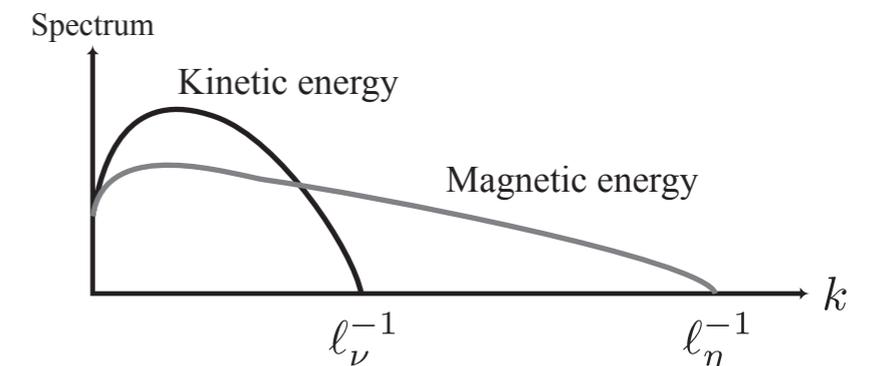
- **Vanish if**  $G_{ub} = G_{bu} = 0$  No cross-interaction response effects

- **Finite cross-interaction responses with same timescales**

$$\begin{aligned} \alpha_X &= \tau_{bu} (-\langle \mathbf{u}' \cdot \mathbf{j}' \rangle + \langle \boldsymbol{\omega}' \cdot \mathbf{b}' \rangle) \quad \tau_{bu} = \tau_{ub} \quad G_{ub} = G_{bu} \neq 0 \\ &= \tau_{bu} \nabla \cdot \langle \mathbf{u}' \times \mathbf{b}' \rangle \quad \text{EMF flux across the boundary} \end{aligned}$$

- **Timescale difference in responses**  $G_{ub} \neq G_{bu}$

$$\tau_{bu} \neq \tau_{ub} \quad Pm \gg 1 \text{ or } Pm \ll 1$$



- **Non-equilibrium properties**  $H_{ub}(\tau, \tau_1) \neq H_{ub}(\tau_1, \tau)$

**Present talk**

# Cross-interaction $\alpha_X$

$$\alpha_X = -\frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 G_{bu}(k, \mathbf{X}; \tau, \tau_1, T) H_{ub}(k, \mathbf{X}; \tau, \tau_1, T) \\ + \frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 G_{bu}(k, \mathbf{X}; \tau, \tau_1, T) H_{bu}(k, \mathbf{X}; \tau, \tau_1, T)$$

Helical functions satisfy  $H_{bu}(\tau, \tau_1) = H_{ub}(\tau_1, \tau)$

$$\alpha_X = -\frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 G_{bu}(k, \mathbf{X}; \tau, \tau_1, T) H_{ub}(k, \mathbf{X}; \tau, \tau_1, T) \\ + \frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 G_{bu}(k, \mathbf{X}; \tau, \tau_1, T) H_{ub}(k, \mathbf{X}; \tau_1, \tau, T)$$

Symmetric and anti-symmetric parts of  $H_{ub}$   
with respect to the exchange of time variables

$$H_{ub}^{(S)}(\tau, \tau_1) = \frac{1}{2} (H_{ub}(\tau, \tau_1) + H_{ub}(\tau_1, \tau))$$

$$H_{ub}^{(A)}(\tau, \tau_1) = \frac{1}{2} (H_{ub}(\tau, \tau_1) - H_{ub}(\tau_1, \tau)) \quad \text{Non-equilibrium effect}$$

$$H_{ub}(\tau, \tau_1) \neq H_{ub}(\tau_1, \tau)$$

$$\alpha_{\mathbf{X}} = -\frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 [G_{ub}(k, \mathbf{X}; \tau, \tau_1, T) + G_{bu}(k, \mathbf{X}; \tau, \tau_1, T)] H_{ub}^{(A)}(k, \mathbf{X}; \tau, \tau_1, T) \\ + \frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 [G_{ub}(k, \mathbf{X}; \tau, \tau_1, T) - G_{bu}(k, \mathbf{X}; \tau, \tau_1, T)] H_{ub}^{(S)}(k, \mathbf{X}; \tau_1, \tau, T)$$

Non-equilibrium alpha

$$\alpha_{\text{neq}} = -\frac{1}{3} \int d^3k \int_{-\infty}^{\tau} d\tau_1 [G_{ub}(k, \mathbf{X}; \tau, \tau_1, T) + G_{bu}(k, \mathbf{X}; \tau, \tau_1, T)] H_{ub}^{(A)}(k, \mathbf{X}; \tau, \tau_1, T)$$

Simple model

$$\mathcal{G}(\tau, \tau_1) = G_{ub}(\tau, \tau_1) + G_{bu}(\tau, \tau_1) \quad \text{Independent of } k$$

$$\alpha_{\text{neq}} = -\frac{1}{3} \int_{-\infty}^{\tau} d\tau_1 \mathcal{G}(\tau, \tau_1) \langle \mathbf{u}'_{00} \cdot \mathbf{j}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau_1)$$

$$\text{where } \langle \mathbf{u}'_{00} \cdot \mathbf{j}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau_1) = \frac{1}{2} [\langle \mathbf{u}'_{00}(\mathbf{x}; \tau) \cdot \mathbf{j}'_{00}(\mathbf{x}; \tau_1) \rangle - \langle \mathbf{u}'_{00}(\mathbf{x}; \tau_1) \cdot \mathbf{j}'_{00}(\mathbf{x}; \tau) \rangle]$$

**Memory effect is crucial** since  $\langle \mathbf{u}'_{00} \cdot \mathbf{j}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau) = 0$

$$\begin{pmatrix} \frac{\partial}{\partial \tau} + \nu k^2 & 0 \\ 0 & \frac{\partial}{\partial \tau} + \eta k^2 \end{pmatrix} \begin{pmatrix} G_{uu}^{ij} & G_{bu}^{ij} \\ G_{ub}^{ij} & G_{bb}^{ij} \end{pmatrix} + i \begin{pmatrix} -2M^{ikm} \int_{\Delta} u_{00}^k & 2M^{ikm} \int_{\Delta} b_{00}^k \\ N^{ikm} \int_{\Delta} b_{00}^k & -N^{ikm} \int_{\Delta} u_{00}^k \end{pmatrix} \begin{pmatrix} G_{uu}^{mj} & G_{bu}^{mj} \\ G_{ub}^{mj} & G_{bb}^{mj} \end{pmatrix} = \delta^{ij} \delta(\tau - \tau') \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

The response functions  $\mathbf{G}_{ub}$  must be an odd function of  $\mathbf{b}_{00}$

$\langle \mathbf{u}'_{00} \cdot \mathbf{j}'_{00} \rangle$  pure scalar  $\longrightarrow$   $G_{ub}$  skew

Dynamic quantity that is odd in  $\mathbf{b}_{00}$  and skew  $\longrightarrow$  Cross helicity  $\langle \mathbf{u}'_{00} \cdot \mathbf{b}'_{00} \rangle$

$$\alpha_{\text{neq}} \sim -\frac{2}{3} \int_{-\infty}^{\tau} d\tau_1 \Upsilon^{(S)}(\mathbf{x}; \tau, \tau_1) \langle \mathbf{u}'_{00} \cdot \mathbf{j}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau_1)$$

where  $\Upsilon(\mathbf{x}; \tau, \tau_1) \equiv \frac{\langle \mathbf{u}'_{00}(\mathbf{x}; \tau) \cdot \mathbf{b}'_{00}(\mathbf{x}; \tau_1) \rangle}{\sqrt{\langle u'^2_{00} \rangle(\mathbf{x}; \tau)} \sqrt{\langle b'^2_{00} \rangle(\mathbf{x}; \tau_1)}}$  Normalised cross helicity

$$\Upsilon^{(S)}(\mathbf{x}; \tau, \tau_1) = \frac{1}{2} [\Upsilon(\mathbf{x}; \tau, \tau_1) + \Upsilon(\mathbf{x}; \tau_1, \tau)]$$

$\longrightarrow$

$$\alpha_{\text{neq}} \sim -\frac{2}{3} \int_{-\infty}^{\tau} d\tau_1 \left( \Upsilon^{(S)}(\mathbf{x}; \tau, \tau_1) \right)^2 \langle \mathbf{u}'_{00} \cdot \boldsymbol{\omega}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau_1)$$

## Non-equilibrium alpha

$$\alpha_{\text{neq}} \sim -\frac{2}{3} \int_{-\infty}^{\tau} d\tau_1 \left( \Upsilon^{(S)}(\mathbf{x}; \tau, \tau_1) \right)^2 \langle \mathbf{u}'_{00} \cdot \boldsymbol{\omega}'_{00} \rangle^{(A)}(\mathbf{x}; \tau, \tau_1)$$

## Validation of $\alpha_{\text{neq}}$ by DNSs

$$\mathbf{g} \parallel \nabla \rho \parallel \mathbf{B}_0 \parallel \boldsymbol{\omega}_F$$

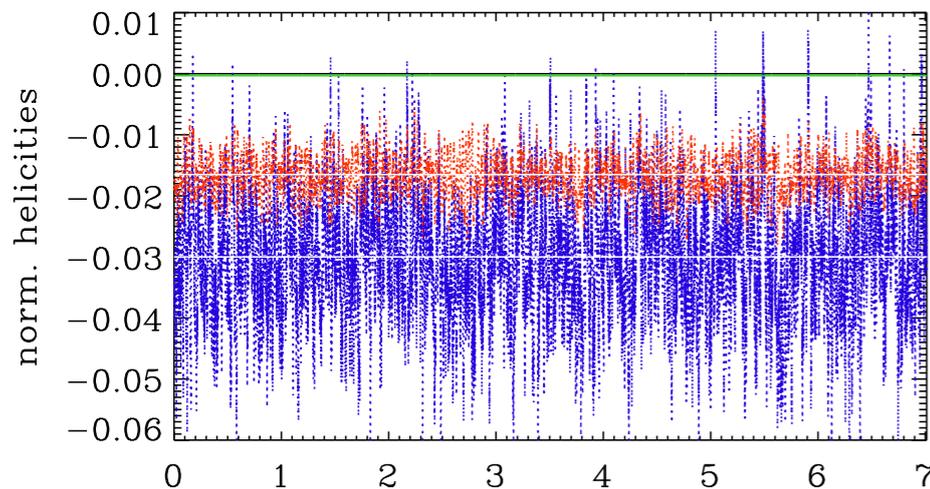
Cross helicity

Non-equilibrium helicity

$$\boldsymbol{\Omega}_F \parallel \nabla \rho \longrightarrow \text{Kinetic helicity}$$

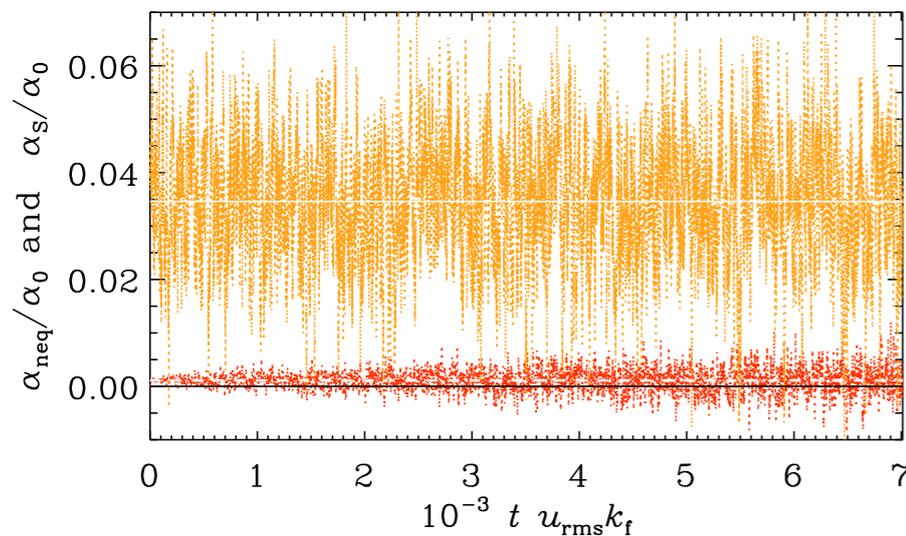
$$\mathbf{B} \parallel \nabla \rho \longrightarrow \text{Cross helicity}$$

## Co-existence of kinetic and cross helicities



	$\frac{g}{c_s^2 k_1}$	$\frac{v_{A0}}{c_s}$	$\frac{\langle \mathbf{u}' \cdot \mathbf{b}' \rangle}{\sqrt{\langle u'^2 \rangle \langle b'^2 \rangle}}$	$\frac{\langle \mathbf{u}' \cdot \mathbf{w}' \rangle}{\sqrt{\langle u'^2 \rangle \langle w'^2 \rangle}}$	$\frac{\langle \mathbf{b}' \cdot \mathbf{j}' \rangle}{\sqrt{\langle u'^2 \rangle \langle w'^2 \rangle}}$	$\frac{\alpha_{\text{neq}}}{\alpha_0}$	$\frac{\alpha_S}{\alpha_0}$	$\frac{u_{\text{rms}}}{c_s}$	$\frac{v_A^{\text{rms}}}{c_s}$
C	0.5	0.01	$-9.8 \times 10^{-3}$	$-1.6 \times 10^{-2}$	$-2.0 \times 10^{-4}$	$7.8 \times 10^{-4}$	$1.8 \times 10^{-2}$	0.10	0.03
A	1.0	0.01	$-1.7 \times 10^{-2}$	$-3.0 \times 10^{-2}$	$-3.3 \times 10^{-4}$	$1.1 \times 10^{-3}$	$3.5 \times 10^{-2}$	0.11	0.04
D	2.0	0.01	$-2.0 \times 10^{-2}$	$-3.6 \times 10^{-2}$	$-2.8 \times 10^{-4}$	$6.1 \times 10^{-4}$	$4.1 \times 10^{-2}$	0.16	0.04
E	0.5	0.10	$-5.5 \times 10^{-2}$	$-1.9 \times 10^{-2}$	$-6.2 \times 10^{-4}$	$-5.6 \times 10^{-3}$	$1.5 \times 10^{-2}$	0.08	0.07
B	1.0	0.10	$-5.3 \times 10^{-2}$	$-3.2 \times 10^{-2}$	$-1.2 \times 10^{-2}$	$2.3 \times 10^{-3}$	$1.8 \times 10^{-2}$	0.09	0.12

TABLE 1. Summary of the simulation results for Runs A–E.



	$k_f$	$\frac{\langle \mathbf{u}' \cdot \mathbf{b}' \rangle}{\sqrt{\langle u'^2 \rangle \langle b'^2 \rangle}}$	$\frac{\langle \mathbf{u}' \cdot \mathbf{w}' \rangle}{\sqrt{\langle u'^2 \rangle \langle w'^2 \rangle}}$	$\frac{\langle \mathbf{b}' \cdot \mathbf{j}' \rangle}{\sqrt{\langle u'^2 \rangle \langle w'^2 \rangle}}$	$\frac{\alpha_{\text{neq}}}{\alpha_0}$	$\frac{\alpha_S}{\alpha_0}$	$\frac{u_{\text{rms}}}{c_s}$	$\frac{v_A^{\text{rms}}}{c_s}$
A	30	$-1.7 \times 10^{-2}$	$-3.0 \times 10^{-2}$	$-3.3 \times 10^{-4}$	$1.1 \times 10^{-3}$	$3.5 \times 10^{-2}$	0.11	0.04
A2	10	$-1.3 \times 10^{-1}$	$-1.2 \times 10^{-1}$	$1.3 \times 10^{-3}$	$-1.7 \times 10^{-2}$	$6.9 \times 10^{-2}$	0.12	0.12
A3	3	$-6.4 \times 10^{-2}$	$-2.1 \times 10^{-1}$	$-3.0 \times 10^{-2}$	$-6.0 \times 10^{-3}$	$5.5 \times 10^{-2}$	0.19	0.09

TABLE 2. Summary of the simulation results for Runs A, A2, and A3.

# Summary of non-equilibrium effects on dynamos

Cross-interaction response functions

$$G_{ub}, \quad G_{bu}$$

Torsional cross correlations

$$\langle \mathbf{u}' \cdot \mathbf{j}' \rangle, \quad \langle \boldsymbol{\omega}' \cdot \mathbf{b}' \rangle$$

Non-equilibrium properties of turbulence

$$\langle \mathbf{u}'_{00}(\mathbf{x}; \tau) \cdot \mathbf{j}'_{00}(\mathbf{x}; \tau_1) \rangle \neq \langle \mathbf{u}'_{00}(\mathbf{x}; \tau_1) \cdot \mathbf{j}'_{00}(\mathbf{x}; \tau) \rangle$$

Co-existence of kinetic and cross helicities

$$\langle \mathbf{u}' \cdot \mathbf{b}' \rangle, \quad \langle \mathbf{u}' \cdot \boldsymbol{\omega}' \rangle$$

# **Beyond heuristic modelling**

# Mean-field equations in compressible MHD

Yokoi, N., J. Plasma Phys. **84**, 735840501 & 775840603 (2018a,b)

Density	$\frac{\partial \bar{\rho}}{\partial t} + \nabla \cdot (\bar{\rho} \mathbf{U}) = -\nabla \cdot \langle \rho' \mathbf{u}' \rangle$	Turb. mass flux
Momentum	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} U^\alpha + \frac{\partial}{\partial x^a} \bar{\rho} U^a U^\alpha \\ = -(\gamma_0 - 1) \frac{\partial}{\partial x^\alpha} \bar{\rho} Q + \frac{\partial}{\partial x^\alpha} \mu S^{a\alpha} + (\mathbf{J} \times \mathbf{B})^\alpha \\ - \frac{\partial}{\partial x^\alpha} \left( \underbrace{\bar{\rho} \langle u'^a u'^\alpha \rangle}_{\text{Reynolds stress}} - \frac{1}{\mu_0} \underbrace{\langle b'^a b'^\alpha \rangle}_{\text{Turb. Maxwell stress}} + U^a \underbrace{\langle \rho' u'^\alpha \rangle}_{\text{Turb. energy flux}} + U^\alpha \underbrace{\langle \rho' u'^a \rangle}_{\text{Turb. mass-energy correl.}} \right) + R_U^\alpha \end{aligned}$	
Internal energy	$\begin{aligned} \frac{\partial}{\partial t} \bar{\rho} Q + \nabla \cdot (\bar{\rho} \mathbf{U} Q) = \nabla \cdot \left( \frac{\kappa}{C_V} \nabla Q \right) - \nabla \cdot (\bar{\rho} \langle q' \mathbf{u}' \rangle + Q \langle \rho' \mathbf{u}' \rangle + \mathbf{U} \langle \rho' q' \rangle) \\ - (\gamma_0 - 1) \left( \bar{\rho} Q \nabla \cdot \mathbf{U} + \underbrace{\bar{\rho} \langle q' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. energy dilatation}} + Q \underbrace{\langle \rho' \nabla \cdot \mathbf{u}' \rangle}_{\text{Turb. mass dilatation}} \right) + R_Q \end{aligned}$	
Magnetic field	$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B} + \langle \mathbf{u}' \times \mathbf{b}' \rangle) + \eta \nabla^2 \mathbf{B}$	Turb. electromotive force

# Some main results of theoretical analysis

## Reynolds and turbulent Maxwell stress

$$\langle \mathbf{u}'\mathbf{u}' - \mathbf{b}'\mathbf{b}' \rangle_D = -\overset{\text{eddy viscosity}}{\nu_K} \mathbf{S} + \overset{\text{cross helicity}}{\nu_M} \mathcal{M} + \overset{\text{inhomogeneous helicity}}{\eta_H} \nabla H \Omega_* + \dots$$

D: deviatoric part

cross helicity

mean velocity strain  $\mathbf{S} = \nabla \mathbf{U} + (\nabla \mathbf{U})^\dagger$  mean magnetic-field strain  $\mathcal{M} = \nabla \mathbf{B} + (\nabla \mathbf{B})^\dagger$

$\Omega_*$ : absolute mean vorticity (mean vorticity + rotation)

## Turbulent electromotive force

$$\begin{aligned} \langle \mathbf{u}' \times \mathbf{b}' \rangle = & \overset{\text{Turb. Mag. Diffusivity}}{-(\beta + \zeta)} \nabla \times \mathbf{B} + \overset{\text{Alpha effect}}{\gamma} \nabla \times \mathbf{U} + \alpha \mathbf{B} + (\nabla \zeta) \times \mathbf{B} \\ & - \chi_\rho \nabla \bar{\rho} \times \mathbf{B} - \chi_Q \nabla Q \times \mathbf{B} - \chi_D \frac{D\mathbf{U}}{Dt} \times \mathbf{B} \quad \text{Compressibility} \end{aligned}$$

Cross-helicity effect      Magnetic pumping

## Turbulent mass flux

$$\langle \rho' \mathbf{u}' \rangle = -\kappa_\rho \nabla \bar{\rho} - \kappa_Q \nabla Q - \kappa_D \frac{D\mathbf{U}}{DT} - \kappa_B \mathbf{B}$$

## Turbulent internal-energy flux

$$\langle q' \mathbf{u}' \rangle = -\eta_Q \nabla Q - \eta_\rho \nabla \bar{\rho} - \eta_B \mathbf{B} \quad + \text{Non-equilibrium effects}$$

# Transport coefficients

Mean fields

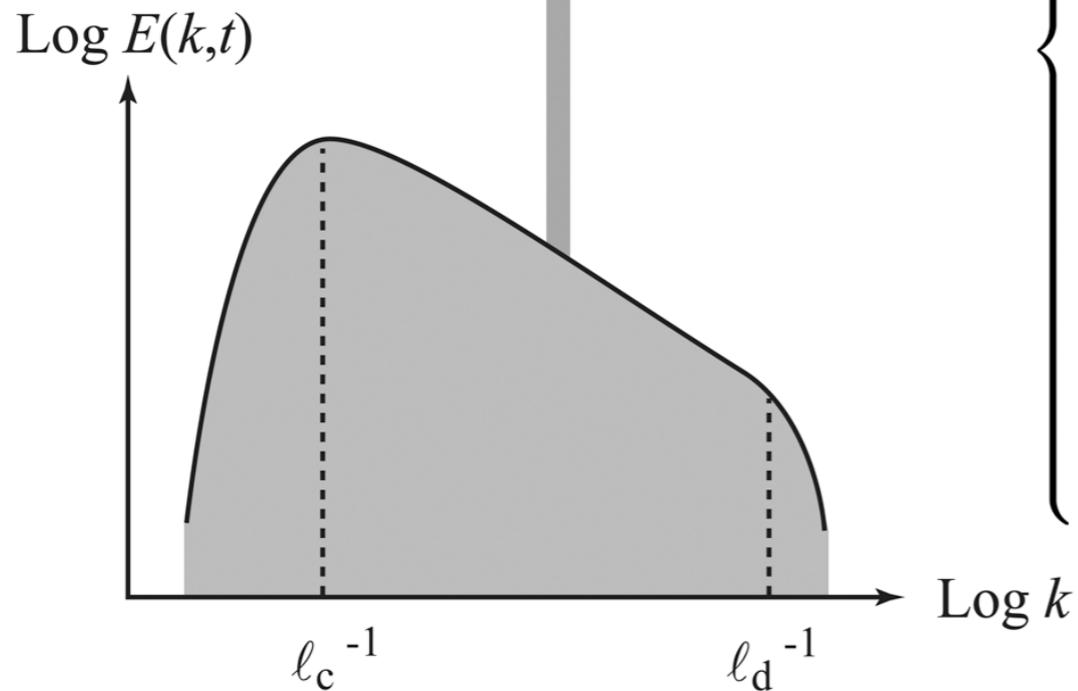
$\bar{\rho}$

**U**

**B**

**Q**

represented by



One-point  
turbulent statistical quantities

$$[K_\rho \equiv \langle \rho'^2 \rangle]$$

$$K \equiv \frac{1}{2} \langle \mathbf{u}'^2 + \mathbf{b}'^2 \rangle$$

$$\varepsilon \equiv \nu \left\langle \left( \frac{\partial u'_a}{\partial x_b} \right)^2 \right\rangle + \eta \left\langle \left( \frac{\partial b'_a}{\partial x_b} \right)^2 \right\rangle$$

$$W \equiv \langle \mathbf{u}' \cdot \mathbf{b}' \rangle$$

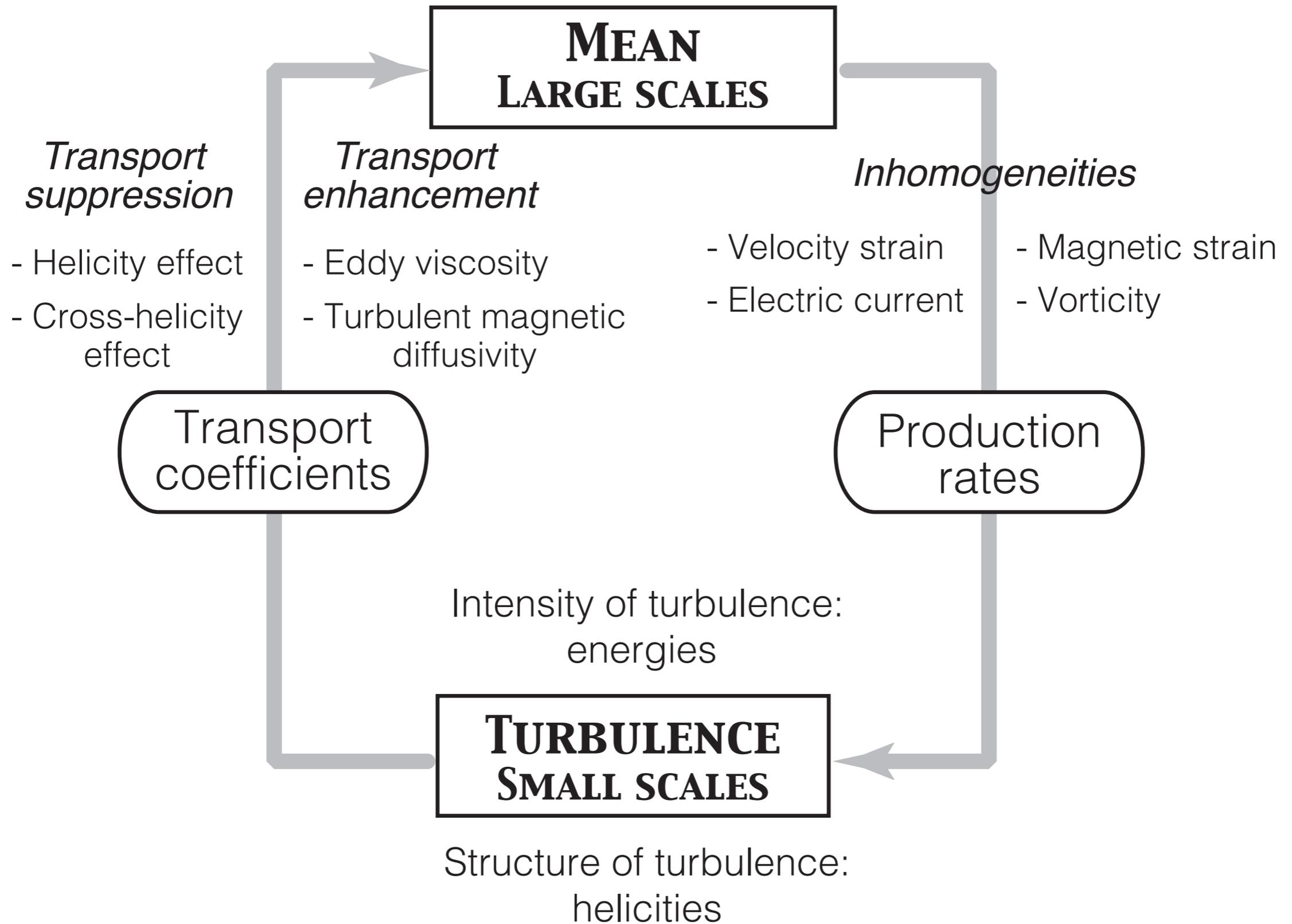
$$[K_R \equiv \frac{1}{2} \langle \mathbf{u}'^2 - \mathbf{b}'^2 \rangle]$$

$$[\varepsilon_W \equiv (\nu + \eta) \left\langle \frac{\partial u'_a}{\partial x_b} \frac{\partial b'_a}{\partial x_b} \right\rangle]$$

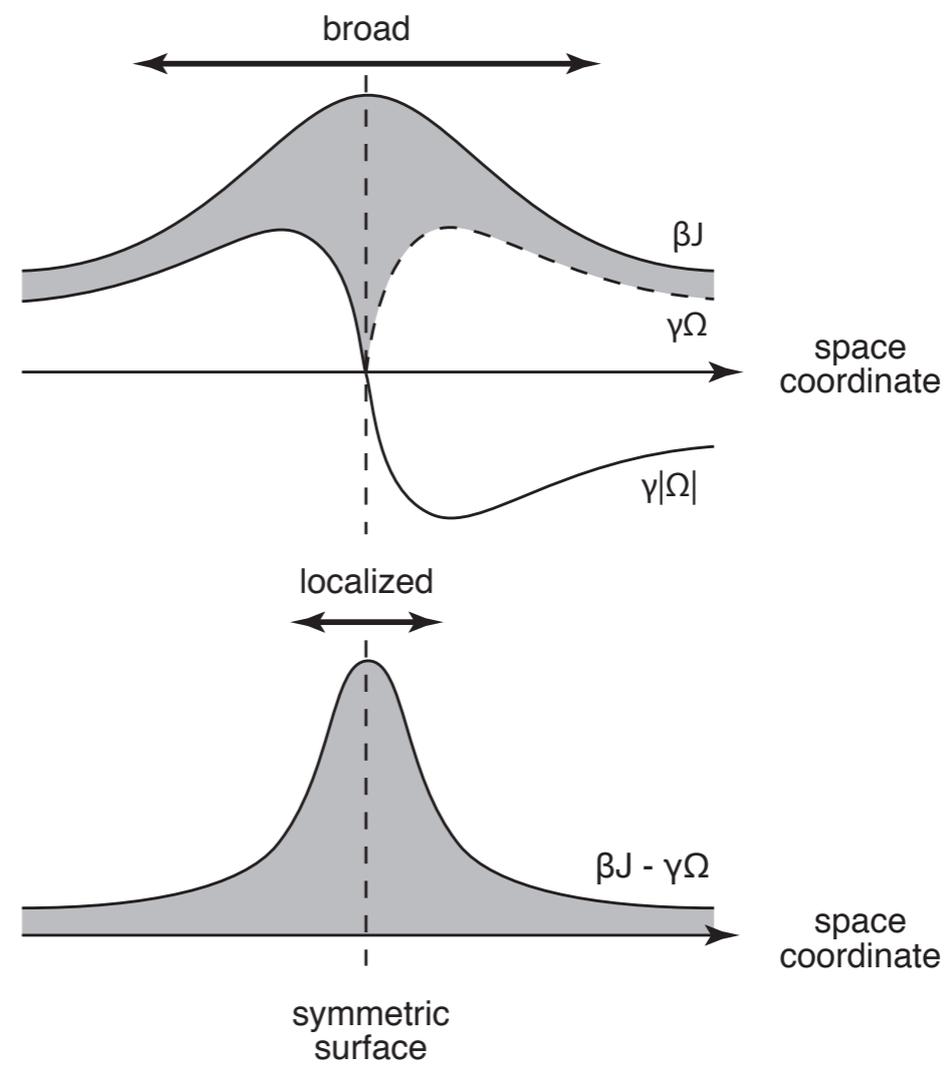
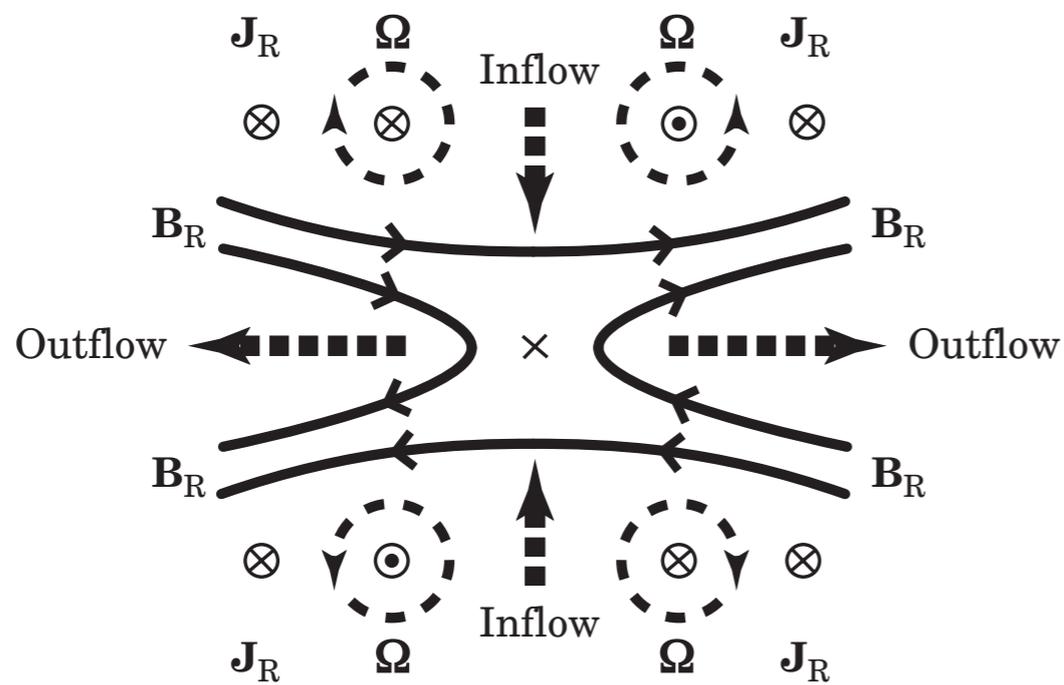
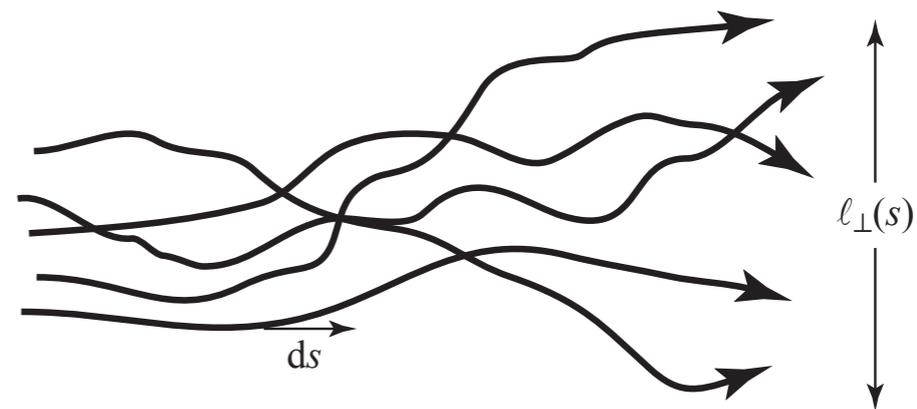
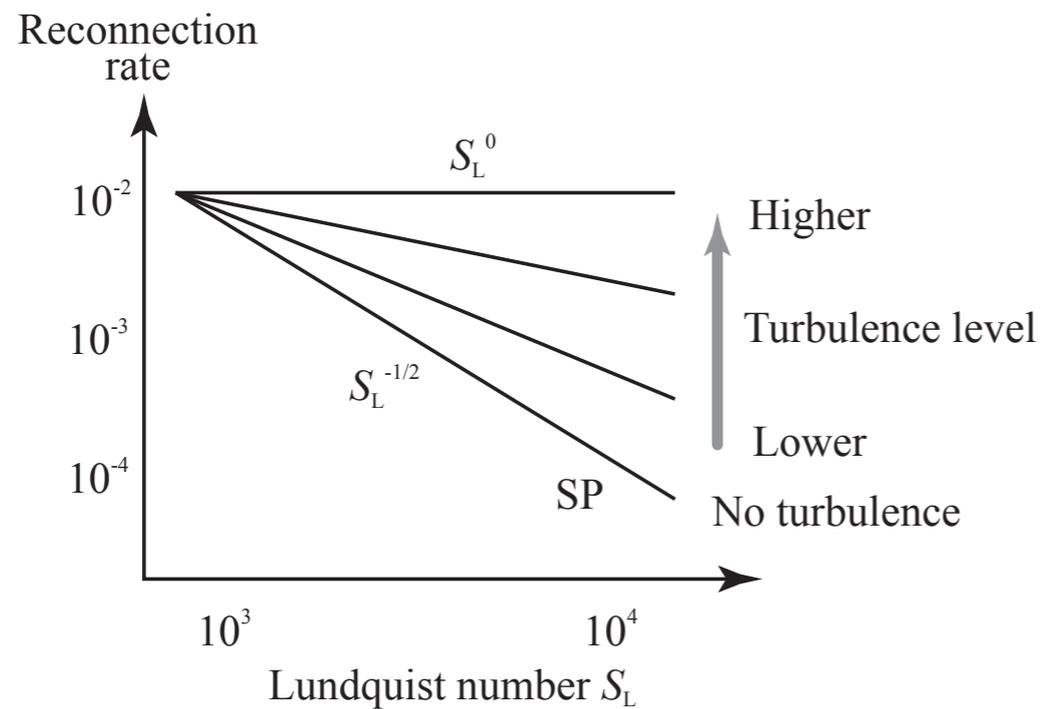
$$[H \equiv \langle -\mathbf{u}' \cdot \boldsymbol{\omega}' + \mathbf{b}' \cdot \mathbf{j}' \rangle]$$

Beyond the heuristic modelling

- Model structures
- Transport coefficients



# **Turbulent magnetic reconnection**



# Turbulence modeling approach to magnetic reconnection

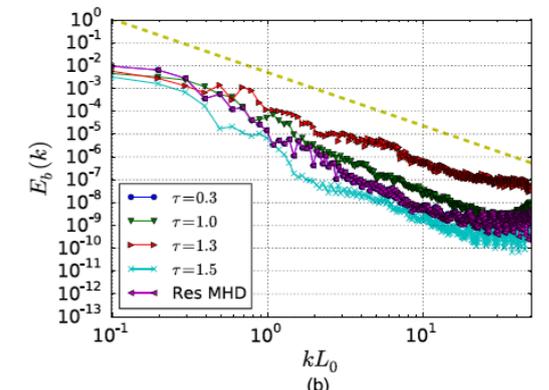
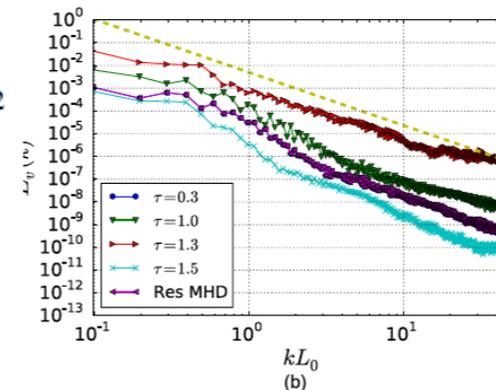
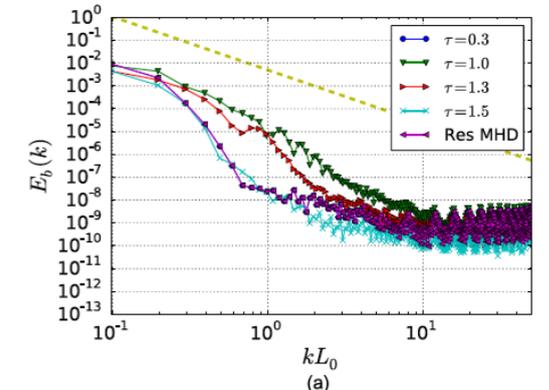
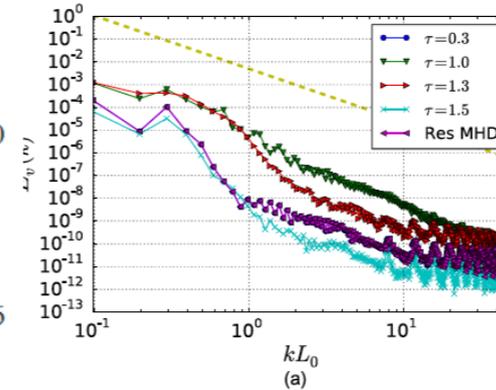
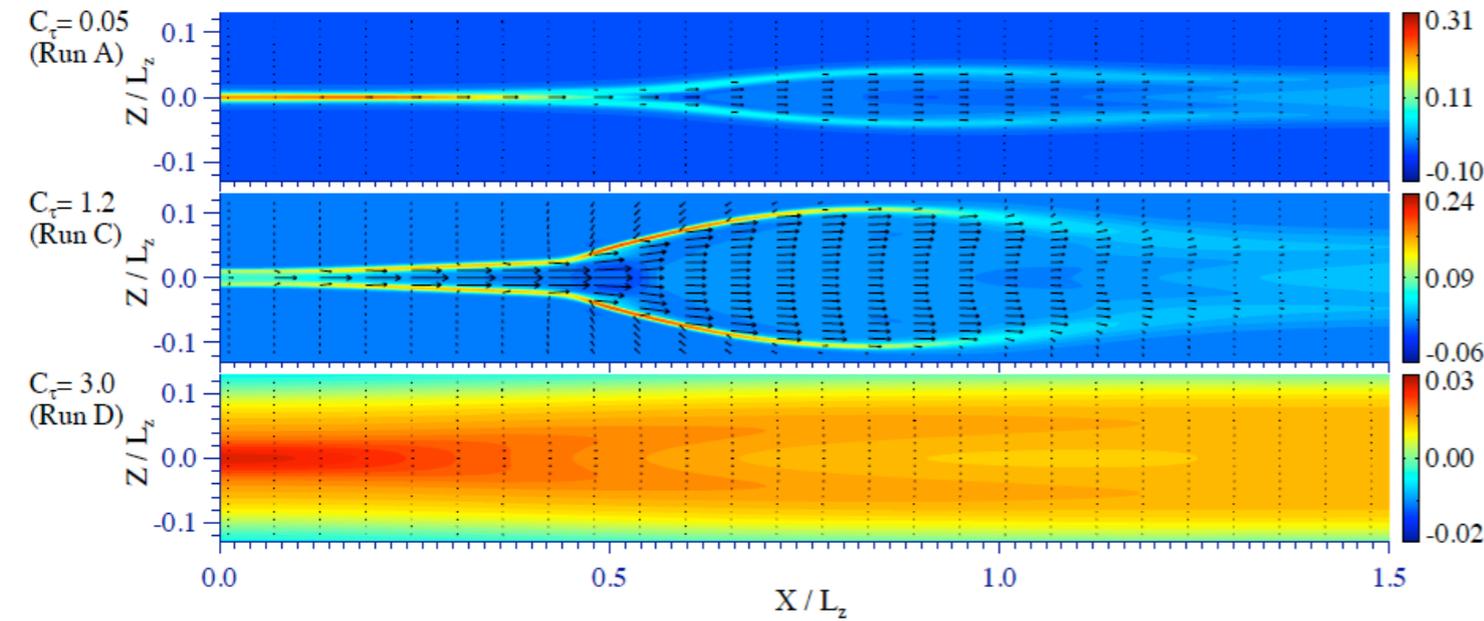
Yokoi & Hoshino, Phys. Plasmas **18**, 111208 (2011)

## Spectral evolution

### Electric-current and flow structures

### Kinetic energy

### Magnetic energy



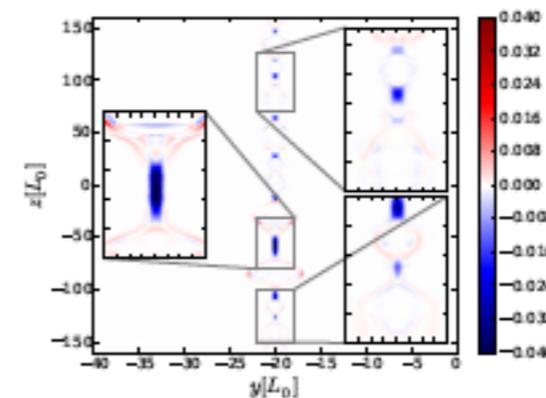
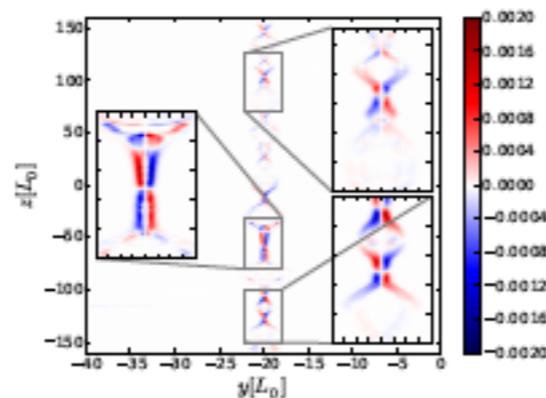
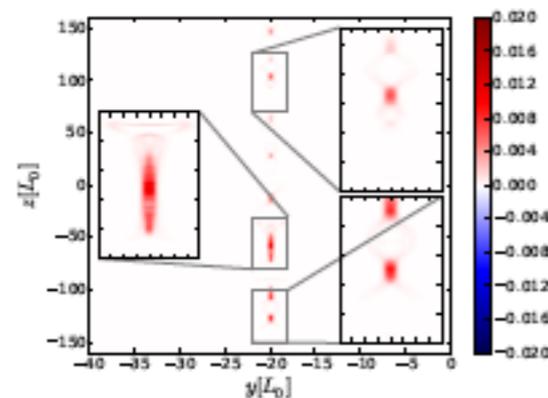
Higashimori, Yokoi & Hoshino, PRL **110**, 255001 (2013)

### Energy

### Cross helicity

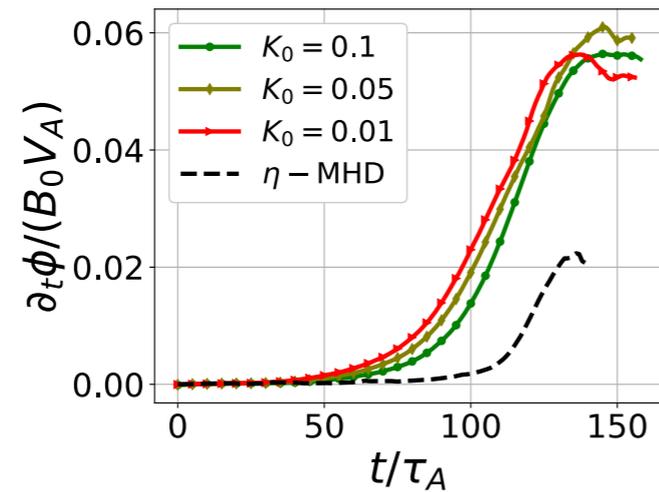
### Helicity

### Guide field

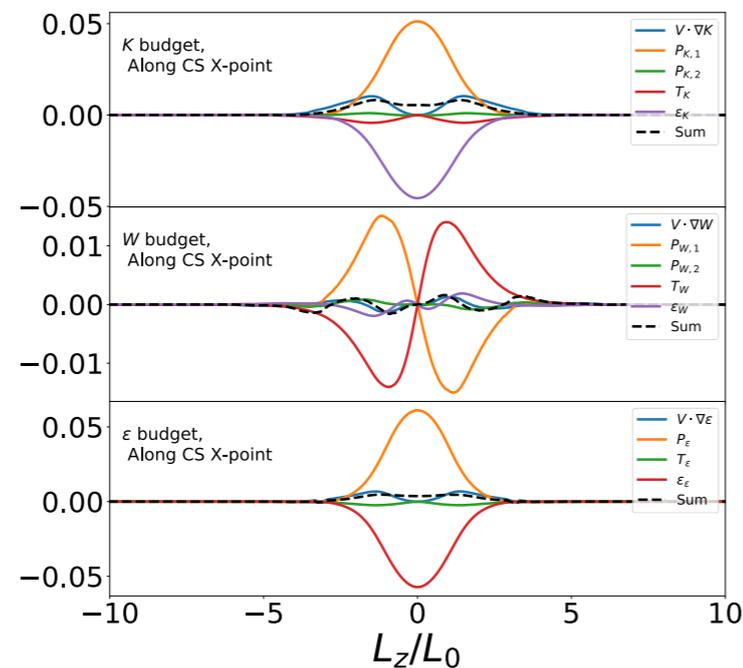


Widmer, Büchner & Yokoi, Phys. Plasmas **23**, 092304 (2016)

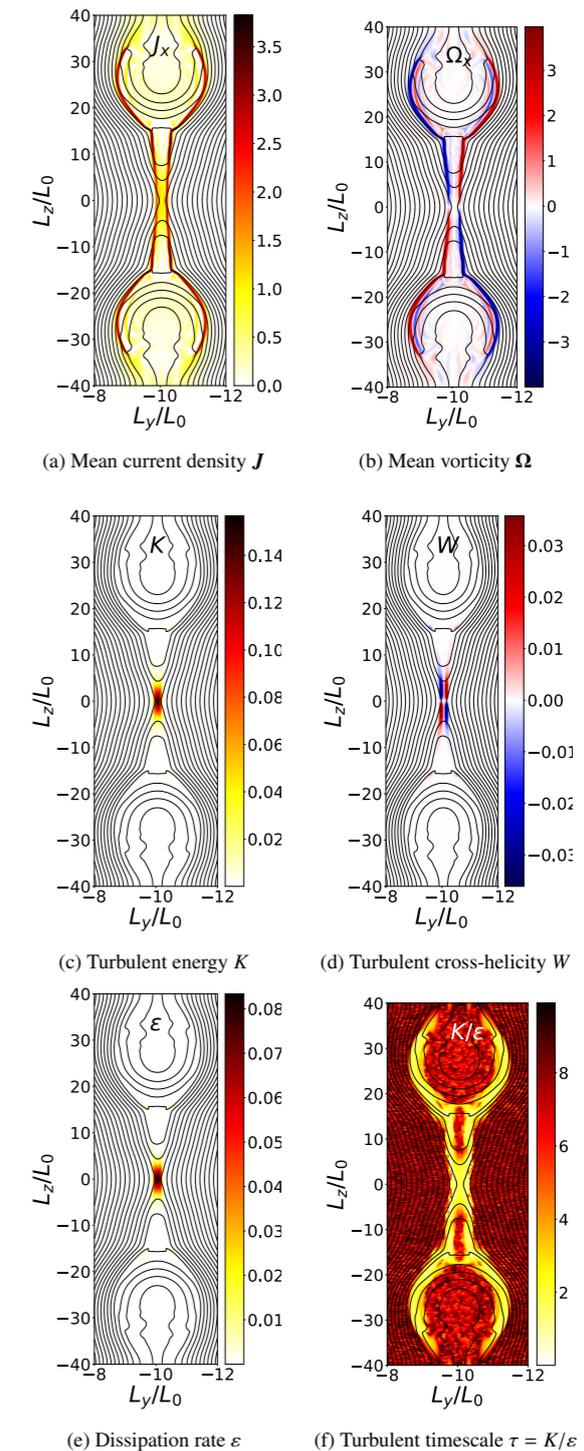
Widmer, Büchner, and Yokoi (2019) Phys. Plasmas, accepted on 25 Sep. 2019



Reconnection rate with changing initial turbulent energy



Budget of turbulent energy  $K$ , cross helicity  $W$ , and energy dissipation rate  $\epsilon$

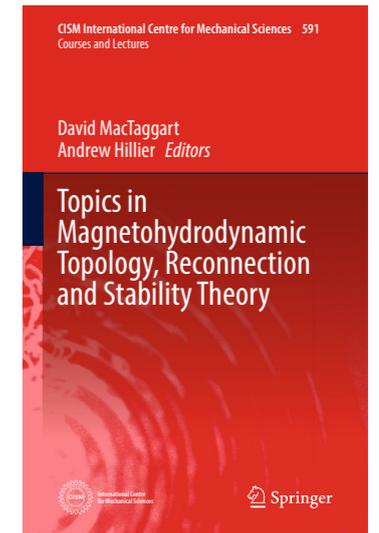


Spatial distributions of mean and turbulent quantities

# Advertisement: Turbulence, Reconnection, Dynamos, Helicities...

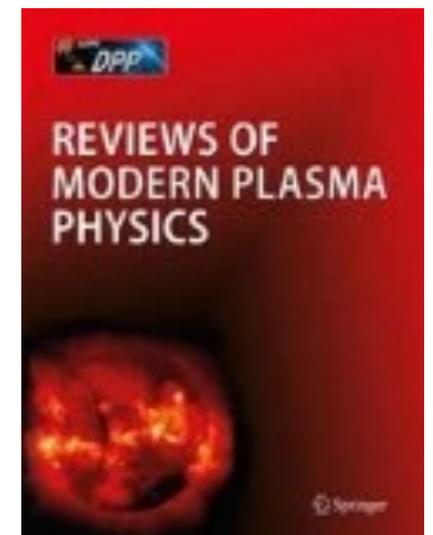
**Yokoi, N.** “Turbulence, transport and reconnection,” Chap. 6 in **Topics in Magnetohydrodynamic Topology, Reconnection and Stability Theory: CISM International Centre for Mechanical Sciences 591 pp.177-265 (Springer, 2020)**

[https://doi.org/10.1007/978-3-030-16343-3\\_6](https://doi.org/10.1007/978-3-030-16343-3_6)



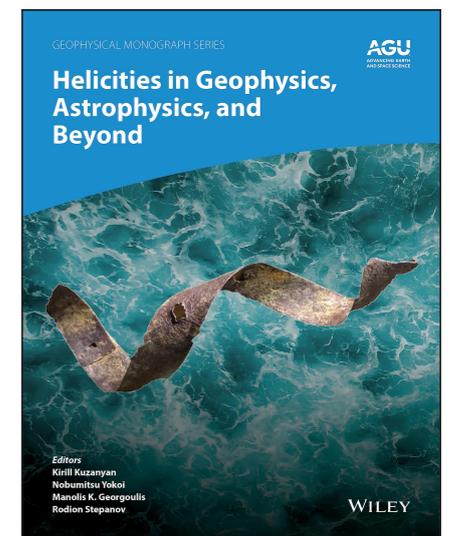
**Pouquet, A. & Yokoi, N.** “Helical fluid and (Hall)-MHD turbulence: a brief review,” **Phil. Trans. Roy. Soc. A 380, 20210081 (2022)**

<https://doi.org/10.1098/rsta.2021.0087>



**Yokoi, N.** “Unappreciated cross-helicity effects in plasma physics: anti-diffusion effects in dynamo and momentum transport,” **Reviews of Modern Plasma Physics 7, 33 (2023)**

<https://doi.org/10.1007/s41614-023-00133-4>



**Kuzanyan, K., Yokoi, N., Georgoulis, M. & Stepanov, R.** (Eds.) **AGU Books 283: Helicities in Geophysics, Astrophysics, and Beyond (Wiley, 2024)**

<https://doi.org/10.1002/9781119841715>

Nobu Yokoi <nobyokoi@iis.u-tokyo.ac.jp>

## Theory

Yokoi, N. "Turbulence, transport and reconnection," Chap. 6 in *Topics in Magnetohydrodynamic Topology, Reconnection and Stability Theory*: CISM International Centre for Mechanical Sciences 591 pp.177-265 (Springer, 2020)  
[https://doi.org/10.1007/978-3-030-16343-3\\_6](https://doi.org/10.1007/978-3-030-16343-3_6)

## Helicity

Pouquet, A. & Yokoi, N. "Helical fluid and (Hall)-MHD turbulence: a brief review," *Phil. Trans. Roy. Soc. A* 380, 20210081-1-18 (2022)  
<https://doi.org/10.1098/rsta.2021.0087>

Yokoi, N. "Transports in helical fluid turbulence," pp.25-50, in Kuzanyan, Yokoi, Georgoulis & Stepanov (eds.) *AGU Book: Helicities in Geophysics, Astrophysics and Beyond* (Wiley, 2023)  
<https://doi.org/10.1002/9781119841715>

Yokoi, N. & Brandenburg, A. "Global flow generation by inhomogeneous helicity," (2016) *Phys. Rev. E*. 93, 033125-1-14 (2016)  
<https://doi.org/10.1103/PhysRevE.93.033125>

## Dynamos

Yokoi, N. "Cross-helicity and related dynamo," *Geophys. Astrophys. Fluid Dyn.* 107, 114-184 (2013)  
<https://doi.org/10.1080/03091929.2012.754022>

Yokoi, N. "Unappreciated cross-helicity effects in plasma physics: Anti-diffusion effects in dynamo and momentum transport," *Rev. Mod. Plasma Phys.* 7, 33-1-98 (2023)  
<https://doi.org/10.1007/s41614-023-00133-4>

## Convection

Yokoi, N., Masada, Y. & Takiwaki, T. “Modelling stellar convective transport with plumes: I. Non-equilibrium turbulence effect in double-averaging formulation,” MNRAS 516, 2718–2735 (2022)

<https://doi.org/10.1093/mnras/stac1181>

Yokoi, N. “Non-Equilibrium Turbulent Transport in Convective Plumes Obtained from Closure Theory,” Atmosphere 14, 1013-1-22 (2023)

<https://doi.org/10.3390/atmos14061013>

## Strong compressibility

### Electromotive force

Yokoi, N. “Electromotive force in strongly compressible magneto-hydrodynamic turbulence,” J. Plasma Phys. 84, 735840501-1-26 (2018)

<https://doi.org/10.1017/S0022377818000727>

### Mass/Heat

Yokoi, N. “Mass and internal-energy transports in strongly compressible magnetohydrodynamic turbulence,” J. Plasma Phys. 84, 775840603-1-30 (2018)

<https://doi.org/10.1017/S0022377818001228>

## Instability

Yokoi, N. & Tobias, S. M. “Magnetoclinicity Instability,” Prog. in Turb. IX, Springer Proceedings in Physics 267, 273-279 (2021)

<https://doi.org/10.1080/03091929.2012.754022> (arXiv:2205.14453)